

THE
DAVID DUNLAP
OBSERVATORY



ETTERS

NEWSPAPER CLIPPINGS

PRINTED ARTICLES

AND PHOTOGRAPHS

Relating to the

DAVID DUNLAP

OBSERVATORY



Presented to
DAVID MOFFAT DUNLAP
by his
MOTHER
CHRISTMAS 1934



THE DAVID DUNLAP OBSERVATORY

In January 1927 Mrs. Dunlap informed Professor C.A. Chant that she was willing to provide an observatory for the University of Toronto, to be erected in the neighborhood of Toronto and to be a memorial to her husband, who died October 29, 1924. She stated however, that the actual work of construction would have to be delayed for three years. This time was spent in selecting a suitable site and in securing information about equipment. The order for the great telescope was given to the firm of Sir Howard Grubb, Parsons and Co., of Newcastle-on-Tyne, England, in May 1930, although the actual contract was not signed until the autumn of that year.

In the summer of 1931 there was some discussion regarding a public announcement of the project.



201 Madison Ave.,
June 1, 1930

Dear Mrs. Dunlap, —

I had a final talk with Mr. Holden last Wednesday and gave him letters of introduction to Sir Charles Parsons and Sir Frank Dyson.

President Palmer had expressed the wish to inform the Board of Governors of the present status of your great project. Dr. Cady and Bruce Macdonald knew in a general way what was proposed, but the others had no knowledge of the matter. The President was afraid that the news of the great telescope might leak out from England to the Board and they would wonder



July 10, 1930

Dear Mrs. Dunlap, —

We do not know just when Mr. Holden will approve the public announcement of our project, though I think he realizes that it cannot be delayed much longer. As you expect to be away in September, it might be well to prepare a statement to be given out. Indeed you spoke about something of this sort.

You could do this yourself, but if you wish I shall work it out in accordance with any suggestions you may give me. Suppose I go out to see you some evening next week and we can discuss the



Mrs. D. A. Dunlap,
93 Highlands Avenue,
Toronto 5.



Mrs. D. A. Dunlap,
Don. Alta Farm,
Todmorden P.O.

Ont.

Letters Regarding Announcement of the Observatory Project



*Mrs. Dunlap and Professor Chaut
At Don-Alida Farm, June 1929*

PROVISIONAL

A GREAT MEMORIAL

The important announcement has just been made that in the near future there will be established in the vicinity of Toronto an astronomical observatory which will rank with the world's greatest institutions of this sort. It will be erected by Mrs. D.A. Dunlap ^{for her husband} as a memorial to her husband David A. Dunlap, who died on October 29, 1924, and will be known as the David Dunlap Observatory.

Astronomy and geology were both favorite studies of Mr. Dunlap, but the former had a peculiar attraction for him. He was an intelligent student of the heavens and was always pleased to share his interest with others. This great project ^{which} has been under consideration for the last five years ~~and~~ will now be brought to completion. In working out the plans Mrs. Dunlap has had the assistance of Professor C.A. Chant.

The outstanding feature of the observatory will be a great reflecting telescope 74 inches in diameter. There is only one larger than this in the world, namely that at Mount Wilson in California. The instrument was ordered some time ago from the firm of Sir Howard Grubb, Parsons and Co., Newcastle-on-Tyne, England. It will be housed in a circular metal building, such construction being best for this purpose. The observatory building will be a beautiful structure in the classic style. The site chosen is on a prominent elevation just east of Yonge St. and one mile south of Richmond Hill. It is in the middle of a ~~one~~-acre plot of land which will be made into a park to be known as the David Dunlap Park.

When the observatory is completed it will be ^{operated by} ~~under~~ the Department of Astronomy of the University of Toronto, while the Park will be developed in a scientific way by the Faculty of Forestry.

It is believed that the new institution will bring distinction to the city, the province and indeed the whole dominion. It will be a noble memorial to a ~~worthy~~ citizen.

noble Canadian

lasting



First Draft of Announcement of Project

New Observatory to Rival World's Best

Wonder Telescope Will Be Situated On Site Near City

DONATES FINE GIFT



MRS. DAVID A. DUNLAP who with her son, D. Moffat Dunlap, will fund an astronomical observatory near Toronto as a memorial to her husband, the late David A. Dunlap. The second largest reflecting telescope in the world will be installed.

Munificent Gift of Mrs. David A. Dunlap and Her Son Will Provide Toronto Area With Astronomical Study Equipment Second Only to That of California

WILL BE BENEFIT TO UNIVERSITY

Department of Astronomy Will Have Facilities Available for Students, and Surrounding David Dunlap Park Is to Be Under Care of Faculty of Forestry

An astronomical observatory which, it is claimed, will rank with the finest in the world will be established in the near future in the vicinity of Toronto, according to announcement made yesterday by Professor Clarence A. Chant, head of the Department of Astronomy of the University of Toronto. The observatory will be known as the David Dunlap Observatory, and will be erected by Mrs. Dunlap and her son, D. Moffat Dunlap, as a memorial to the late Mr. Dunlap, who died on Oct. 29, 1924.

The outstanding feature of the observatory will be a large reflecting telescope seventy-four inches in diameter. It is stated that there is only one greater aperture in the world, namely, that on Mount Wilson, in California. The instrument is already under construction by the firm of Sir Howard Grubb, Parsons & Company, Newcastle-on-Tyne, England. It will be housed in a circular metal building, such construction being best for this purpose.

In Classic Style. The observatory building will be a beautiful structure in the classic style. It will be erected on a suitable site near Toronto—probably near Thornhill—in the midst of a large acreage which will be converted into a park to be known as the David Dunlap Park. When the observatory is completed it will be under the Department of Astronomy of the University of Toronto, while the park will be under the Faculty of Forestry.

Astronomy and geology were both favorite studies with the late Mr. Dunlap. He was a keen student of the stars and enjoyed sharing his knowledge with others. This project has been under consideration for the past five years, and is now being brought to completion. In working out the plans Mrs. Dunlap has been assisted by Professor Chant. No further information was given out officially beyond that which was contained in the statement. Professor Chant said the exact site of the building had not yet been determined, nor was the

price known. Mrs. Dunlap courteously declined to add anything whatever to the official statement.

The Department of Forestry had not heard anything of the gift, and did not know it was to care for the grounds until this article, who might have known something of the project were out of the city last night.

Though meteorological buildings come under the Department of Marine and Ottawa, apparently observatories do not, as Ottawa's latter nothing of the Dunlap gift.

Handsome Gift. The meaning of this exceedingly handsome gift by Mrs. Dunlap and her son to the University of Toronto was explained to The Globe last night by A. R. Haxford, K.C., Vice-Chairman of the Royal Astronomical Society of Toronto, who has long been interested in the study of astronomy. "Anything that Mrs. Dunlap will do can be depended upon to be the very best," said Mr. Haxford, in expressing his pleasure at the news.

"Both Mr. and Mrs. Dunlap came to our meetings," explained the Vice-Chairman, who recalled that Mr. Dunlap was a lawyer and mining man. Since the death of her husband Mrs. Dunlap has been exceedingly interested.

Two years ago last fall, when leaders of the Royal Astronomical Society of England visited Toronto, Mrs. Dunlap entertained the delegates. Vice-Chairman Haxford said such an observatory to be adequately located would be built on high ground, probably up Young Street, about Thornhill. The site of the De La Salle Institute, at Ancaster is the highest point in York County, and Mr. Haxford thought this distance away from Toronto might be inexpedient, as it would be advisable to have it reasonably near to the University of Toronto. It would likely be built away from Young Street, in order to get away from railway and electrical vibrations and also from mists.

Mount Wilson Telescope. Speaking of the telescope at Mount Wilson, Mr. Haxford said it has a diameter of 100 inches. The telescope at the Dominion Observatory at Victoria, B.C., has a diameter of 72 inches, the glass alone weighing four tons. The

glass has a thickness of 14 inches, the whole being so delicately poised that it can be moved very easily. Dr. J. A. Brashear of Pittsburgh, Pa., supplied the lenses and mirrors. British telescopes have also been renowned for generations.

Mr. Haxford declared that Ontario astronomers have been trying and hoping for twenty years for this great gift, made possible by Mrs. Dunlap and her son. The gift is particularly welcome at this time, because the Toronto Observatory is being dismantled. This observatory was erected in 1882 to view the transit of the planet Venus.

At the University of Toronto, Mr. Haxford continued, there is a special department for the teaching of physical and astronomical sciences and to teach the subject adequately and at first hand, it is absolutely indispensable to have something like the gift of Mrs. Dunlap and her son. For the University of Toronto department information is now obtained from observatories outside of Canada or the Dominion observatories at Victoria or at Ottawa. The telescope at the Toronto Observatory is only six inches in diameter and not fit to cope with the greater problems of the local scientists.

The telescope, such as is described in the official statement, according to Vice-Chairman Haxford will permit astronomers to view the moons of Mars, five of the satellites of Jupiter, eight at least of the satellites of Saturn, and probably astronomers might get a glimpse of the new planet, Pluto discovered at Mount Wilson last year, and which is the outermost planet of the solar system.

Mr. Haxford said that Professor Chant of the University of Toronto had published a book a year ago on astronomy. This book which has had a wide-wide circulation, and has been translated into at least two European languages, was dedicated to Mrs. Dunlap. In this book Mr. Haxford is pictured as looking through his home-made set, which has a diameter of nine inches. "Mrs. Dunlap is wonderfully interested in astronomy, and this explains her handsome gift," concluded Mr. Haxford.

THE GLOBE, TORONTO, THURSDAY, JANUARY 1, 1931.

MAIL & EMPIRE, DECEMBER 30, 1930

A MUNIFICENT GIFT

Establishment in the neighborhood of Toronto of an astronomical observatory in connection with the university is the most important scientific accomplishment of recent years. It will be of incalculable value to those devoted to a study of the objects of the sciences, and will make of the city a centre for eminent men whose chief interest is in the solution of the actual problems. This observatory is to be built and equipped by Mrs. David A. Dunlap and her son Mr. D. Moffat Dunlap, as a memorial to the late Mr. Dunlap, who was an ardent student of astronomy.

It will be equipped to magnificently as to rank with the greatest similar observatories anywhere in the world. The building that is to house the delicate instruments of the astronomer is to be in keeping with the nature of its duties and will stand in the midst of an extensive acreage to be devoted to park purposes. The observatory itself will be under the Department of Astronomy of the University of Toronto, while the surrounding area will be developed in a scientific way by the Faculty of Forestry.

The science of astronomy I said to have had its origin with the Chaldeans who according to Greek historians were able to predict eclipses with considerable accuracy. All down the centuries great names have been associated with this most interesting of the sciences. Little by little the movement of the planets has been made known by observers and with increasing knowledge the legends and end of an eclipse may be announced to the minute. The great telescope to be installed in the Toronto observatory will be announced, perhaps not before the year 1931.

The new observatory will be a wonderful thing for Toronto, where there are many amateur astronomers, but still more valuable to the University, which then may take its place as a centre for the study of the mysteries of its stars vault above. In their munificent gift Mr. D. Dunlap and his son have secured, and will have, the gratitude not only of the people of this city but of the whole Dominion.

Born to Astronomers

To the Editor of The Globe. What a delightful piece of news was reported to your issue of Dec. 30 with reference to a new telescope which will be presented to Toronto in the near future. It is not only an instrument which will be an honor to the city, and a fitting memorial to one who, without doubt, must have many times when surveyed the heavens suffered from lack of adequate means to delve into the hidden mysteries of the universe, but an instrument that will be truly an incentive for our rising generation to take up this sublime science of last night.

I, like one of us who are not of the technical side give our whole support as much as we possibly can to bring about the best application of this splendid apparatus which will not only afford the professional student means of studying the higher branch of science, but will give also ways and means will be to aid to the uninitiated human to see and usually through this instrument and thereby bring a greater number of it down studying this particular branch of science if only in an amateur way.

Best Wishes. H. A. Robinson.

Observatory Planned for City As Memorial to D. A. Dunlap

Widow to Erect Plant With Second Largest Telescope in World.

One of the finest astronomical observatories in the world, with a giant reflecting telescope 74 inches in diameter, the second largest telescopic aperture in the world, will be erected in the vicinity of Toronto in the near future, according to an announcement made yesterday by Prof. C. A. Chant, head of the department of astronomy at the University of Toronto.

The observatory will be erected by Mrs. D. A. Dunlap and her son, D. Moffat Dunlap, as a memorial to the late David A. Dunlap, who died Oct. 29, 1924, and will be known as the "David Dunlap Observatory." A tentative site has been selected a few miles north of the city. Prof. Chant stated that the observatory, which was given several months ago to the firm of Sir Howard Grubb, Parsons and Company, Newcastle-on-Tyne, England, will be three years in the making. Prof. Chant estimated it would be up and one half years before the observatory is in operation. Plans have not yet been drawn, he said, and construction will not likely begin for some time.

The observatory will comprise two buildings, one a circular metal structure for the telescope and the other in classic style for offices and laboratories. There is only one telescope of greater aperture in the world than the instrument planned, that on Mount Wilson in California.

The acreage surrounding the observatory will be converted into a park which will be known as the David Dunlap park. The university's department of astronomy will have charge of the observatory while the forestry faculty will develop the park.

Astronomy and geology were both favorite studies of the late David Dunlap, but the former had a peculiar attraction for him. The new project, Prof. Chant stated, has been under consideration for the past five years. In working out the plans Mrs. Dunlap has had the assistance of Professor Chant, one of Canada's best known astronomers.

Toronto Daily Star, Dec. 31, 1931.

Gift Observatory for Toronto Will House World's Second Largest Telescope

Star Dec. 31, 1930



MRS. DAVID DUNLAP DONATES MAGNIFICENT ASTRONOMICAL RESEARCH CENTRE AS MEMORIAL TO HER HUSBAND—SEE UNCERTAIN.

As a memorial to her husband, late David Dunlap, Mrs. David Dunlap and her son, D. Moffat Dunlap, have donated a magnificent observatory to the University of Toronto, which will house the world's second largest telescope. The telescope will be 74 inches in diameter, second only to that at Mount Wilson observatory, California, which is 100 inches. The exact site of the new ob-

servatory is, as yet, uncertain, but speculation centres about the Alcona tract and W. G. and H. A. Marsh farms at Stop 23 Yonge St., which were purchased in 1923 for a purpose

which will be erected in memory of the late David Dunlap, who died Oct. 29, 1924, and will be known as the David Dunlap Observatory. Astronomy and geology were both favourite studies of Mr. Dunlap, but the former had a peculiar attraction for him. He was a keen student of the heavens, and always liked to share his knowledge with others. This project has been under consideration for the last five years and will now be brought to completion. In working out the plans, Mrs. Dunlap has had the assistance of Prof. C. A. Chant, head of the Department of Astronomy of the University of Toronto.

Toronto Daily Star, Dec. 31, 1930.

DUNLAP OBSERVATORY GIFT TO UNIVERSITY

Second Largest Telescope in World To Be Erected Here as Memorial

An astronomical observatory which will rank as one of the finest in the world, and which will contain a telescope with an aperture 74 inches in diameter, the second largest in the world, will be erected in the vicinity of Toronto, in the near future. According to an announcement made yesterday by Prof. C. A. Chant, head of the department of astronomy of the University of Toronto.

Housed in a circular metal building, with offices, libraries and studies in an adjoining structure of beautiful classic architecture, the observatory will be the gift to the university of Mrs. Dunlap and her son, D. Moffat Dunlap, as a memorial to the late D. A. Dunlap, who died Oct. 29, 1924.

The observatory will be situated in the midst of a large acreage on north Yonge St., which will be converted into a park, to be known as David Dunlap Park.

The observatory will be under the control of the department of astronomy of the university, while the park will be developed by the faculty of forestry. This is considered fitting, as the late Mr. Dunlap was interested in astronomy and geology.

It is understood that the university has been hampered for some time by the lack of a suitable observatory. The equipment erected in 1882 to view the transit of Venus is being demolished. Students of astronomy at the university have been raising information derived from foreign sources, chiefly from the Dominion observatory at Ottawa. The telescope of the Toronto observatory is only 28 inches in diameter and quite inadequate to cope with the requirements of the students.

The contract for the new telescope instrument for the David Dunlap observatory was let some time ago to the firm of Sir Howard Grubb, Parsons and Co., Newcastle-on-Tyne, England.

NEW OBSERVATORY FOR USE WHOLLY IN RESEARCH WORK

Mrs. Dunlap's Magnificent Gift an Outstanding Attraction to the University

SITE IS UNCERTAIN.

The magnificent gift to the University of Toronto of an astronomical observatory with the second largest reflecting telescope in the world by Mrs. Dunlap and her son, D. Moffat Dunlap, as a memorial to her husband, David Dunlap, cannot fail to be an outstanding attraction to the university for this part of the world.

Prof. Clarence A. Chant indicated in an interview to-day. The Toronto telescope, 74 inches in diameter, will be two inches larger in diameter than the one at the Victoria observatory in British Columbia, and will have many improvements of design on this great instrument, especially in case of operation. In all essential respects, the Toronto observatory here will resemble the huge instrument in Victoria. One difference here will be lower walls.

In reply to questions, Dr. Chant explained that the new observatory, when it is established here, will have no connection with the Dominion government and the government meteorological department.

One will be used for purely astronomical research," Dr. Chant explained, "and will be for research in geology, and will be for research in geology, we might make some observations to help them out in connection with weather observations."

In Seeking Site

At the present moment, Dr. Chant was emphatic, the site of the building is entirely uncertain, except that there is a farm close to Yonge St., about ten miles north, that the University can fall back on. But, at the present moment, Dr. Chant is making enquiries that have been brought to his attention. There is a farm close to Yonge St., about ten miles north, that the University can fall back on. But, at the present moment, Dr. Chant is making enquiries that have been brought to his attention. There is a farm close to Yonge St., about ten miles north, that the University can fall back on. But, at the present moment, Dr. Chant is making enquiries that have been brought to his attention.

Science News-Letter Washington D.C. Jan. 10, 1931

Another Huge Telescope Ordered for Canada

CANADA will soon have two of the world's three largest telescopes. This became known with the announcement to Science Service by Prof. C. A. Chant, head of the Department of Astronomy of the University of Toronto, that a 74-inch reflecting telescope has been ordered for the University from Sir Howard Grubb, Parsons and Co. in England. Among existing telescopes, only the 100-inch reflector at the Mt. Wilson Observatory, in California, is larger. The figure refers to the diameter of the great mirror which concentrates the light of the star at which it is pointed.

According to Dr. Chant's announcement, the telescope will form the main feature of the David Dunlap Observatory.

At present Canada has the world's second largest telescope. This is a 72-inch reflector at the Dominion Astrophysical Observatory, Victoria, and it will become the third largest when the Dunlap instrument is completed. Third largest at present is the 69-inch reflector at the Perkins Observatory of Ohio Wesleyan University, Delaware, Ohio. The great mirror for this telescope,

made of glass manufactured at the Bureau of Standards, Washington, is now being completed at Pittsburgh. The telescope is expected to be in operation by next summer.

Two telescopes at present share the honors for the fourth largest. The 60-inch at Mt. Wilson is used by a 60-inch at the South African station of the Harvard College Observatory.

All of these instruments will be far surpassed in size by the 200-inch reflector now being made in southern California for the California Institute of Technology. Such an undertaking is this, however, that it will probably be several years before it is completed. Another project for a huge reflector, perhaps even bigger than 200 inches, has the backing of a group of men in Miami, Florida, but as yet no actual work has been done on it.

Science News-Letter, January 10, 1931

Nature, (London Eng.) Jan. 17, 1931

JANUARY 17, 1931]

NATURE

News and

The important announcement has just been made that in the near future there will be established in the vicinity of Toronto an astronomical observatory which will rank with the world's greatest institutions of this kind. It will be erected by Mrs. D. A. Dunlap and her son, D. Moffat Dunlap, as a memorial to the late David A. Dunlap, who died on Oct. 29, 1924, and will be known as the David Dunlap Observatory. Astronomy and geology were both favourite studies of Mr. Dunlap, but the former had a peculiar attraction for him. He was a keen student of the heavens, and always liked to share his knowledge with others. This project has been under consideration for the last five years and will now be brought to completion. In working out the plans, Mrs. Dunlap has had the assistance of Prof. C. A. Chant, head of the Department of Astronomy of the University of Toronto.

The outstanding feature of the David Dunlap Observatory will be a large reflecting telescope seventy-four inches in diameter. There is only one of greater aperture in the world, namely, that on Mount Wilson in California. The instrument was ordered some time ago from the firm of Sir Howard Grubb, Parsons and Co., of Newcastle-on-Tyne. It will be housed in a circular metal building, such construction being best for this purpose. The observatory building will be a beautiful structure in the classic style. It will be erected on a suitable site near Toronto, in the midst of a large acreage which will be converted into a park, to be known as the David Dunlap Park. When the observatory is completed, it will be housed in the Department of Astronomy of the University of Toronto, for the park will be developed in a scientific way by the Faculty of Forestry. The new institution will bring distinction to the University, the city, the province, and indeed the whole Dominion. It will be an enduring memorial to a worthy citizen.

Telegraphic Address:
PHUSIS, WESTSTRAND, LONDON
(Two Words).
Telephone Number:
GERRARD 9830

BAG/FG, N.

Prof. C. A. Chant,

The University,

TORONTO, Canada.

Dear Prof. Chant,

I am very glad to have the information you have been good enough to send me relating to the Observatory which is to be established in connection with the University of Toronto by Mrs. Dunlap as a memorial to her husband. I am inserting the notes in this week's issue of "NATURE" and am sure that many readers will be interested to learn of this magnificent gift.

With all good wishes for 1931,

Very sincerely yours,

C. A. Gregory

THE JOURNAL

OF

THE ROYAL ASTRONOMICAL SOCIETY
OF CANADA

Vol. XXV, No. 1

JANUARY, 1931

Whole No. 200

A GREAT MEMORIAL

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The new institution will bring distinction to the university, the city, the province and indeed the whole dominion. It will be an enduring memorial to a worthy citizen.

Note.—The above announcement was made public on December 30, 1930. A great observatory in the midst of a fine park will be an appropriate memorial to the late Mr. Dunlap, and it will likewise be a splendid gift to the university and the community. Further information will be given as the project is carried out.

C. A. CHANT.

From the Journal of the Royal
Astronomical Society of Canada
C. A. Chant, Editor

Better from the Editor of "Nature", See printed note on page 14

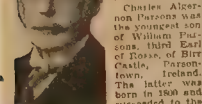
YEARS OF BRITISH TRADITION WILL ENTER NEW TELESCOPE

Regretted Sir Charles Parsons
Died Before Toronto
Reflector Made

A DIFFICULT TASK

By C. A. CHANT
Professor of Astro-Physics, University
of Toronto

The recent death of Sir Charles Parsons, while on a pleasure cruise in the West Indies reveals some interesting incidents in the history of the telescope building.



Sir Charles Parsons was the youngest son of William Parsons, third Earl of Rosse, of Birr Castle, Fermanagh, Ireland. The latter was born in 1800 and succeeded to the title on his father's death in 1841. He graduated at Oxford with first class honours in mathematics. While a youth, he showed a special taste for mechanics, and developed a remarkable manual dexterity in carrying out his ideas. He loved to visit the great industrial works which were then springing up in England. On one occasion while being conducted over a large factory in the north of England, he exhibited such an intelligent interest in what he saw that the proprietor, being in need of a foreman, offered him the position. The visitor gently declined the compliment and the incident led to a pleasant dinner and a valued friendship.

Having leisure and means, Lord Rosse looked about for the best way to employ his mechanical ability. He admired exceedingly the work of Sir William Herschel, who became world famous through his construction of great telescopes and his skill and zeal in using them. Lord Rosse realized that the construction of mighty instruments for exploring the heavens would demand much time, much wealth and the very limit of his mechanical skill and he decided to make this the chief business of his life.

There are two types of telescopes, refractors and reflectors. In the former a lens collects the light coming from the celestial object and converges it to form the image. In the latter a concave mirror performs this function. The largest telescopes are reflectors.

Now the production of the mirror of a reflector is the most difficult and delicate part of the whole procedure and Herschel had left no doubt of the way he went about it. He made the material then used for mirrors was an alloy of copper and tin, which is very hard and brittle quite unlike either of its constituents. It is hard to cast and difficult to work but it takes a fine polish and does not tarnish easily.

Was Scientific Wonder
Lord Rosse, after much experimenting, undertook to make a mirror six feet in diameter and five

inches thick—for larger than any attempted up to that time. At every stage he met immense difficulties, but he overcame them all and completed his telescope in 1845. It was one of the wonders of the scientific world. It was pictured in the papers and magazines and in all the books on astronomy, and Sir Charles became a mecca to which the faithful from every country made pilgrimages. It remained unequalled in size until about 15 years ago.

Thomas Grubb was another clever Irish mechanic, also born in 1800, at Killybegs. Early in the century he established, near Dublin, works for the manufacture of astronomical instruments, small machine tools, etc. He was original and facetious and acquired great skill in practical optics. Among the products of his shop was the original equatorial for nearly 40 magnetic stations, one of which was established in Toronto in 1840 and has been in operation ever since, though removed in 1885 to Agincourt. Lord Rosse became acquainted with Grubb and frequently sought advice from him. Grubb's last and most important production was the notable four-foot reflector for Melbourne, Australia, which was completed in 1847. Until quite recently it was the largest telescope in the southern hemisphere. The present writer had the pleasure of inspecting it in 1922.

Thomas Grubb retired from active work in 1868 and was succeeded by his son Howard, who extended the business making many instruments, large and small. In 1918 the works were moved to St. Albans in Hertfordshire, England, in connection with the manufacture of periscopes, as the firm had made a specialty of these instruments from the date of the first British submarine.

Refracted steam turbine
Sir Charles Parsons now comes into the story. He was born in 1854 and received some of his education from the able young astronomer employed by his father to observe with

the great six-foot telescope, and then he went to Cambridge, where he graduated in 1876. In his father's workshops he learned to love mechanical purity and on leaving the university, he went as an apprentice into the Armstrong Ordnance works at Elswick.

He devoted all his spare time to the study of steam and in 1884 he built a ten horse-power turbine steam engine the first practically successful engine of this kind made. It received little attention, but the inventor labored incessantly to perfect it, and in 1889 the firm of C. A. Parsons and Co. was formed to produce steam turbines and electrical equipment.

The business progressed quickly and modestly as its head was averse to spectacular advertising. However, he saw his opportunity on the occasion of Queen Victoria's Diamond Jubilee in 1897. He equipped a yacht 100 feet long and nine feet wide with turbine engines of about 2000 horse-power, and at the naval review it raced up and down the line of warships at the unprecedented speed of nearly 40 miles an hour. Nothing further was needed. The use of the steam turbine for warships and good passenger vessels is now almost universal.

The invention and perfection of the steam turbine is by far the most important advance in steam engineering since the time of Watt. Immense works have been established at Newcastle-on-Tyne for the manufacture of marine and stationary turbines and electrical machines. The great Cunard ship now being built will be propelled by turbine engines of 30,000 horse-power.

Working on Toronto Reflector
In the reaction after the war, Sir Howard Grubb's firm became involved in financial difficulties, and what could no longer be sustained.

The firm was acquired by Sir Charles in 1920 and the name was changed to Sir Howard Grubb, Parsons and Co. Large and well-equipped shops were erected beside the electrical works at St. Albans and the firm was ready to construct astronomical equipment of the largest size.

In 1921, Sir Charles acquired what is now known as the Parsons Optical Glass Works at Derby, and thus the firm can supply its own glass for all purposes. Recently a 36-inch reflector was completed for the Royal Observatory at Edinburgh, while a 40-inch reflector and a 24-inch reflector for Stockholm are almost finished. Some work has been done on a 15-inch reflector for the Soviet government of Russia.

In May, 1929, the order was given to the firm for the immense 74-inch reflector for the David Dunlap Observatory, to be erected near Toronto. About three years will be required for its construction.

Sir Charles Parsons passed away on board ship in the harbor of Kingston, Jamaica, at the age of 76. It is to be regretted that he did not live to see Toronto's mammoth telescope completed for which he built many commercial machines of larger dimensions and indeed was accustomed to carrying through great projects yet it is certain that he would have had special delight in producing a telescope carrying his father's historic masterpiece of 1845.

The Star (Weekly) (Toronto) March 31, 1931.



ROYAL YORK HOTEL
TORONTO

Canadian Pacific Hotels

May 23rd, 1931.

Mrs. D. A. Dunlap,
93 Highland Drive,
Toronto, Ont.

Dear Mrs. Dunlap:-

It was a great pleasure to me to meet you personally on Thursday night, and to express briefly then my appreciation of your magnificent gift for the advancement of Astronomy. I can assure you, for I know personally most of the eminent Astronomers of the world and their views on any addition to the facilities for prosecuting their science, that your great contribution will be highly appreciated by them and that the name of your husband and yourself will be perpetuated by such a gift and by the work done by the David Dunlap Observatory, as I think a gift to no other science could accomplish.

I recall to mind the name of James Lick who bequeathed a large sum to establish an observatory near San Francisco nearly fifty years ago. Lick was a contractor of San Francisco who would certainly have been forgotten long ago if it had not been for this gift which, wisely administered, has made the name of the Lick Observatory perhaps the most celebrated in the world. I feel sure that your generous gift, under the wise management of Professor Chant and the University, will make the name of your husband widely known, not only to all Astronomers, but to the whole world of science and culture. I wish to thank and congratulate you again as a Canadian Astronomer for your generous gift and very wise memorial.

*Letter from Dr. J. S. Plaskett
regarding the site chosen
for the Observatory*

- 2 -

I had the pleasure this morning of going with Professor Chant to the proposed site of the Observatory, and I was very favourably impressed with its special suitability for the purpose. The position on an eminence sloping away in every direction is particularly desirable, as the air drainage thereby produced tends to improve the "seeing" conditions materially, as well as reduce the daily range of temperature variations, also an important factor. I did not see in our journey to the spot any location at all approaching in desirability the one chosen. I believe however, it would be desirable to secure some additional land to the north so as to control the whole summit and approaches to it for at least a quarter of a mile in every direction.

May I wish you sincerely much pleasure and happiness in watching the progress of this wonderful memorial, and in following after completion the valuable work it will undoubtedly produce.

Yours sincerely,

J. S. Plaskett

Director,
Dominion Astrophysical Observatory,
Victoria, B. C.

THE ROYAL SOCIETY OF CANADA PROCEEDINGS FOR 1931

RESOLUTION REGARDING THE DAVID DUNLAP OBSERVATORY

It was moved by Dr. Plaskett, seconded by Dr. King, that the members of the Royal Society of Canada desire to express their appreciation of the notable contribution to science which is being made by Mrs. D. A. Dunlap in undertaking the erection of a well-equipped astronomical observatory, which will supply additional facilities for research to the University of Toronto.

They recognize that this great observatory will be an appropriate and enduring memorial to the donor's late husband; they also believe that in the years to come it will lead to a large increase in our knowledge of the universe and will bring distinction to our country. Carried.

Resolution passed May 22, 1931

An Astronomical Observatory for Toronto



An outline of a project prepared by a committee
representing the University of Toronto and the
Royal Astronomical Society of Canada



UNIVERSITY OF TORONTO PRESS



201 Madison Ave.,
April 28 1932

Dear Mrs. Dunlop, -

Herewith is the leaflet which I spoke
to you about last evening. It was prepared
in 1921 and describes the scheme we had
at that time.

You will find the news I have
regarding similar to the people. The statistics
are ten years, or more, old, but the flocking
of the people to a telescope still continues.
In 1931 the number of visitors to the observatory
at Victoria, B.C., was 24,540, which was fewer
than in previous years.

Very sincerely yours

"a (hand)"



201 Madison Ave.,
May 23, 1932

Dear Mrs. Dunlop, -

Herewith are three prints from
the films exposed the other day. You
can tell me what you think of them
when next I see you. I enclose also
a picture of Prof. & Mrs. de Sitter, which
I thought you might like. It was taken
at Dr. Burton's home at Weston.

Very sincerely yours

"a (hand)"

P.S. Hope this may catch
you before you get to
the olden



At 93 Highland Ave., May 19, 1932



Professor & Mrs. W. de Sitter of Leyden Holland
Taken in Dr. Burton's garden Nov 1931

University of Toronto
TORONTO 8
DEPARTMENT OF ASTRONOMY



Mrs. D. A. Dunlop,
93 Highland Ave.,
Toronto 5

HOTELS OF DISTINCTION
CANADIAN NATIONAL RAILWAYSCHATEAU LAURIER
OTTAWA

OTTAWA, ONT.

May 26, 1932

Dear Mrs. Dunlop, —

I am attending the meeting of the Royal Society of Canada. Sir Robert Falconer is president and gives his address this evening.

At noon Premier Bennett gave the members (about 125) a luncheon at The Forestry Club, and near me at the table was J. MacKenzie Bell, a geologist, who was much interested in our observatory project, and who tries to claim relationship to you. He says there are some people to whom you wish to consider yourself related. He is a charming person anyway.

I understand that extra land has been definitely secured.

I return on Saturday afternoon. Will address this to Highland Ave.

Very sincerely yours
C.A. Chant

Lo-Hone Bay,
July 20, 1932

Dear Mrs. Dunlop, —

I was pleased to learn that you had the children with you again, and that the excitement had not proved too much for you.

We came up here on Saturday, June 25, and have had cool weather much of the time since then. Mr. Holden told me that the family would migrate to Cranby Bay on Tuesday, June 28. By the way, on June 24 I got some photographs of him, several of which Mrs. Holden thinks good. I shall send you three at the first opportunity. The advertisement calling for tenders for the Administration Building had appeared before I left and the tenders were to be received up to July 11. I have not read anything further. The Property Committee of the Board of Governors was authorized to award the tender, but this was not to be done until Mr. Holden had been

THE BOARD OF GOVERNORS.
OFFICE OF THE CHAIRMAN,
REV. CANON CODY, D.D., LL.D.



603 Jarvis St.

TORONTO, 17th June 1932

Dear Mrs. Dunlop:

The Board of the University has sent to Mr. Holden the official acknowledgment of his letter authorizing the Board to call for tenders for the Administration Building of the Dunlop Memorial Observatory. May Todd, as Chairman of the Board, our deep appreciation of your splendid generosity in this whole matter? On Wednesday morning I visited the site near Richmond Hill in company with Professor Chant and Col. MacParr. It is magnificent. There is a natural elevation which gives an unobstructed view in all directions. The Building will have a commanding position and the park which will be created around them



Mrs. D.A. Dunlop,
Highland Ave.,
Pore Dale
Toronto

HOLDEN, MURDOCH, WALTON & BEATTY

JOHN B. HOLDEN, K.C.
JAMES Y. MURDOCH, K.C.
WILLIAM H. BEATTY
WILLIAM S. WALTON
P. C. FINLAY, B.A.
S. H. ROBINSON, B.A.

BARRISTERS & SOLICITORS
SUITE 603 4 ROYAL BANK BUILDING
20 KING STREET EAST

TELEPHONE ELGIN 91
CABLE ADDRESS: BARTLAW, TORONTO
ADDRESSES ALL CORRESPONDENCE TO THE FIRM

TORONTO 2

August 9th, 1932.

Mrs. Jessie B. Dunlap,
Mattawa,
Ont.

Dear Mrs. Dunlap:

RE OBSERVATORY.

Mr. Holden has asked us to write you and give you a report on the progress in this matter since your departure.

The deeds have now been executed by the University of Toronto and same will be registered tomorrow. The contract for the work will also be signed tomorrow.

We have also this day paid to the University of Toronto the sum of \$50,000 and delivered to the University of Toronto as security for the additional payment, bonds totalling \$90,000. We will not go into details in respect to the payments and bonds as we feel that you are quite familiar with all the facts.

The University will now be in a position to proceed at once with the work and Dr. Moure mentioned to the writer the fact that arrangements had been made with you and Mr. Holden that no information would be given out to the newspapers except with your approval. He also furnished us with several photostats showing the proposed Great Telescope and Administration Building and three pages of typewritten material which he states Professor Chant proposed giving to the newspapers. We understand that Professor Chant collaborated with you in the preparation of this material to be given to the newspapers, and accordingly we are forwarding same to you together with one copy of the photostat in order that you may consider same and let us know if it meets with your approval. Would you please let us have a reply by return mail.

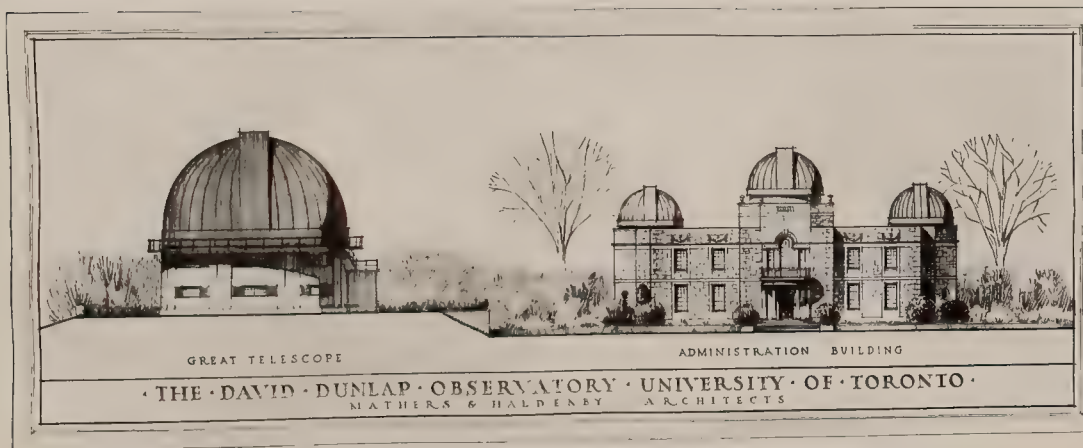
Trusting you are having an enjoyable vacation, we remain

Yours very truly,

HOLDEN, MURDOCH, WALTON & BEATTY.

Per: *[Signature]*

PCB/E
ENCLOS.



Reduced Copy of Photostat mentioned in above Letter

THE DAVID DUNLAP OBSERVATORY

In one of the accompanying illustrations is presented the western elevation of the two chief buildings of the David Dunlap Observatory, the erection of which will be commenced soon. The Observatory is being provided as a memorial to the late David A. Dunlap by his widow, Mrs. Jessie D. Dunlap, with whom is associated their son, Mr. Moffat Dunlap. When completed it will be presented to the University of Toronto and will be conducted by the Department of Astronomy.

At the left of the picture is a circular building 61 feet in diameter, surmounted by a hemispherical dome, which can be revolved by an electric motor. The frame work is of steel. The walls are 2 feet thick and hollow, being sheathed on the inside and outside with galvanized iron. The inner and outer surfaces of the dome, also 2 feet apart, will be covered with a special kind of papiermache, and over the outer surface will be put a layer of sheet copper for protection against the weather.

This building will house a huge telescope which will be next to the largest in the world. It is of the reflecting type, that is, in place of a lens it will use a concave mirror to collect the rays of light from a star and bring them to a focus. This mirror will have a clear aperture of 74 inches and will be fashioned from a disc of glass 76 inches in diameter and 12 inches thick and weighing about 2-1/2 tons.

The telescope was ordered upwards of 2 years ago from the firm of Sir Howard Grubb, Parsons and Company, Newcastle-on-Tyne, England, and the construction is well advanced. It was estimated that the great instrument would be finished in 3 years but the final figuring of the mirror may require a longer time. The casting of such an immense block of glass, the long process of annealing, and then the grinding and polishing of its surface correct to half a millionth of an inch, sometimes takes longer than anticipated; but every effort is being made to avoid delay.

Typed Document mentioned on opposite page

continued on next page

-2-

The building is also being supplied by the same firm which is a subsidiary of Charles A. Parsons and Company, makers of steam turbines used all over the world on land and sea. It will be shipped to Toronto, it is expected, about a year hence, after it has been set up and tested in England. This summer, however, it is intended to put in the cement foundation of the building and to erect the immense pier which will carry the 50 ton telescope.

At the right of the picture is shown the Administration Building. The plans for it have been prepared by Mathers and Haldenby, Architects, Toronto. It will contain offices, library, lecture room, reception room, laboratories, computing rooms and workshop. According to the plans it will be 91 feet long by 49 feet wide, with semi-octagonal projections on each end. On the roof are three domes, those at the ends being 21 feet, that at the centre 25 feet, in diameter. In one of the smaller domes it is proposed to mount the 19-inch reflecting telescope recently constructed at the University of Toronto by Professor R. K. Young, while, in the other small one, there will be 3 astronomical cameras on a single mounting. The chief parts of these cameras are already on hand and the mounting will be made in the Observatory workshop. In the central dome it is hoped to mount a 10-inch telescope of the refracting (or ordinary) type. This is to be used for the observation of planets, comets, occultations, and double stars; and it will also be available for public purposes.

According to the specifications the Administration Building will be constructed from Credit Valley limestone with trimmings of Queenston or Indiana stone. The plans show a beautiful square entrance hall and stairway to be finished in marble.

In the other picture is shown the proposed ground plan of the buildings. The site for them is in the midst of a 177-acre plot of land about a mile south of Richmond Hill and 1/2 a mile east of Yonge Street. This large acreage, now farmland, is to be made into a park to be known as the David Dunlap Park. It will be developed in a scientific way by the Faculty of Forestry and, in the course of years,

should yield results of interest and value, besides providing a pleasant resort for the people. The dotted curves are contour lines which show the elevations above sea level. It is proposed to grade the land so that the round building will be on a circular platform 800 feet above sea level; while about 5 feet lower will be the area on which the Administration Building will be placed, and, in front of it, will be an area 3 feet lower still. These three areas will be bordered by hedges.

The great Dominion Government Observatory near Victoria, B.C., whose telescope is 72 inches in diameter, has been in operation 14 years, and has become one of the world's outstanding institutions of research. Every year it attracts scientific workers from distant lands and some 30,000 visitors make a pilgrimage to it. The David Dunlap Observatory hopes, also, to make contributions of permanent value to astronomical science and to bring distinction to the University and the Country.

Copy

September 17, 1931

Dear Mr. Holden.

Last evening I went out to Don-Alda Farm and had a long conversation with Mrs. Dunlap and Moffat. I showed them about a dozen photographs of the new telescope, which afford a visual proof of the progress outlined in the last quarterly statement from the firm. I should like to show them to you if it should be convenient and you would wish it.

We also discussed what could be done to help along our project, and though we realized that the actual construction is a considerable way off, we all hoped that the ceremony of turning the sod or laying the corner stone might take place before Sir Robert Falconer retires from the presidency (July 1, 1932).

I mentioned the proposal which Mrs. Dunlap had made some time ago, namely, that the University be asked to undertake the construction. She thinks this would relieve her of much worry; and I assured her that if this arrangement was made the University would do nothing of importance without her full approval. Supt. LePan has the staff and they have had the experience which would allow the work to be carried on efficiently and without any "extras".

Then we talked about an architect. Mrs. Dunlap did not wish to name anyone. Moffat suggested Paisley, who was architect for St. Andrew's College, but his mother and I felt inclined to accept LePan's responsible suggestion of the firm of Mathers and Haldenby, which he had found very satisfactory. Even though construction cannot begin for a long time, it seems to me that the only way to make progress is to choose the architect and thus have some authoritative person to

h. Months will be required before the plans are ready. I expect Young and I may have to visit some architect to get suggestions regarding the best construction. Mrs. Dunlap shares my views on this. I think any unreasonable obligation will be in-
tact, even though construction is delayed, objection to the proposal.

You can take time to look at the photographs I have enclosed. I have asked to take them to you, either in the form of a letter or by mail. I am sorry to hear of your illness and hope you will soon be able to return to your usual work. It is my sincere hope that you will soon be able to return to your usual work.

Very sincerely yours

J. A. Clark

Toronto
Ontario

Mrs. D. A. Dunlap,
Don-Alda Farm,
T O D M O R D E N,
Ont.

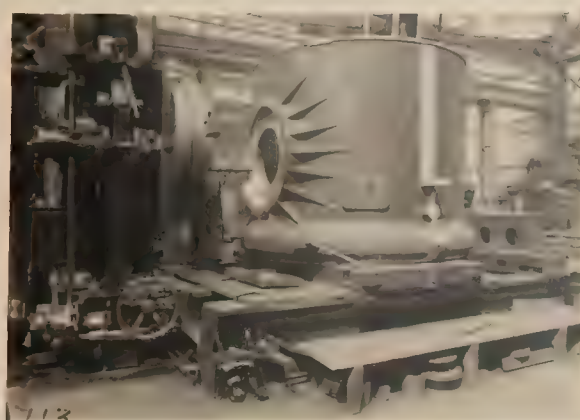


Copy of letter to Mr. HOLDEN discussing Architects and the Progress of the Telescope - 3
In the pages following are Photographs illustrating the construction of the 74-inch Telescope



No. 12

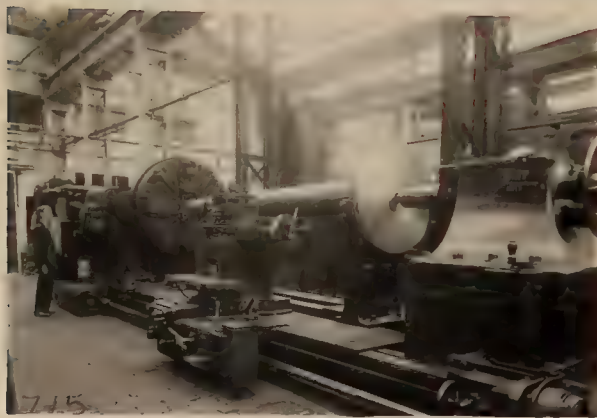
Boring out the Centrepiece of the Tube. April 1931



No. 13

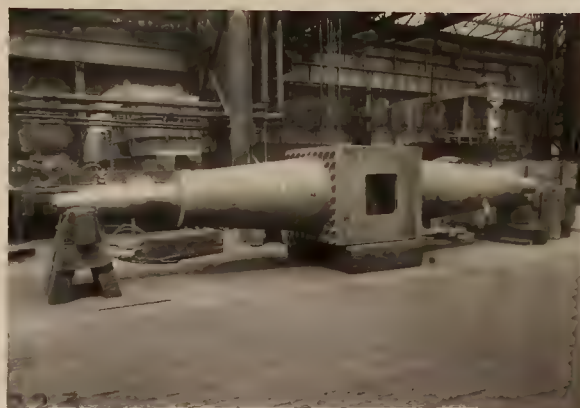
Machining the facing on the Centrepiece in the Kerr's Machine May 1931

Outside diameter, 8 ft. 7 in., Length, 6 ft. 9 in., Weight $4\frac{1}{2}$ tons



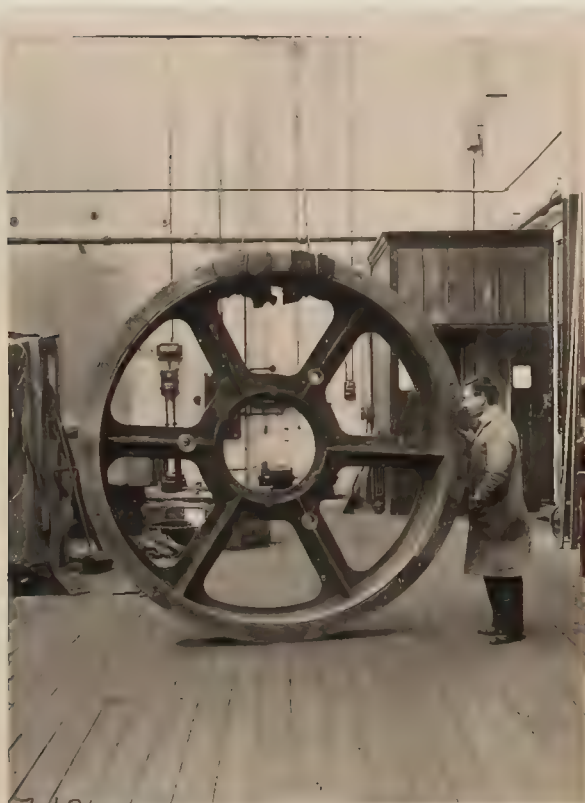
No. 115

Turning the Declination Axis - May 1931



No. 123

The Polar Axis assembled in the Shops - July 1931



No. 118

Cast Steel Cell for the Main Mirror - June 1930



No. 125

The Declination Axis fitted to the Centrepiece July 1931



No. 736.

Boring the Bush of the Sidereal Circle - August 1931



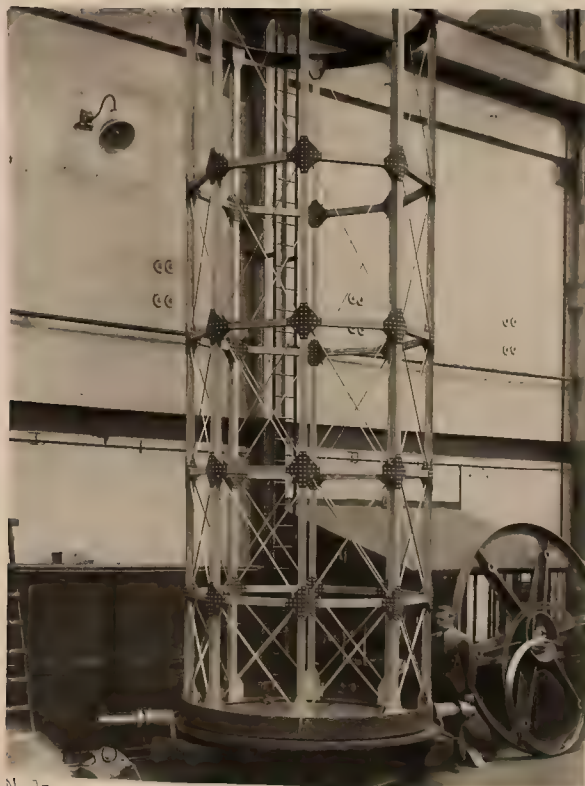
No. 754

Model of the 7 1/4 inch Telescope and Dome November 1931



No. 158.

Constructing the lattice Tube - August 1931 (Note size from Workman)



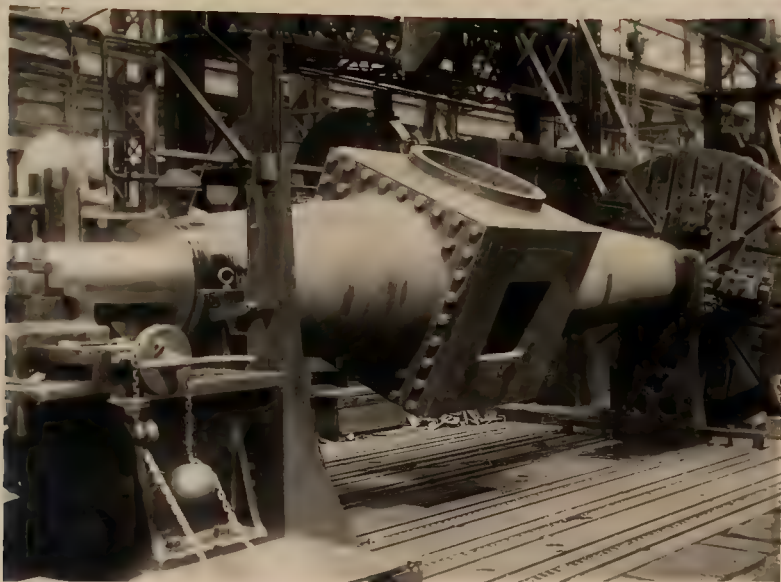
No. 750

The Completed Tube, Sidereal Circle at Right October 1931



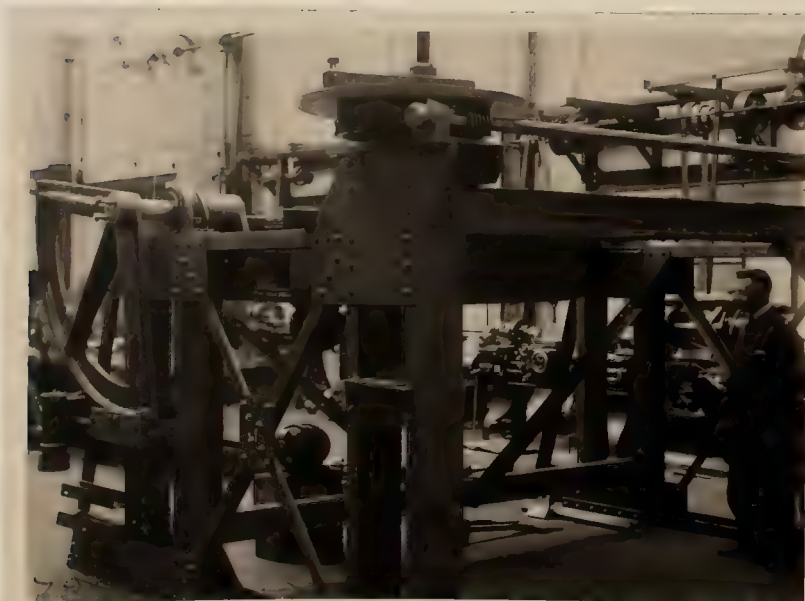
No. 751.

The Tube Mounted on Centrepiece - October 1931



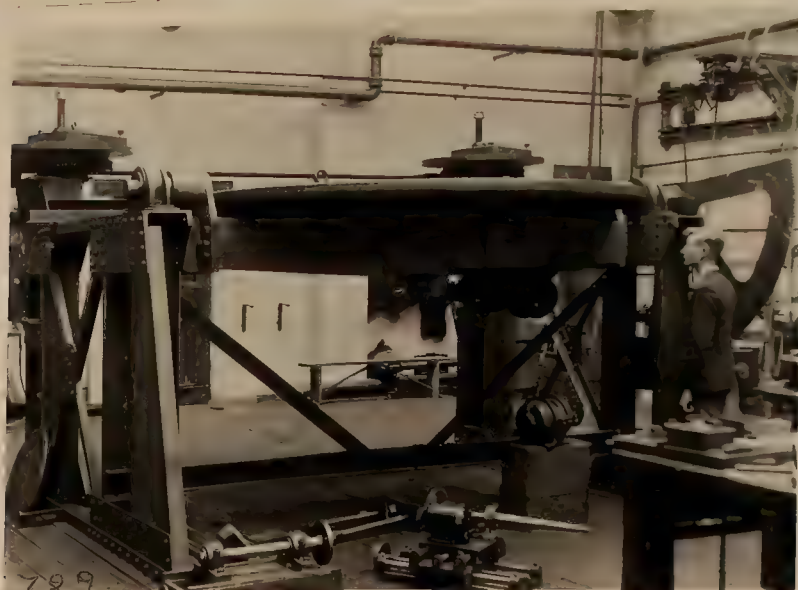
No. 127.

*Turning the Polar Axis, Weight about 10 tons July 1931
(See page 22)*



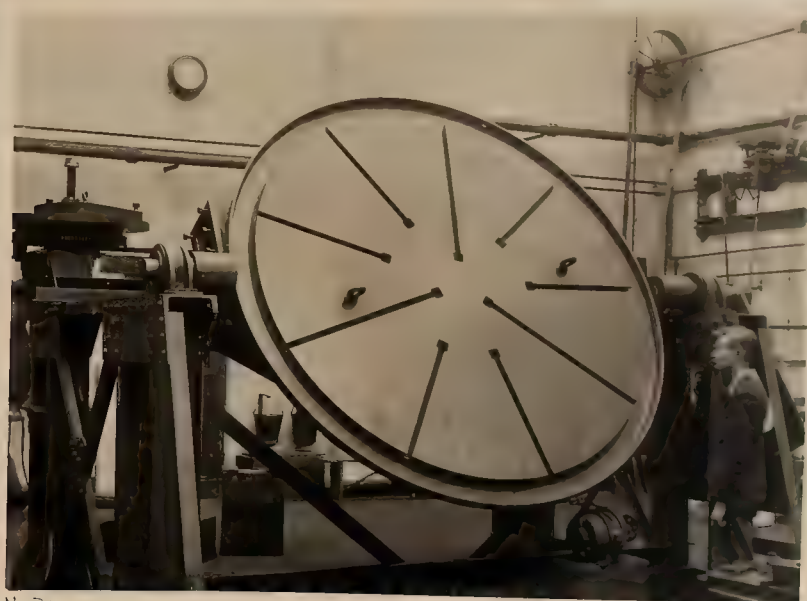
No. 128.

*The Machine for grinding, polishing and testing the great
Mirror, Rear View showing the cranks August 1931*



729
No 729

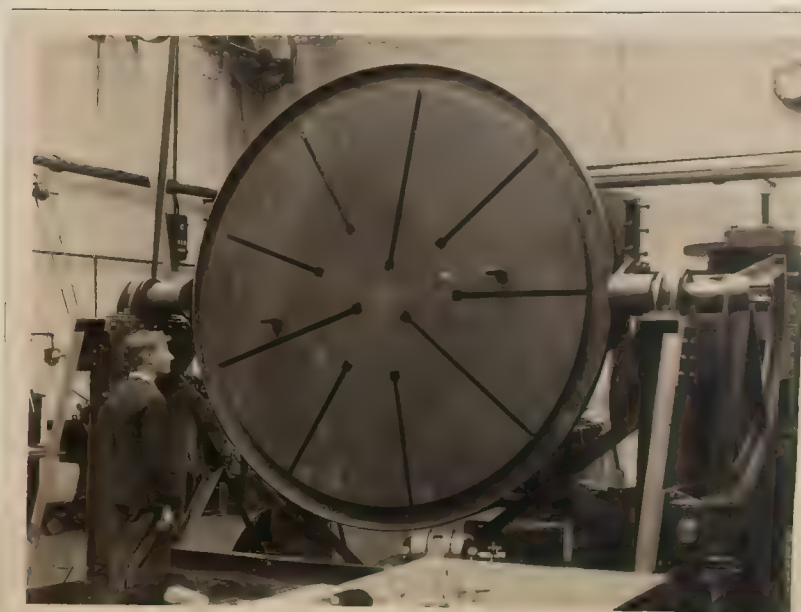
*Machine for working the
Great Mirror, view from front,
table horizontal - August 1931*



No. 730

*The same machine
with table tilted -
August 1931*

*The same machine, with table
vertical, - August 1931*



No. 731

Edinburgh Telescope



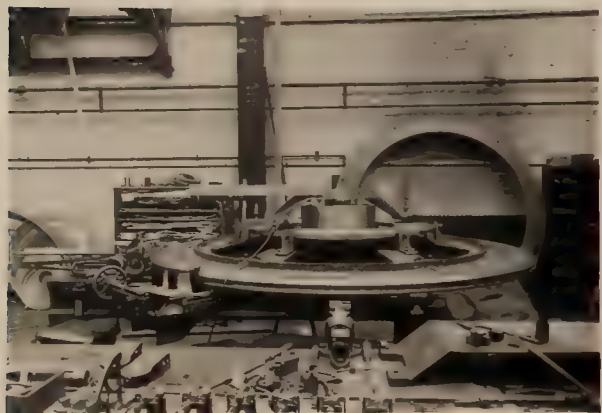
No. 76.

Turning the Driving Circle in the bath - October 1931



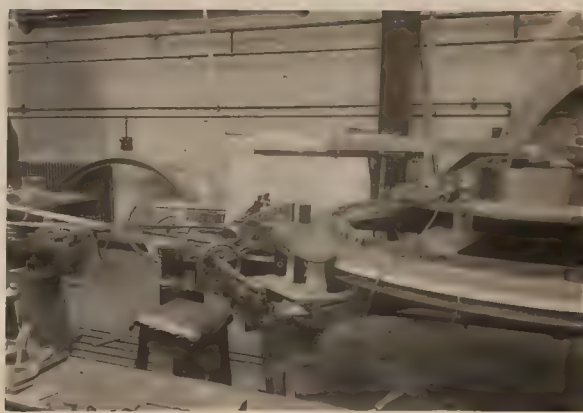
No. 786

Gashing the Driving Circle, view from behind cutter gear - February 1932



No 78.

Gashing the Driving Circle, side view February 1932



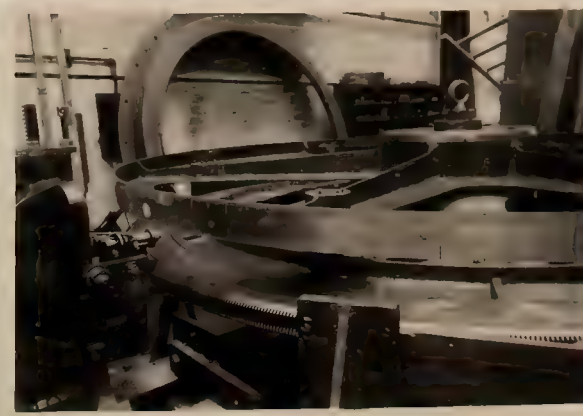
No 788

Gashing the Driving Circle, close up view showing the hob - February 1932



No 802

Graduating the Sidereal Circle, using the driving circle as master - February 1932



No 804

Graduating the Sidereal Circle, near view February 1932



No 822.

Polar Axis complete with circles, ready for hoisting into place, close up view ~ May 1932



No 824.

Polar Axis, ready for hoisting into place; view from S.E. showing shear legs, May 1932



No 825

Polar Axis being hoisted into place, the lift just commencing ~ May 1932



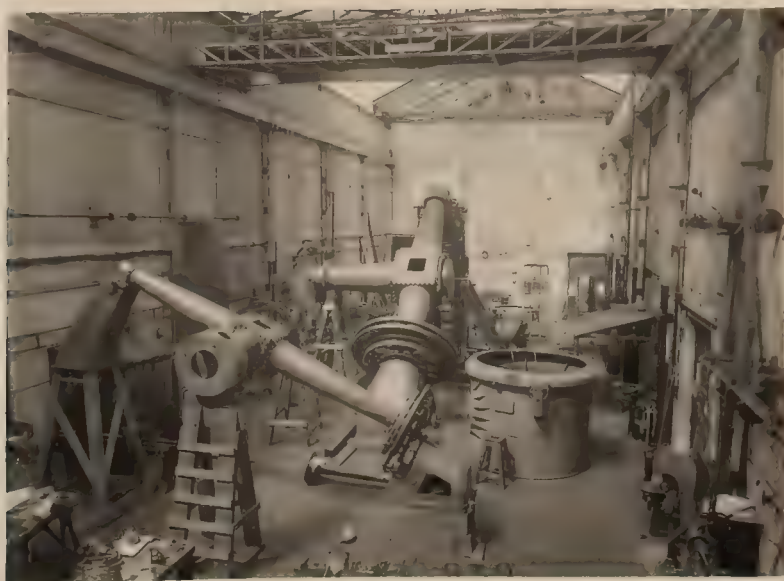
No. 827

*Hoisting Polar Axis into place, top ready
for lowering - May 1932*



No. 828

*Hoisting Polar Axis into place; lowering
the Axis - May 1932*



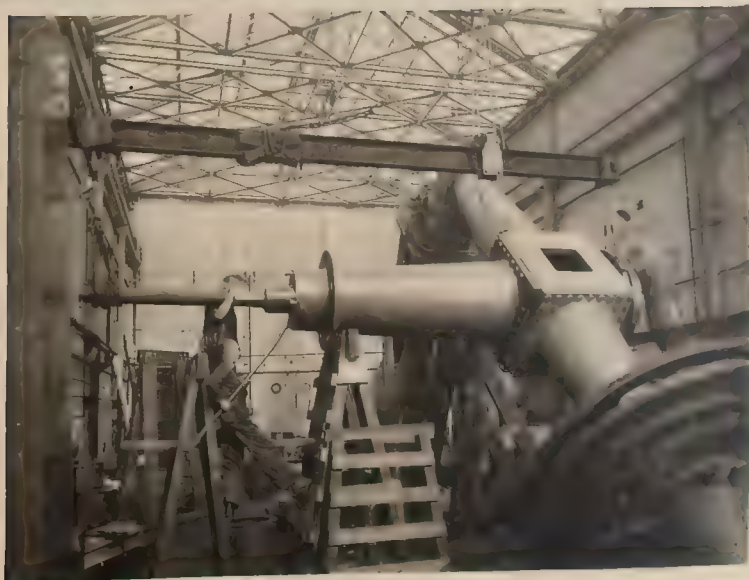
No. 829

*General view in shops, showing polar axis in place and declination axis fitted, centrepiece
at right. The polar axis on its stand in front is for the twin 16-inch astrographic telescope
for Bryden University - June 1932*



No 849

Inserting Declination Axis; axis entered - August 1932



No 851

Inserting Declination Axis; axis home - August 1932



*Moffat Dunlap, C.A. Chant, Mrs. Dunlap.
Eclipse Camp, St. Alexis des Monts, Quebec August 31, 1932*

While the construction of the Telescope and dome was proceeding in England preparations for the erection of the buildings were going forward at home.

★ ★ ★

From the Journal of the Royal Astronomical Society of Canada, July-August issue
Page 277.

THE DAVID DUNLAP OBSERVATORY

This great project is going forward steadily. The 74-inch reflector was ordered in May, 1930, and its mounting is well advanced towards completion. The contract for the 61-foot dome to cover it was given in November, 1931, and rapid progress is being made on it. Tenders for the erection of the Administration Building have been invited by advertisement in the press and construction will proceed actively during the summer.

★ ★ ★

There was a total eclipse of the sun on August 31, and an expedition had gone from the University of Toronto to St. Alexis des Monts P.Q. to observe it.

St. Alexis des Monts, P.Q.,

August 14, 1932

Dear Mrs. Dunlap,

I thought I would let you know how our eclipse preparations are proceeding.

Dr. Young and I drove from Toronto on Aug. 1 & 2, arriving here at 4 p.m. We stopped over at Cornwall & the road was pleasant.

On arrival we went to Hotel Boulanger where we were expected. It is a modest place, not many fills, but we are thriving. The work is going well. Dr. Young's brother Albert & son (a young man) came a few days earlier. The family rented a cottage at 'La Cache', 5 miles away. Then Mr. & Mrs. Fleming arrived by motor on Sunday afternoon. The men are excellent workers and the apparatus and their houses are almost erected. Quite a show!

On the 23rd three more men came to rehearse the programme & on the 25th Mrs. Chant comes by train. They will all be at the hotel. Immediately

From C.A. Chant
St. Alexis des Monts,
P.Q.



Mrs. D. A. Dunlap,

Matthewa P.O.,

Ont.

St. Alexis des Monts, P.Q.

August 13, 1932

Dear Mrs. Dunlap,

Just a bit more information about this place. Yesterday I had the opportunity to visit a little more about three miles from the village and found there an excellent summer hotel built just a year ago. It is called Le Glacé. In my last letter I was not rather drastic about accommodation around here but if you can make I am content to stop over here there will be no lack of a pleasant place to stay at and the rates are very reasonable too.

If you can make it by St. Paulin this morning or tomorrow would not you and bring you up. Son and daughter will also

President's Office.

The Hon. and Rev. B. J. Gody, B.D., T.T.D.

President



August 14, 1932

Mrs. D. A. Dunlap,
55 Richmond St. W.,
Toronto.

Dear Mrs. Dunlap:

I am glad to hear that the contractor is well under way with the Administration Building. We feel sure that it will be a fine addition to the observatory. There should be a formal laying of the corner stone at an early date before the work is completed. I hope that it will be a success. I note that it will be a fine addition to the observatory. I hope that it will be a success. I note that it will be a fine addition to the observatory. I hope that it will be a success.

Our children's debts are being paid every time. I am glad to hear that the contractor is well under way with the Administration Building. We feel sure that it will be a fine addition to the observatory. There should be a formal laying of the corner stone at an early date before the work is completed. I hope that it will be a success. I note that it will be a fine addition to the observatory. I hope that it will be a success.

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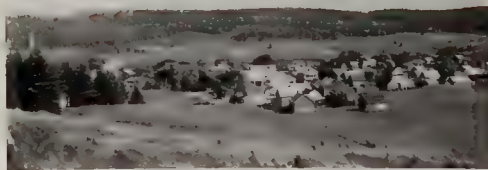
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H. J. Gody

President

Eclipse Camp, St. Alexis des Monts.

Mrs. Dunlap visited the eclipse camp, having driven from Montreal, accompanied by Moffat and Mr. Fromings.



St. Alexis from Eclipse Hill, looking east.



The Eclipse Camp, Aug. 11.



The 40-foot Camera, Aug. 18.



Mrs. Dunlap and Moffat, in front of "Chateau Einstein"



*Jean Young Counting Seconds
(Professor Burton behind)*



*Dr. Young explaining the Movie Outfit.
Dr. Young, Mr. Fromings, Mrs. Dunlap, Mrs. Chant, Moffat.*



Mrs. Chant and her Shadow Bands Sheet. Aug. 31



*F.W. Burton, Mrs. Dunlap, Moffat, Professor Burton, Mr. Fromings
and Mrs. L.V. King, beside the sheet. Aug. 31.*



Nov. 12, 1932

Dear Mrs. Dunlap, —

Herewith are prints from
the negatives you chose. An
enlargement of that one of
Mr. Holden in his library will be
ready shortly.

Very sincerely yours

C. A. Chant

see
opposite
page.

P.S. I was at the Observatory
this morning (Saturday)
and got another roll
of films. C.A.C.



In Library.

Mr. John Holden, at 7 Thornwood Rd. Toronto, June 24, 1932.



At Sun Porch.

INTERNATIONAL ASTRONOMICAL UNION

FOURTH GENERAL ASSEMBLY

CAMBRIDGE, U.S.A.

SEPTEMBER 2-9, 1932

September 5, 1932

Dear Mrs. Dunlap, —

I suppose you reached Montreal safely on
Wednesday evening and then went on to Toronto. Mrs.
Chant and I left St. Alexis on Thursday at 2 p.m.
and drove to Montreal with Mr. & Mrs. Herring. At
8¹⁰ p.m. we boarded the train for Boston and reached
this place at 9 a.m. Friday morning. We are stopping
in Eliot Hall, one of the Radcliffe College dormi-
tories, and a pleasant place it is. Radcliffe, as
you know, is the women's college in connection
with Harvard.

There is a large attendance of members of the
I.A.U. I posted to you the program of the meeting
as well as a booklet describing the Harvard Observatory.
I have met many of my old friends & made
some new ones.

Yesterday afternoon we drove to Oak Ridge,
25 mi. N.W., to the new station of Harvard Observatory,
and had a picnic supper in the woods — but
it was not as satisfactory a picnic as we
have on the rocks of Georgian Bay. The "corner
brick" was laid by Sir Frank Dyson, a copper
(one)

HARVARD COLLEGE OBSERVATORY
CAMBRIDGE

Dr. and Mrs. Harlow Shapiro
would be happy to have you attend
the Garden Party at the Observatory
on Saturday afternoon, September
third, from four to six, and
join the excursion of the
International Astronomical Union
to the Observatory station on
Oak Ridge on Sunday afternoon,
September fourth.

To Mrs. J. J. Dunlap
93 Highland Avenue
Toronto, Ontario

The Journal of the Royal Astronomical Society of Canada. September, 1932

THE DAVID DUNLAP OBSERVATORY

By C. A. CHANT

(With Plate IV and V)

IN Plate IV is shown the western elevation of the two chief buildings of the David Dunlap Observatory, the erection of which will commence immediately. As already stated in this JOURNAL, it is being provided as a memorial to the late David A. Dunlap by his widow, Mrs. Jessie D. Dunlap, with whom is associated their son, Mr. Moffat Dunlap. When complete it will be presented to the University of Toronto, and will be conducted by the Department of Astronomy.

The circular building at the left, surmounted by a dome, is 61 feet in diameter. The framework is of steel with hollow walls two feet thick, sheathed inside and outside with galvanized iron. The inner and outer surfaces of the dome will be covered with a special kind of papiermache, with sheet copper on the outside. This building will house the 74-inch reflecting telescope which was ordered in May, 1930, from Sir Howard Grubb, Parsons and Co., of Newcastle-on-Tyne, England. The mounting is almost completed, but the date when the large mirror will be ready is uncertain. Every effort is being made to prevent undue delay. The dome is being supplied by the same firm and is almost completed.

The Administration Building is shown at the right. The plans

for it have been prepared by Mathers and Haldenby, Architects, Toronto, and the contract for its erection has been awarded to Sullivan and Fried, a Toronto company. It will contain offices, library, lecture room, reception room, laboratories, computing rooms and workshop. It will be 91 feet long by 49 feet wide, with a semi-octagonal projection on each end. The two end domes are 21 feet, the centre one 25 feet, in diameter. In one of the smaller domes will be mounted the 19-inch reflecting telescope constructed at the University by Prof. R. K. Young, while in the other will be three astronomical cameras on a single mounting. The chief parts of these cameras are already on hand and the mounting will be made in the observatory workshop. In the central dome it is intended to place a 10-inch refractor, which will be used for the observation of planets, comets, occultations, and double stars, and will also be available for public purposes.

The Administration Building will be constructed of Credit Valley limestone with trimmings of Queenston stone, while the square entrance hall and stairway will be finished in travertine.

In Plate V is shown the ground plan of the buildings. They will be located in the midst of a 177-acre plot of land near the village of Richmond Hill, about twelve miles north of Toronto. This large acreage, now farmland, is to be made into a park to be known as the David Dunlap Park. It will be developed in a scientific way by the Faculty of Forestry. The dotted contour lines show the elevation above sea-level. The great dome will be on a platform 800 feet above the sea, on an area five feet lower the Administration Building will be placed, and in front of the latter will be an area three feet lower still. These areas will be bordered with hedges, and decorated with shrubs and flowers.

PLATE IV

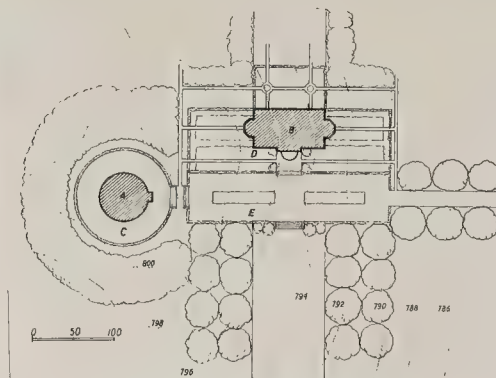


THE DAVID DUNLAP OBSERVATORY

The above picture, from a drawing supplied by the architect, shows the western elevation of the David Dunlap Observatory, construction of which will begin immediately. The dome of the left is 61 feet in diameter and will cover the 74-inch reflector. The Administration Building on the right will be 91 feet long by 49 feet wide with semi-octagonal projection at the ends. The building will be erected with Credit Valley limestone. In one of the smaller domes will be mounted a 19-inch reflector, in the other a battery of photographic telescopes, and in the centre dome will be a 10-inch refractor.

Journal of the Royal Astronomical Society of Canada 1932

PLATE V.



GROUND PLAN OF THE DAVID DUNLAP OBSERVATORY

A is the 61 feet building for the great reflector, on a platform C, which is 800 feet above sea-level. On the other side, 796 feet in elevation, is the Administration Building B. The area D has an elevation of 792 feet. The dotted contour lines show the present elevations. The hill falls away in all directions, the summit being about 75 feet above the surrounding country.

The President and Governors of the University of Toronto request the honour of your presence at the laying of the corner-stone of the

Administration Building of the Dunlap Observatory

on Saturday, the tenth of September, at three o'clock, by Mr. David Moffat Dunlap.

Through the kindness of Mrs. Dunlap tea will be served after the ceremony.

THE SITE IS SOUTH OF RICHMOND HILL ON THE EAST SIDE OF YONGE STREET

Invitation Card to the Laying of the Corner-Stone.

Sep. 10 - 3 p.m.

Opening remarks.

Prof. Chant

Premier Henry

M. Ferguson

Sir Wm. Mulock

Mayor of Toronto

Laying Stone

Prayer (Dr. Roberts)

Order of Proceedings in President Cody's Handwriting

At the Laying of the Corner-Stone - September 10, 1932



President Cody Addressing the Gathering.



2. President Cody
4. Sir W. Mulock
5. Mrs. John S. Moore
7. Miss Montgomery
8. Mrs. Frank Cochran
9. Mrs. T. McHenry
10. Mrs. D. A. Dunlap
11. Dr. D. B. Macdonald

12. Mr. Justice Kelly
13. Mrs. S. H. Ferguson
14. C. A. Chant
15. Mrs. S. S. Henry
16. Mrs. H. D. Warren
17. Mrs. D. M. Dunlap
18. Premier S. S. Henry
19. Mrs. F. N. S. Stiller



1. E. W. Haldenby
3. President Cody
6. Frank Cochran
7. Alice Cochran
8. Dr. D. B. Macdonald
9. John S. Moore
10. Sir F. Stupart
12. Mrs. D. A. Dunlap

13. Mrs. J. F. Ingles
16. Miss Montgomery
17. Mrs. S. H. Ferguson
18. Mrs. T. McHenry
19. Mrs. W. D. Ross
20. Mrs. S. S. Henry
21. Mr. Justice Kelly
22. C. A. Chant



Mrs. Dunlap presenting the Deed to Dr. Bruce Macdonald.



The Corner-Stone (with Copper Box on it) before being laid.

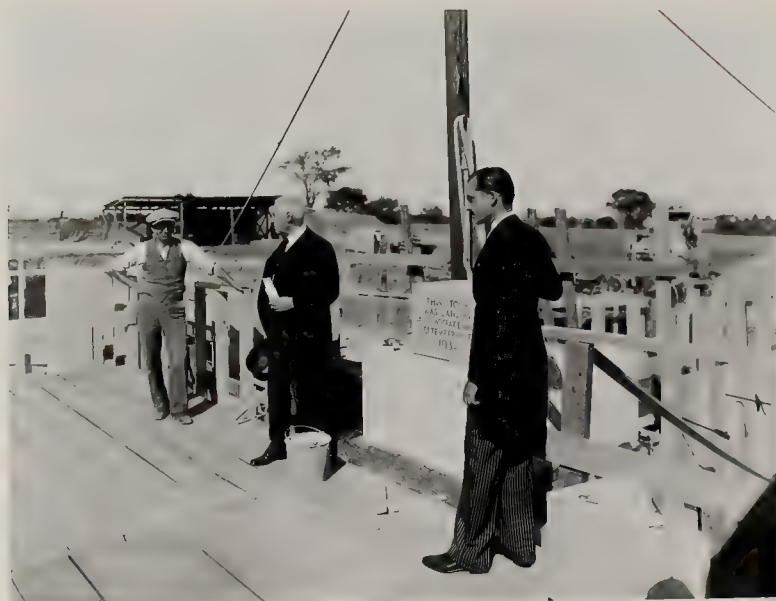


President Cody

David Moffat Dunlap E.W. Haldenby

Mr. E.W. Haldenby presented a silver-Trowel to Moffat Dunlap who laid the Corner-stone and said.

To the glory of God
and in honoured memory
of My Father,
David Alexander Dunlap,
I lay this corner-stone.



Wilson
(Stonecutter)

President Cody

D.M. Dunlap



Premier Henry Addressing the Sathering.



1. President Cody
2. Dr. D. B. Macdonald
3. Premier Henry
4. Dr. W. A. Parks
5. Alice Cochran
6. John S. Moore
7. Sir W. Mulock
8. Mrs. Edward Dunlap
9. Mrs. John S. Moore
10. Mrs. Frank Cochran

11. Mr. Mac Donell
12. Miss Montgomery
13. Mrs. D. A. Dunlap
14. Mrs. D. B. Macdonald
15. Mrs. T. Mc Kenzie
16. Mrs. S. H. Ferguson
17. Mr. Justice Kelly
18. C. A. Chant
19. Mrs. W. D. Ross



Sir William Mulock Speaking.



1. Mr. Moffat Dunlop
2. President Cody
3. Sir W. Mulock
4. Dr. D. B. Macdonald

5. Miss Ingle
6. S. H. Ferguson
7. Miss Mary Dunlop
8. W. J. Dunlop



Mrs. Dunlop

Mrs. S. H. Ferguson



After the Ceremony - going to the Refreshment Tent.



1. A. D. Lee
2. Mrs. Mayis Budd
3. Mrs. Williams
4. Doris Dobson
5. Mrs. H. D. Warren
6. Dr. F. N. S. Sturge
7. Mrs. D. A. Dunlop
8. Mrs. S. H. Ferguson

9. S. H. Ferguson
10. W. A. Wright
11. Dr. R. H. Hays
12. Dr. D. B. Macdonald
13. Rev. L. S. Stewart
14. Dr. F. W. Macdonald
15. Mrs. S. H. Ferguson

Trees A, B, C were later removed for the site of a new building.

81 GLENGOWAN ROAD
LAWRENCE PARK

Oct. Sept. 10 1932

Dear Mr. Dunlap

It is a matter of deep regret that I am unable to be present today at the laying of the corner-stone of the Observatory. But I wish to say that I regard the occasion as being one of the highest importance not only for the University of Toronto but for the advancement of Science by earnest hope & strong expectation that the Observatory will be a centre from which there will go forth new wisdom trained in the most efficient manner, and that from it there will be announced from time to time discoveries and astronomical facts of primary value

From Sir Robert Falconer.

First wishes of the Corporation of the City of Toronto on the magnificent gift to the City of Toronto of the University of Toronto.

From the Mayor and City Clerk of Toronto.

The Evening Telegram, September 12, 1932

"THE OLD GUARD"



WHEN THE CORNERSTONE OF DUNLAP OBSERVATORY, University of Toronto, was laid at Richmond Hill, Ontario, there were assembled on the same platform, then Sir Howard Ferguson, High Commissioner for Canada in London, Dr. H. J. Cody, president of the University of Toronto, and Professor C. A. J. Oudemans, president of the Royal Astronomical Society. Also as undergraduates, certain distinguished young men, prominent and now there together, for they were officials of University College, Trinity and St. Michael's, and St. George's. The photographs show them in their undergraduate guise.

THE MAIL AND EMPIRE,
TORONTO, MONDAY, SEPTEMBER 12, 1932

Lay Corner-Stone of Dunlap Observatory



Prominent figures in the laying of the corner-stone on Saturday of the Dunlap Observatory, which will contain the second largest telescope in the world, are shown here. Above is Major Dunlap laying the corner-stone with a silver trowel, and on the left is Dr. H. J. Cody, president of the University of Toronto. Below is Dr. L. Bruce Macdonald, new chairman of the board of governors of the university, and on the right is Mrs. D. A. Dunlap, donor of the observatory in memory of her husband.

THE TORONTO DAILY STAR, SATURDAY, SEPTEMBER 10, 1932



LAY CORNER-STONE FOR DUNLAP OBSERVATORY

Laid the corner-stone of the Dunlap Observatory, which will contain the second largest telescope in the world, are shown here. Above is Major Dunlap laying the corner-stone with a silver trowel, and on the left is Dr. H. J. Cody, president of the University of Toronto. Below is Dr. L. Bruce Macdonald, new chairman of the board of governors of the university, and on the right is Mrs. D. A. Dunlap, donor of the observatory in memory of her husband.

TORONTO GETS BIG TELESCOPE

C. D. Gibb of Newcastle-on-Tyne, in Charge of Dunlap Memorial, Here To-day

Seventy-four-inch Instrument Slightly Larger Than One at Saanich

Toronto will shortly boast an astronomical telescope that will be slightly larger than the one established on Little Saanich Mountain, near here, it was learned with the arrival in the city today of C. D. Gibb, director of Sir Howard Grubb, Parsons & Company of Newcastle-on-Tyne, builders of the instrument.

Mr. Gibb, who is accompanied by T. J. Bell of Toronto, Canadian representative of the company, is staying at the Empress Hotel.

The Toronto University Observatory will be known as the Dunlap Memorial. The construction of this telescope has been under way for about a year. It will be built on a hill at approximately 1,000 feet above sea level.

Mr. Gibb and Bell this morning went out to Saanich to examine the 72-inch telescope which is in operation at the astronomical observatory here. As compared with the 72-inch telescope here, the Toronto observatory will be equipped with a 74-inch instrument.

At the present time the local telescope is the second largest in the world, being exceeded in size by Mount Wilson, near Pasadena, Calif., which boasts a 100-inch telescope.

A 200-inch astronomical telescope, it is understood, is being planned for southern California, but it is reported that trouble is being experienced in obtaining the mirrors.

Mrs. Ray Williams, organizing secretary of the European astronomical tour party now in the city, says the largest telescope in Great Britain is a 36-inch one at Greenwich Observatory. A 36-inch instrument has just been completed at Edinburgh to replace a 30-inch one there.

Victoria Times-B.C.

Aug. 16, 1932

TORONTO TO HAVE LARGE TELESCOPE

To Be Erected As Memorial To David A. Dunlap, By His Widow

TORONTO, Aug. 16. — Details were made public last night of the erection of the second largest telescope in the world at an estimated cost of \$500,000, as a memorial to David A. Dunlap by his widow.

Work on the administration building, one of the two chief buildings being erected as part of the David Dunlap Observatory, will start immediately, officials of the University of Toronto said. The administration building, which is a 61 feet in diameter, will house the large telescope, virtually all of which is now being constructed in England. The mirror alone weighs 5,000 pounds.

The two buildings will be located in the centre of a 177-acre plot of land a mile south of Richmond Hill and half a mile east of Yonge street in Markham township. The acreage, now in farmland, will be made into a park to be known as the David Dunlap Park.

The round building will be on a circular platform 800 feet above sea level while the administration building will be on a level five feet below and in front of it will be an area three feet lower.

Reflecting Type
The telescope, larger than the one in the Dominion Government observatory at Victoria, B. C., is of a new type, that is, in place of a lens it will use a concave mirror to collect the rays of light from a star and bring them to a focus. This mirror will have a clear aperture of 74 inches, and will be fashioned from a disk of glass 76 inches in diameter and 12 inches thick.

The telescope was ordered more than two years ago from the firm of Sir Howard Grubb, Parsons & Co., Ltd., of Newcastle-on-Tyne, England, and the construction is well advanced. It was estimated that it would be finished in three years, but the final figuring of the mirror may require a longer time.

In one of the smaller domes in the administration building it is proposed to mount the 12-inch reflecting telescope recently constructed by the University of Toronto by other small one, while in the astronomical cameras on a single mounting. The chief parts of the telescope, however, already on hand, and the mounting will be made in the factory workshop. In the central telescope it is hoped to mount a 10-inch type. The reflecting telescope, however, will be used for observation of planets, comets and will also be available for public purposes.

Kitchener Record.
Ont. Aug. 16, 1932

\$500,000 Telescope Is Widow's Gift In Memory of Husband

TORONTO, Aug. 16.—(C.P.)—Details of the erection near Richmond Hill, north of here, of the second largest telescope in the world, at an estimated cost of \$500,000, were made public here.

Officials of the University of Toronto, to which Mrs. Jessie Dunlap is donating the David Dunlap Observatory in memory of her husband, announced construction would start at once on the two main buildings of the plant.

On a circular platform 800 feet above sea level, a round building 61 feet in diameter will be built to house the huge telescope, nearly all parts of which are being made in England. Larger than the one in the Dominion Government Observatory, Victoria, B.C., the telescope will be of the reflecting type and will have mirrors weighing 5,000 pounds.

The telescope building and an administration building to be erected at a cost of \$125,000, will be located in the centre of a 177-acre plot which will be known as the David Dunlap Park.

TORONTO WILL HAVE GIGANTIC TELESCOPE

Mrs. Jessie Dunlap to Make \$500,000 Gift

TORONTO, Aug. 16.—(C.P.)—Details of the erection near Richmond Hill, north of here, of the second largest telescope in the world, at an estimated cost of \$500,000, were made public last night.

Officials of the University of Toronto, to which Mrs. Jessie Dunlap is donating the David Dunlap Observatory in memory of her husband, announced construction would start at once on the two main buildings of the plant.

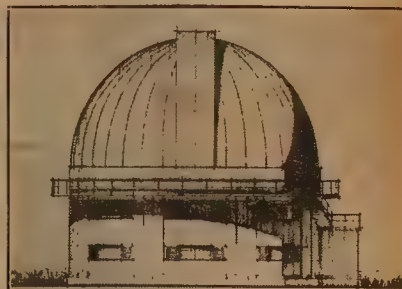
On a circular platform 800 feet above sea level, a round building 61 feet in diameter will be built to house the huge telescope, nearly all parts of which are being made in England. Larger than the one in the Dominion Government Observatory, Victoria, B.C., the telescope will be of the reflecting type and will have mirrors weighing 5,000 pounds.

The telescope building and an administration building to be erected at a cost of \$125,000, will be located in the centre of a 177-acre plot which will be known as the David Dunlap Park.

Dunlap Memorial Observatory and Administration Building



ADMINISTRATION BUILDING



GREAT TELESCOPE

Full details of the second largest observatory in the world, erection of which will begin on property a mile south of Richmond Hill immediately, were released last night. It is being provided as a memorial to the late David A. Dunlap, by his widow, Mrs. Jessie D. Dunlap. The administration building, shown in the upper sketch, will cost \$125,000. Below is shown the observatory itself, work on the telescope and other parts of which has been proceeding in England for two years and may require another year yet. When completed, the memorial and 177-acre park in which it is situated will be turned over to the University of Toronto.

Vancouver Province.
B.C. Aug. 16, 1932

Montreal Star. Que.
Aug. 16.

Second Largest Telescope Will Be At Toronto

TORONTO, Aug. 16. — Details of the erection near Richmond Hill, north of here of the second largest telescope in the world, at an estimated cost of \$500,000, were made public last night.

Officials of the University of Toronto to which Mrs. Dunlap is donating the David Dunlap Observatory in memory of her husband, announced that construction would start at once on the two main buildings of the plant.

On a circular platform 800 feet above sea level, a round building 61 feet in diameter will be built to house the huge telescope, nearly all parts of which are being made in England. Larger than the one in the Dominion Government Observatory, Victoria, B.C., the telescope will be of the reflecting type and will have mirrors weighing 5,000 pounds.

The telescope building and an administration building to be erected at a cost of \$125,000, will be located in the center of a 177 acre plot which will be known as the David Dunlap park.

WORK BEGINS AT ONCE

Second Largest Telescope in World Will Be in Toronto

TORONTO, Aug. 16.—Details of the erection near Richmond Hill, north of here, of the second largest telescope in the world, at an estimated cost of \$500,000, were made public tonight.

Officials of the University of Toronto to which Mrs. Jessie Dunlap is donating the David Dunlap Observatory in memory of her husband, announced that construction would start at once on the two main buildings of the plant.

On a circular platform 800 feet above sea level, a round building 61 feet in diameter will be built to house the huge telescope, nearly all parts of which are being made in England. Larger than the one in the Dominion Government Observatory, Victoria, B.C., the telescope will be of the reflecting type and will have mirrors weighing 5,000 pounds.

The telescope building and an administration building to be erected at a cost of \$125,000, will be located in the center of a 177 acre plot which will be known as the David Dunlap Park.

Toronto Mail and Empire Aug. 16, 1932

Montreal Gazette. Que.
Aug. 16, 1932

Ft. William Times-Journal
Ont. Aug. 16, 1932

Toronto Globe. Aug. 17, 1932

New Canadian Observatory to be Second Largest in the World



NEW DUNLAP OBSERVATORY AT RICHMOND HILL, ONTARIO, FOR UNIVERSITY OF TORONTO ASTRONOMY DEPARTMENT. Architect's drawing of the new University of Toronto Astronomy world, and at the RIGHT the will cost \$500,000 as a gift of the late David A. Dunlap, who was Dunlap Observatory to be built at Richmond Hill, Ontario, for the University of Toronto. At LEFT, the structure \$125,000 administration building, Dunlap estate, through Mrs. Jessie Dunlap, is the site of the second largest in the 177 acres. The total layout, which city of Toronto, to commemorate the plans.

Welland-Pt. Colborne Eve. Tribune. Ont. Aug. 17, 1932

Observatory Erection To Begin Immediately

Work on the administration building of the David Dunlap Observatory a mile south of Richmond Hill is to begin immediately, was announced last night. The observatory, costing \$500,000, and housing the second largest telescope in the world, is being erected as a memorial to David A. Dunlap by his widow.

The administration building, one of the two chief structures, costs about \$125,000, but the greater part of the expense is involved in the erection of a circular building 61 feet in diameter 800 feet above sea level, and the 74-inch mirror telescope that is to be installed there. The two buildings will be in the center of a 177-acre plot half a mile south of Yonge street in Markham township. The area is to be developed into a park to be known as the David Dunlap Park.

The great telescope was ordered about two years ago from England and is expected to be completed in another year's time. A 12-inch reflecting telescope, constructed by the University of Toronto at the University of the smaller domes in the administration building, while in the astronomical cameras.

Construction of Dunlap Memorial Observatory Started



CONSTRUCTION OF THE \$500,000 DAVID DUNLAP MEMORIAL OBSERVATORY, the gift of Mrs. D. A. Dunlap to the University of Toronto, is commencing on the 177-acre site near Richmond Hill. Architects' drawings shown above are of the administration building (right) and the observatory proper, which will house a 74-inch mirror telescope, the second largest in the world.

Toronto Telegram. August 16, 1932

OBSERVATORY TO BE ERECTED NEAR RICHMOND HILL



The world's second largest telescope will be erected in the new Dunlap Observatory of the University of Toronto, which will be located in Markham township just south of Richmond Hill. The above layout shows architects' drawing, at the left the structure to house the great telescope and at the right the administration building.

Construction of a building which will house the second largest telescope in the world, and an administration building which will contain laboratories, lecture room, library, and a workshop, will begin almost immediately. The total cost of the observatory, including land, buildings and apparatus, is approximately \$500,000.

The observatory will be a memorial to the late David A. Dunlap, who was a keen student of astronomy, and will be built through funds provided by his widow, Mrs. Jessie D. Dunlap, with whom is associated their son, Moffat Dunlap. Upon completion it will be presented to the University of Toronto, and will be conducted by the Department of Astronomy, of which Prof. C. A. Chant is the head. Mathews & Haldenby, architects, prepared the plans for the Administration Building and for the renovation of an old red brick farm-house on the property which will be used as living quarters for students and those in charge of the apparatus.

The buildings will lie in the midst of a 177-acre plot of land, 800 feet above sea level, and about one mile south of the Village of Richmond Hill and one-half mile east of the Yonge Street Highway. This large acreage will be developed by the Faculty of Forestry and will be known as the David Dunlap Park.

The main telescope is of the reflecting type. In place of a lens it will be equipped with a concave mirror to collect light rays from a star and bring them to focus. This mirror will have a clear aperture of seventy-four inches, and will be fashioned from a disc of glass twelve inches thick and weighing about two and one-half tons. The order for the telescope was placed two years ago, and its construction is well advanced, although it will not be completed until next summer. Special machinery was built to fashion the huge bearings of the apparatus, and to polish the glass. The centre of the immense block of glass, the grinding and the grinding and polishing of its surface to a correctness of one-half millionth of an inch, is a lengthy process, but every effort is being made to have the apparatus in working order by December, 1933.

The structure which will contain the huge telescope will be a circular building sixty-one feet in diameter, surmounted by a metal hemispherical dome, which will be revolved by an electric motor. The framework is of steel. The walls will be two feet thick, and hollow, being sheathed on the inside and outside with galvanized iron in order that temperature changes will not affect the delicate mechanism of the telescope. The inner and outer surfaces of the dome into two feet apart, will be covered with a special mixture of papier-mache and the outer surface will be covered by a layer of sheet copper for protection against the weather. The building is being constructed in England by the firm of Sir Howard Grubb, Parsons & Company, Newcastle-on-Tyne, which is also building the telescope. The structure will be shipped to Richmond Hill in about a year after it has been

set up and tested in England. This summer, however, the cement foundation of the building will carry the immense pier which will carry the telescope will be erected.

The Administration Building, constructed from Credit Valley limestone with trimmings of Queenston or Indiana stone, will be ninety-one feet long by forty-nine feet wide, with semi-octagonal projections on each end. On the roof will be a small dome at either end and a large dome at the centre. In one of the small domes will be mounted the 19-inch reflecting telescope recently constructed by Professor R. K. Young of the University of Toronto, while in the other small dome there will be three astronomical cameras on a single mounting. In the central dome a ten-inch telescope of the refracting type probably will be mounted. This is to be used for the observation of planets, comets, occultations and double stars, and will be available for public purposes. The building will be equipped with the most modern computing instruments, and the entrance hall and main staircase will be finished in marble.

The David Dunlap Observatory at Toronto is to house the second largest telescope in the world and is to be in the centre of the park of that name. The telescope will be considerably larger than the one in the Dominion Government Observatory at Victoria, B.C., and when it is in operation Canada will be well-equipped to observe the wonders of the heavens. Thanks are due to Mrs. Jessie Dunlap, who has provided the funds for the institution, which is in memory of her late husband.

New Landmark To Be Erected on Yonge Street Highway

DETAILS ANNOUNCED OF BIG OBSERVATORY AT RICHMOND HILL

Second-Largest Telescope in World Will Be Housed in Building Being Constructed and Tested in England

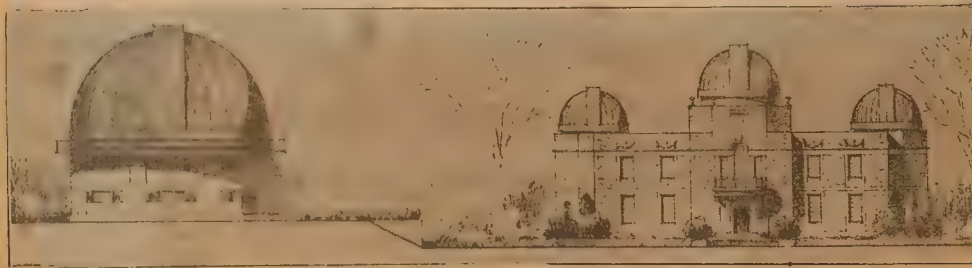
Construction of a building which will house the second largest telescope in the world, and an administration building which will contain laboratories, lecture rooms, libraries and a workshop, will begin almost immediately about one mile south of Richmond Hill. The total cost of the observatory, including land, buildings and apparatus, is approximately \$500,000.

The observatory will be a memorial to the late David A. Dunlap, who was a keen student of astronomy, and will be built through funds provided by his widow, Mrs. Jessie D. Dunlap, with whom is associated their son, Moffat Dunlap. Upon completion it will be presented to the University of Toronto and will be conducted by the Department of Astronomy, of which Prof. C. A. Chant is the

Penetanguishene
Herald. Ont.
Aug. 18, 1932

Richmond Hill Liberal. Ont. Aug. 18, 1932

NEW CANADIAN OBSERVATORY TO BE SECOND LARGEST IN THE WORLD



NEW DUNLAP OBSERVATORY AT RICHMOND HILL, ONTARIO, FOR UNIVERSITY OF TORONTO ASTRONOMY DEPARTMENT. Architects' drawing of the new which will be the second largest in the world, at the right the 177-acre site of this observatory covers. Dunlap estate, through Mrs. Jessie Dunlap, who has provided the funds for the institution, which is in memory of her late husband. The total layout, which Dunlap and her son, to the University of Toronto, is commemorated in the name of the David Dunlap Park.

Owen Sound Sun Times. Ont. Aug. 17, 1932

New Canadian Observatory To Be Second Largest In The World



NEW DUNLAP OBSERVATORY AT RICHMOND HILL, ONTARIO, FOR UNIVERSITY OF TORONTO ASTRONOMY DEPARTMENT. Architects' drawing of the new which will be the second largest in the world, at the right the 177-acre site of this observatory covers. Dunlap estate, through Mrs. Jessie Dunlap, who has provided the funds for the institution, which is in memory of her late husband. The total layout, which Dunlap and her son, to the University of Toronto, is commemorated in the name of the David Dunlap Park.

Peterborough Examiner. Ont. Aug. 18, 1932

ings of the apparatus and to polish the glass. The casting of and polishing of its surface to the immense block of glass, the annealing and the grinding and polishing of its surface to a correctness of one-half millionth of an inch, is a lengthy process, but every effort is being made to have the apparatus in working order for December, 1933.

The structure which will contain the huge telescope will be a circular building, sixty-one feet in diameter, surmounted by a metal hemispherical dome which will be revolved by an electric motor. The framework is of steel. The walls will be two feet thick, and hollow, being sheathed on the inside and outside with galvanized iron in order that temperature changes will not affect the delicate mechanism of the telescope. The inner and outer surfaces of the dome, also two feet apart, will be covered with a special mixture of papier-mache and the outer surface will be covered by a layer of sheet copper for protection against the weather. The building is being constructed in England by the firm of Sir Howard Grubb, Parsons & Company, Newcastle-on-Tyne, which is also building the telescope. The structure will be shipped to Richmond Hill in about a year after it has been set up and tested in England. This summer however, the cement foundation of the building will be laid and the immense pier which will carry the 50-ton telescope will be erected.

The Administration Building, constructed from Credit Valley limestone, with trimmings of Queenston or Indiana stone, will be ninety-one feet long by forty-nine feet wide, with semi-octagonal projections at each end. On the roof will be a small dome at either end and a large dome at the centre. In one of the small domes will be mounted the 19-inch reflecting telescope recently constructed by Professor R. K. Young of the University of Toronto, while in the other small dome there will be three astronomical cameras on a single mounting. In the central dome a ten-inch telescope of the refracting type probably will be mounted. This is to be used for the observation of planets, comets, occultations and double stars, and will be available for public head. Mathews & Haldenby, architects, prepared the plans for the administration Building and for the renovation of an old red brick farm-house on the property, which will be used as living quarters for students and those in charge of the apparatus.

The buildings will lie in the midst of a 177-acre plot of land 800 feet above sea level, and about one mile south of the village of Richmond Hill and one half mile east of the Yonge street Highway. This large acreage will be developed by the Faculty of Forestry and will be known as the David Dunlap Park.

The main telescope is of the reflecting type. In place of a lens it will be equipped with a concave mirror to collect light rays from a star and bring them to focus. This mirror will have a clear aperture of seventy-four inches, and will be fashioned from a disc of glass twelve inches thick and weighing about two and one-half tons. The order for the telescope was placed two years ago, and its construction is well advanced, although it will not be completed until next summer. Special machinery was built to fashion the huge bearings of the apparatus, and to polish the glass. The centre of the immense block of glass, the grinding and the grinding and polishing of its surface to a correctness of one-half millionth of an inch, is a lengthy process, but every effort is being made to have the apparatus in working order by December, 1933.



LE NOUVEL OBSERVATOIRE DUNLAP DE L'UNIVERSITÉ DE TORONTO — Il sera construit à Richmond Hill, Ontario. Ce sont les plans de l'architecte. A gauche on voit la bâtisse qui abritera le grand télescope mobile qui sera projeté à travers la coupole pour les observations astronomiques. A droite, on aperçoit l'édifice des bureaux d'administration. Cet observatoire a été donné à l'Université par Mme Jessie Dunlap et son fils, en mémoire de feu David Dunlap, astronome réputé.

Montreal Le Devoir. Que. Aug. 18, 1932

New Canadian Observatory to Be Second Largest in the World



Architect's drawing of the new Dunlap Observatory; at left, the structure housing a new telescope, which will be the second largest in the world, and at the right, the \$125,000 administration building. The site of this observatory covers 177 acres. The total layout, which will cost \$500,000, is a gift of the Dunlap estate, through Mrs. Jessie Dunlap and her son, to the University of Toronto, to commemorate the late David A. Dunlap, who was greatly interested in the science of the stars.

Pembroke Standard-Observer. Ont. Aug. 18, 1932



Architect's drawing of the new Dunlap Observatory; at left, building. The site of this observatory covers 177 acres. The total layout, which will cost \$500,000, is a gift of the Dunlap estate, through Mrs. Jessie Dunlap and her son, to the University of Toronto, to commemorate the late David A. Dunlap, who was greatly interested in the science of the stars.

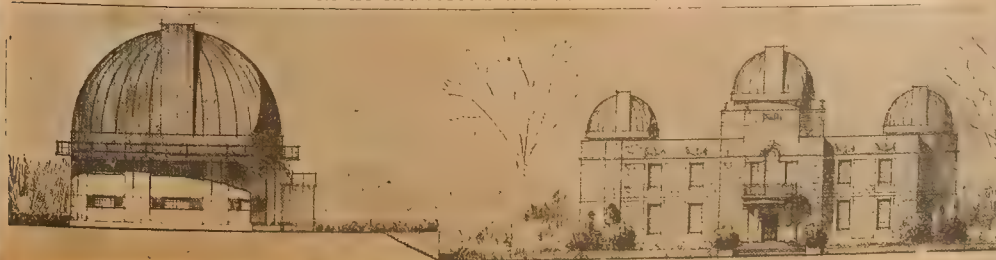
Kingston Whig-Standard. Ont. Aug. 17, 1932



NEW CANADIAN OBSERVATORY TO BE SECOND LARGEST IN THE WORLD. Architect's drawing of the new Dunlap Observatory; at LEFT, the structure housing a new telescope, which will be the second largest in the world, and at the RIGHT the \$125,000 administration building. The site of this observatory covers 177 acres. The total layout, which will cost \$500,000, is a gift of the Dunlap estate, through Mrs. Jessie Dunlap and her son, to the University of Toronto, to commemorate the late David A. Dunlap, who was greatly interested in the science of the stars.

Woodstock Sentinel-Review. Ont. Aug. 17, 1932

NEW CANADIAN OBSERVATORY AT RICHMOND HILL, ONTARIO, TO BE SECOND LARGEST IN WORLD.

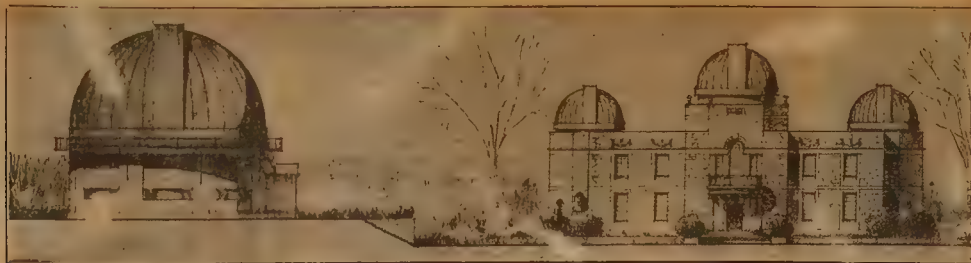


St. Thomas Times-Journal.

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Ont. August 18, 1932

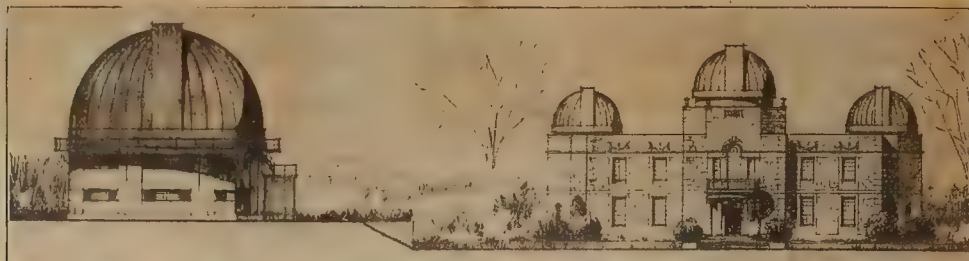
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NEW DUNLAP OBSERVATORY AT RICHMOND HILL, ONTARIO, FOR UNIVERSITY OF TORONTO ASTRONOMY DEPARTMENT
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St. Catharines Standard. Ont. Aug. 19.

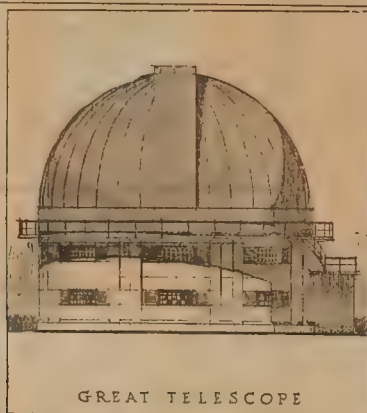
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Salt Reporter. Ont. Aug. 19.

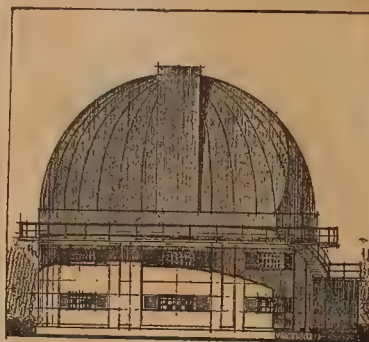
Work To Start On Giant Observatory



GREAT TELESCOPE

The second largest telescope in the world will be housed at the new Dunlap Observatory at Richmond Hill, Ontario. The observatory will house a large telescope the mirror of which alone weighs 5000 pounds. In addition there will be a large administration building which will house the observatory's offices and other necessary buildings. It is estimated that the total cost will be in the neighborhood of \$500,000. Above is a drawing of the great observatory by the architect, Mathers and Haldenby.

Un immense observatoire



Le deuxième observatoire en importance au monde est sur le point d'être commencé à Richmond Hill, Ontario. Il sera élevé à la mémoire de feu David A. Dunlap par son épouse Mme Jessie D. Dunlap. Le coût de la construction sera d'environ \$500,000. L'observatoire comprendra un immense télescope dont le miroir pèsera à lui seul 5,000 livres. Dans les bureaux de l'administration il y aura deux télescopes plus petits et deux caméras astronomiques. Voici les plans de la bâtisse tels que préparés par les architectes Mathers et Haldenby.

Kitchener Record. Ont. Aug. 19

Three Rivers Le Nouvelliste. Que. August 19



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Sudbury Star. Ont. Aug. 20.

LARGE TELESCOPE COMING TO CANADA

Will Be Housed in Memorial
Observatory Near
Toronto.

Toronto, Aug. 20.—There will be completed next year at Newcastle-on-Tyne, England, the second largest telescope in the world. It will be shipped to Canada and erected in a new observatory located on high ground a few miles north of Toronto. The observatory is being built as a memorial to the late David A. Dunlap, who in his lifetime was a keen student of astronomy, funds being provided by his widow and son. Upon completion, it will be presented to the University of Toronto and will be conducted by the Department of Astronomy of the institution. The building comprising the observatory will stand in the midst of a 177-acre plot of land, 800 feet above sea level, and the property will be beautified by the Faculty of Forestry of the University and will be known as the David Dunlap Park.

The main telescope is of the reflecting type. In place of a lens it will be equipped with a concave mirror to collect light rays from a star and bring them to focus. This mirror will have a clear aperture of seventy-four inches, and will be fashioned from a disc of glass twelve inches thick and weighing about two and one-half tons. The order for the telescope was placed two years ago, and its construction is well advanced, although it will not be completed until next summer. Special machinery was built to fashion the huge bearings of the apparatus and to polish the glass. The casting of the immense block of glass, too annealing and grinding and polishing of its surface to a correctness of one-half millionth of an inch, is a lengthy process, but every effort is being made to have the apparatus in working order by December, 1933.

The structure which will contain the huge telescope will be a circular building, sixty-one feet in diameter, surmounted by a metal hemispherical dome, which will be revolved by an electric motor. The framework is of steel. The walls will be two feet thick, and hollow, being sheathed on the inside and outside with galvanized iron in order that temperature changes will not affect the delicate mechanism of the telescope. The inner and outer surfaces of the dome, also two feet apart, will be covered with a special mixture of paper-mache and the outer surface will be covered by a layer of sheet copper for protection against the weather.

*Fredericton
Gleaner. N.B.
Aug. 22.*

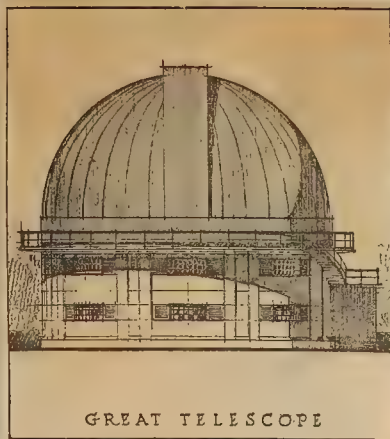
Big Observatory Planned

TORONTO, Aug. 20. — Approximately \$500,000 will be spent on the construction of a new observatory near Richmond Hill, north of Toronto. The station will be known as the David Dunlap Observatory and is being donated to the University of Toronto by Mrs. Jessie Dunlap. It will contain the second largest telescope in the world, larger by two feet than the one in the Dominion Observatory in Victoria. Telescope building and administration building will cost approximately \$125,000.

*Vancouver Journal
of Commerce. B.C.
Aug. 22.*

One more public benefaction attributable to the mining industry of Northern Ontario is the telescope to be erected at the David Dunlap Observatory at Toronto. The telescope will be the second largest in the world, considerably larger than the one in the Dominion Government Observatory at Victoria, B.C., and when it is in operation Canada will be well equipped to observe the wonders of the heavens. Mrs. Jessie Dunlap is providing the funds for the institution, which is in memory of her late husband, who was interested, with the Timmins brothers syndicate, in the LaRoe and Hollinger mines.

*Sudbury Star. Ont.
Aug. 20.*

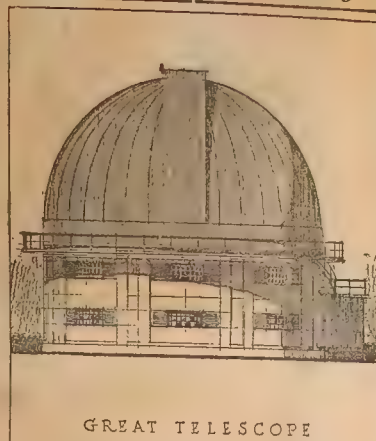


GREAT TELESCOPE

WORK TO COMMENCE ON NEW GIANT OBSERVATORY—

The second largest observatory in the world is about to be erected near Richmond Hill, Ontario, as a memorial to the late David A. Dunlap, by his widow, Mrs. Jessie D. Dunlap. This observatory will house a huge telescope, the mirror of which alone weighs 5,000 pounds. In addition there will be an administration building which will contain two smaller telescopes and three astronomical cameras. It is estimated that the total cost will be in the neighborhood of \$500,000. Above is a drawing of the great observatory by the architects, Mathers & Haldenby.

New Giant Observatory



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Hamilton Spectator. Ont. Aug. 22

Fredericton Gleaner. N.B. Aug. 22.

NEW CANADIAN OBSERVATORY TO BE SECOND LARGEST IN THE WORLD



Architect's drawing of the new and largest in the world, and at the Dunlap Observatory; at LEFT, the structure housing a new telescope, which will be the second largest in the world, and at the RIGHT the \$125,000 administration building. The site of this observatory covers 177 acres. The total layout, which will cost \$500,000, is a gift of the Dunlap estate, through Mrs. Jessie Dunlap and her son to the University of Toronto, to commemorate the late David A. Dunlap, who was greatly interested in the science of the stars.

Gloucester Bay Gazette. C.B. Aug. 20.

Observatory to Be Second Largest in World



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St. John Telegraph-Journal. N.B. Aug. 20.

HALF MILLION OBSERVATORY IS PLANNED

TORONTO, Aug. 18. Details of the erection of a new observatory at Richmond Hill, north of Toronto, of the second largest telescope in the world at an estimated cost of \$500,000, are made public here.

Officials of the University of Toronto, to which Mrs. Jessie Dunlap is donating the David Dunlap Observatory in memory of her husband, announced construction would start at once on the two main buildings of the plant.

On a circular platform 800 feet above sea level, a round building 61 feet in diameter will be built to house the huge telescope, nearly all parts of which are being made in England. Larger than the one in the Dominion Government observatory, Victoria, B.C., the telescope will be on the reflecting type and will have mirrors weighing 5,000 pounds.

The telescope building and an administration building to be erected at a cost of \$125,000, will be located in the centre of a 177-acre plot, which will be known as the David Dunlap Park.

New Westminster British Columbian B.C. Aug. 23

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Rapid City Reporter. Man. Aug. 24.

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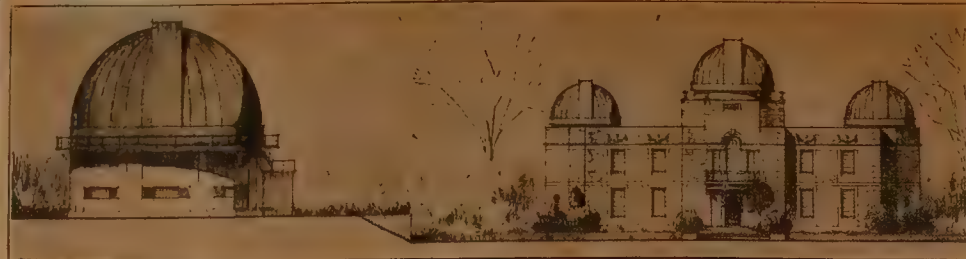
Blairmore Enterprise Alta. Aug. 25

Orillia News-Letter. Ont. Aug. 24

Eaton's Enterprise. Sask. Aug. 25

Eston Press Sask. Aug. 25.

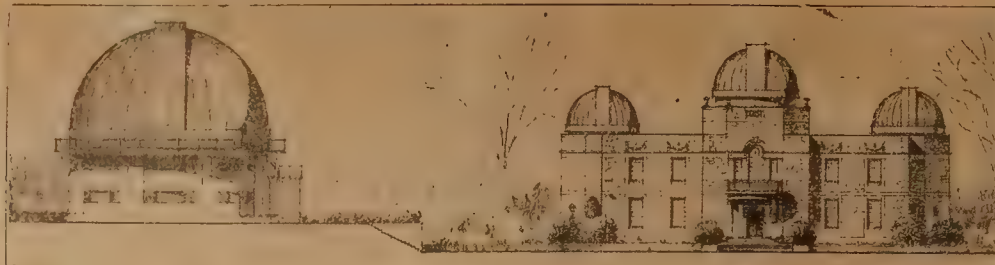
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Edmonton Journal Alta. Aug. 23

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New Glasgow Chronicle N.S. Aug. 23

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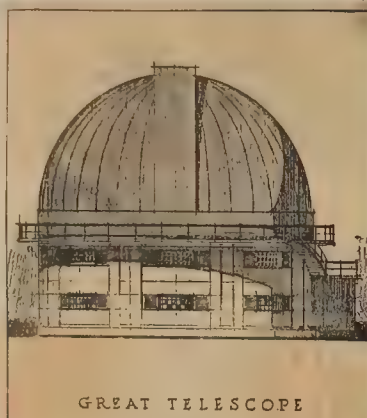
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Niagara Falls Review Ont. Aug. 24.

Canora Courier. Sask. Aug. 25

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Indian Head News.
Sask. Aug. 25.

Kinistino Representative
Sask. Aug. 30

La Fleche Press
Sask. Aug. 25

Luseland Dispatch
Sask. Aug. 25

Rouleau Enterprise
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NEW CANADIAN OBSERVATORY TO BE SECOND LARGEST IN THE WORLD

Architect's drawing of the new Dunlap Observatory; at left, the structure housing a new telescope, which will be the second largest in the world and at the right the \$125,000 administration building. The site of this observatory covers 177 acres. The total layout, which will cost \$500,000, is a gift of the Dunlap estate, through Mrs. Jessie Dunlap and her son, to the University of Toronto, to commemorate the late David A. Dunlap, who was greatly interested in the science of the stars.

Calgary Herald. Alta. Aug. 25

Shaunavon Standard
Sask. Aug. 25

Build Huge Telescope

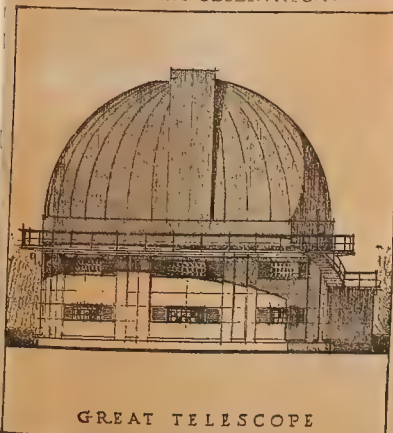
Second Largest Telescope In the World To Be Erected Near Toronto

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Officials of the University of Toronto to which Mrs. Jessie Dunlap is donating the David Dunlap Observatory in memory of her husband, announced construction would start at once on the two main buildings of the plant.

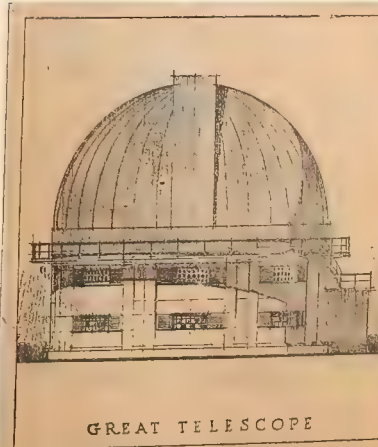
On a circular platform 800 feet above sea level, a round building 61 feet in diameter will be built to house the huge telescope, nearly all parts of which are being made in England. Larger than the one in the Dominion Government observatory, Victoria, B.C., the telescope will be on the reflecting type and will have mirrors weighing 5,000 pounds.

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NEW GIANT OBSERVATORY

GREAT TELESCOPE

The second largest observatory in the world is about to be erected near Richmond Hill, Ontario, as a memorial to the late David A. Dunlap, by his widow Mrs. Jessie D. Dunlap. This observatory will house a huge telescope, the mirror of which alone weighs 5,000 pounds. In addition there will be an administration building which will contain two smaller telescopes and three astronomical cameras. It is estimated that the total cost will be in the neighborhood of \$500,000. Above is a drawing of the great observatory by the architects, Mathers and Haldenby.



GREAT TELESCOPE

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Alameda Dispatch
Sask. Aug. 26

Calgary Albertan. Alta.
Aug. 26

Medicine Hat News. Alta. Aug. 29

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Toronto Will Have Gigantic Telescope

MRS. JESSIE DUNLAP TO MAKE \$500,000 Gift

TORONTO, Aug. 16.—(C.P.)—Details of the erection near Richmond Hill, north of here, of the second largest telescope in the world, at an estimated cost of \$500,000, were made public last night.

Officials of the University of Toronto to which Mrs. Jessie Dunlap is donating the David Dunlap Observatory in memory of her husband, announced that construction would start at once on the two main buildings of the plant.

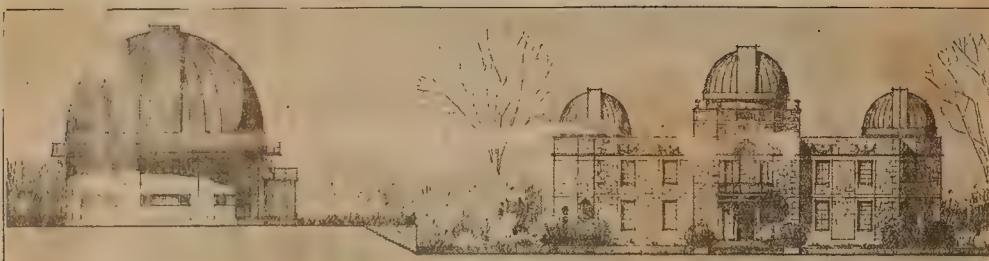
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Moosomin World-Spectator. Sask. Aug. 31

Elrose Review Sask. Sept. 1

St. John's Telegram Nfld. Sept. 1

New Canadian Observatory To Be Second Largest in the World

Architectural plans of the new Dunlap observatory showing at left the structure to house a new telescope, which will be the second largest in the world. At the right is shown the \$125,000 administration building. The site of the observatory covers 177 acres. The total layout which will cost \$500,000 is a gift of the Dunlap estate, through Mrs. Jessie Dunlap, to the University of Toronto to commemorate the late David A. Dunlap, who was greatly interested in the science of the stars.

Regina Leader-Post. Sask. Sept. 2

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Sashatchewan Valley News. Rosthern, Sask. Sept. 7

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The Viking News Alta. Sept. 7

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Pincher Creek Echo Sept. 8

RICHMOND HILL
Will Lay Cornerstone
Observatory Sept. 10
Richmond Hill, Aug. 30.—Announcement that the corner stone of the administration building of the Dunlap Observatory at Richmond Hill will be laid on Saturday, Sept. 10, was made by Dr. H. J. Cody, president of the University of Toronto. Details of the ceremony will be announced later.

Toronto Telegram
Aug. 30.

LAY STONE SEPT. 10
The president of the University of Toronto, Dr. H. J. Cody, announces that the corner-stone of the administration building of the Dunlap Observatory at Richmond Hill will be laid on Saturday afternoon, Sept. 10, at 3 o'clock by Moffat Dunlap. Details will be announced later.

Toronto Daily Star
Aug. 30

Lay Corner-stone Saturday
of Dunlap Observatory

The president of the University of Toronto, Dr. H. J. Cody, announces that the corner-stone of the Administration Building of the Dunlap Observatory at Richmond Hill will be laid on Saturday afternoon, Sept. 10, at three o'clock by Moffat Dunlap. Details of the ceremony will be announced later.

Toronto Mail and
Empire. Ont. Aug. 30

The huge reflecting telescope now under construction for the Dunlap Observatory, University of Toronto, will be the second largest in the world.

St. John Times-
Globe. N.B. Aug. 31

Announcement that the corner stone of the administration building of the Dunlap Observatory at Richmond Hill will be laid on Saturday, Sept. 10, was made by Dr. H. J. Cody, president of the University of Toronto. The cornerstone will be laid by Moffat Dunlap, who with his father, Mrs. D. A. Dunlap, gave the funds for the building and equipment.

Tottenham Sentinel
Ont. Sept. 1, 1932

The huge reflecting telescope now under construction for the Dunlap Observatory, University of Toronto, will be the second largest in the world.

Montreal Gazette
Que. Sept. 2, 1932

The mirror, for the huge reflecting telescope now being built for the Dunlap Observatory, University of Toronto, will weigh 5,000 pounds.

Owen Sound Sun
Times. Ont. Sept. 3, 1932

The mirror, for the huge reflecting telescope now being built for the Dunlap Observatory, University of Toronto, will weigh 5,000 pounds.

Brandon Sun. Man.
Sept. 5, 1932

Markham Twp., Ont.
The following contracts have been awarded in connection with the erection of concrete, steel and stone administration building at David Dunlap Observatory: elec. Standard Electric Contracting and Engineering Co., 195 Victoria St., Toronto; street steel, Dominion Bridge Co. Ltd., 1139 Shaw St., Toronto; reinf. steel, Truscon Steel Co. of Canada Ltd., Federal Bldg., Toronto; cut stone, Nicholson and Curtis, Leaside; light proof window shade, Higgin Mfg. Co., 33 McCaul St., Toronto; gravel, J. E. Montgomery, 95 Butterwood Ave., Mount Dennis. Owner, University of Toronto, archts., Mathers and Haldenby, 96 Bloor St. W., Toronto; gen. contr., Sullivan and Fried Ltd., 81 Victoria St., Toronto

Contract Record and
Engineering Review Sept. 7

Will Lay Corner Stone.
Corner stone of the administration building of the Dunlap Observatory on Richmond Hill will be laid on Saturday at 3 p.m. by Moffat Dunlap. Dr. Cody, as president of the University of Toronto, will preside. Prof. C. A. Chant will describe the proposed building; Rev. Dr. Richard Roberts will offer prayer, and addresses will be given by Sir Wm. Mulock, Premier Henry and Hon. G. Howard Ferguson.

Toronto Telegram
Sept. 7, 1932

The mirror, for the huge reflecting telescope now being built for the David Dunlap Observatory, University of Toronto, will weigh 5,000 pounds.

St. Catharines
Standard Ont. Sept. 8, 1932

President of University
to Open New Observatory

Under the direction of Rev. Canon H. J. Cody, president of the University of Toronto, the cornerstones of the new Dunlap Observatory, a gift to Toronto by Mrs. D. A. Dunlap, will be laid to-morrow afternoon at 3 o'clock. Moffat Dunlap will lay the stone after presentation of the trowel by Canon Cody. Before the laying of the stone, Prof. C. A. Chant will describe the proposed building and its purpose. Following the ceremony, addresses will be delivered by Sir William Mulock, Premier Henry and Hon. G. Howard Ferguson. The site of the observatory is south of Richmond Hill on the east side of Yonge Street.

Toronto Mail and
Empire. Ont. Sept. 9

EDUCATIONAL and parliamentary dignitaries will gather at Richmond Hill, Ont., tomorrow at the ceremonies attending the laying of the corner-stone of the administration building of the David Dunlap Observatory of the University of Toronto. Canon H. J. Cody, president of the university, will be in charge of the ceremonies.

Sault Ste. Marie Star
Ont. Sept. 9, 1932

TO LAY CORNERSTONE
(Canadian Press Dispatch)
RICHMOND HILL, Ont., Sept. 9.—Educational and parliamentary dignitaries will gather here tomorrow at the ceremonies attending the laying of the cornerstone of the administration building of the David Dunlap Observatory of the University of Toronto. Canon H. J. Cody, president of the University, will be in charge of the ceremonies.

Owen Sound Sun
Times. Ont. Sept. 10, 1932

LAY CORNERSTONE OF OBSERVATORY

Provincial Officials Pay Tribute
To Mrs. Dunlap For Princely Benefaction

TORONTO, Sept. 12.—The foundation stone of the Dunlap Observatory at Richmond Hill was laid by Moffat Dunlap on Saturday afternoon in memory of his father D. A. Dunlap and the deed of 177 acres surrounding it was handed to the University of Toronto by the donor, Mrs. Dunlap. In the presence of the premier of the province, the high commissioner to Great Britain and distinguished astronomers, professors and business men the ceremonies took place.

"This is a red letter day in the history of the university," said Dr. H. J. Cody, who presided at the ceremony, and declared that although the University of Toronto was a provincial institution, its benefactors had been numerous. He pointed out that all the investments of the university were gilded, the sums were invested in Dominion, provincial or high-class municipal securities, and inferred that in a time of wavering stocks and money markets, the university stood sound and secure financially. No money was invested in mortgages, none in industrial stocks.

Expressing the gratitude of the university to Mrs. Dunlap for her princely benefaction, he commented that her husband the late David Dunlap, had used every effort to popularize interest in astronomy. The situation of the new observatory was ideal, 500 feet above the sea.

Interested in Idea.
Professor A. C. Chant told of Mr. Dunlap's interest in the project. Following a lecture which he had given on astronomy and in which he had stated the need for a great observatory, Mr. Dunlap had come to him to bespeak his interest in the idea.

Hon. G. Howard Ferguson, Canadian High Commissioner, commented: "I am greatly impressed with the magnitude and extent of this gift through which Mrs. Dunlap is erecting a memorial to her husband. It must be a proud moment for her to see her son laying the corner-stone."

Announcement was made that England's Astronomer Royal, Sir Frank Dyson, as well as Eddington and Jeans and others in the galaxy of astronomers, were expected to be present at the opening of the new observatory.

The gratitude of the province was voiced by Premier Henry and a word of praise by Chancellor Sir William Mulock. Dr. D. Bruce Macdonald received the gift from the hands of Mrs. Dunlap on behalf of the university. Later Mrs. Dunlap entertained the guests at tea.

Kitchener Record. Ont. Sept. 12

Observatory Stone Is Laid

DUNLAP MEMORIAL.

Ferguson, Henry, Cody Are
Heard at Ceremony

TORONTO, Sept. 12. In the presence of the premier of the province, the high commissioner to Great Britain, and distinguished astronomers, professors and business men, the foundation stone of the Dunlap Observatory at Richmond Hill was laid by Moffat Dunlap on Saturday afternoon in memory of his father D. A. Dunlap and the deed of 177 acres surrounding it was handed over to the University of Toronto by the donor.

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Toronto Observatory
An observatory that will house the second largest telescope in the world is to be constructed near Richmond Hill, Toronto, district, and work on the administration building is to commence immediately, according to officials of the University of Toronto. The huge telescope, now under construction in England, is expected to be finished within the next year and the observatory, when fully completed, will be turned over to the University of Toronto. The observatory, which is in the nature of a memorial to the late David A. Dunlap, will cost \$200,000. The telescope, which will be 74 inches in diameter, will be mounted on a single mounting.

The two main buildings of the observatory will be erected in the centre of a 177-acre plot, one foot south of the village of Richmond Hill. The whole area will be developed by the faculty of forestry, University of Toronto, and will be known as the "David Dunlap Park."

The administration building, which plans were prepared by Mathers & Haldenby, architects, will contain laboratories, lecture rooms, libraries and a workshop. A circular building 61 feet in diameter will house the huge telescope, the mirror of which alone weighs 5,000 pounds.

The telescope was ordered over two years ago from the firm of Sir Howard Grubb, Parsons and Company, Newcastle-on-Tyne, England, and work on it is expected to be completed within a year. It is of the reflecting type and the convex mirror which is used instead of a lens in this kind of telescope will have a clear aperture of 74 inches, glass 76 inches in diameter and 12 inches thick.

Three domes will grace the upper structure of the administration building, and in each of two of these a smaller dome will probably be mounted which will be complementary to the great main dome in the separate building. The remaining dome will house three astronomical cameras on a single mounting.

Edmonton Journal
Alta. Sept. 12, 1932

Toronto University's
Observatory

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Funds for the project are being provided by Mrs. D. A. Dunlap, widow of the late David A. Dunlap, with whom is associated the Moffat Dunlap. The completed observatory will be presented to the University of Toronto and will be conducted by the department of astronomy of which Professor C. A. Chant is head.

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Border Cities Star
Windsor, Ont. Sept. 12, 1932

Winnipeg Free Press
Man. Sept. 13, 1932

LAY CORNERSTONE

RICHMOND HILL, Ont., Sept. 2. —(C.P.) Educational and public-spirited donors will gather here tomorrow at the ceremonies attending the laying of the corner stone of the administration building of the David Dunlap Observatory of the University of Toronto. Canon H. J. Cody, president of the university, will be in charge of the ceremony.

Brantford Expositor Ont.

Sept. 13, 1932

LOCAL UNIVERSITY FUNDS
TRIPLY SAFEGUARDED.

The appearance of the Manitoba University trust funds was a double present to Dr. Cody's mind when he resurveyed his audience at the dedication of the Dunlap Memorial that the trust funds of the University of Toronto are absolutely intact. For though the local seat of learning is a state institution it has attracted benefactions amounting to between four million and five million dollars. This large endowment has come in the main from the late Mr. E. C. Whittier, the Macays, the Dunlaps and the Rockefeller foundation.

It is reassuring to learn that these trust funds are wholly invested in gilt-edged issues—mainly in Dominion of Canada and Province of Ontario bonds, and in the higher grade debentures of Eastern Canadian municipalities. The University of Toronto has no mortgages and no corporation bonds at all. The investments have been made with the greatest care, and the securities thus purchased are protected behind a triple wall of caution.

None of the securities is in the hands of University officers. They are all in the custody of the Canadian Bank of Commerce, in whose deposit vaults they rest. The securities and the revenue produced by them are subject to continuous auditing by Messrs. Clarkson, Gordon and Bellworthy. In addition to this there is a periodic checking of the securities by both the bank and the auditor. In the last analysis the bank is responsible for their safety, and if such a misfortune occurred as the disappearance of one of the bonds, the bank would have to make good the loss.

We think it worth while to publish these facts in view of reassuring the public as to the extra-ordinary precautions taken to protect the trust funds belonging to the state university.

Toronto Mail and Empire

Sept. 14, 1932

President H. J. Cody of the University of Toronto delivered the opening address on Saturday when the corner stone of the administration building, David Dunlap Observatory, at Richmond Hill, was laid by David Moffatt Dunlap, only son of Mrs. Dunlap and the late David Dunlap.

Brantford Witness.

Ont. Sept. 14, 1932

Dunlap Observatory Corner-Stone Is Laid



With impressive ceremony, the corner-stone of the Dunlap Observatory was laid by D. M. Dunlap at Richmond Hill Saturday. The observatory is in memory of the late D. A. Dunlap, who took a deep interest in astronomy, and is the gift of Mrs. Dunlap to the University of Toronto. (1) Dr. D. Bruce MacDonald, chairman of the board of governors of the university, (2) D. M. Dunlap, son of the late D. A. Dunlap, laying the cornerstone with a silver trowel.

Peterborough Examiner. Ont. Sept. 15, 1932

Huge Telescope For
University Of Toronto

Second Largest In World Will Cost About \$500,000

Details of the erection near Richmond Hill, north Toronto, of the second largest telescope in the world, at an estimated cost of \$500,000, have been made public.

Officials of the University of Toronto to which Mrs. Jessie Dunlap is donating the David Dunlap Observatory in memory of her husband, announced that construction would start at once on the two main buildings of the plant.

On a circular platform, 800 feet above sea level, a round building 61 feet in diameter will be built to house the huge telescope, nearly all parts of which are being made in England. Larger than the one in the Dominion Government Observatory, Victoria, B.C., the telescope will be of the reflecting type and will have a mirror weighing 5,000 pounds.

The telescope building and an administration building to be erected at a cost of \$125,000, will be located in the centre of a 177-acre plot which will be known as the David Dunlap Park.

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Empress Express. Alta.

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Wetaskiwin Times Alta.

Sept. 22, 1932

Huge Telescope For
University Of Toronto

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DUNLAP OBSERVATORY CORNER-STONE IS LAID



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Sarnia Observer. Ont. Sept. 14, 1932

Qu'Appelle Progress. Sask.

Sept. 22, 1932

Wapella Post. Sask.

Sept. 15, 1932



Photographs of the Foundation, taken on September 12, 1932 (Facing North West)

Air Photograph, looking North-east Oct. 2, 1932

C.N.R.
Narrow road
(North boundary)



Old lane
2nd. Concession
of Markham Tp.
or Bayview Ave.
(East Boundary)
Farm house and
orchard.

Lane from
Yonge St.
C.N.R. Track
(West boundary of property)

2nd. concession
(East boundary)

Power line
during construction

C.N.R.



South boundary

Lane from Yonge St.

The white strip
is ploughed land

C.N.R. Track
(West boundary)

Narrow Road (North boundary)

Air Photograph, looking East. October 2nd., 1932

East boundary →
of property



Air Photograph, looking East, Oct. 2, 1932. Photographs by C.A. Chant
By Airways Limited.



C.A.C. beside the Corner
Stone looking E. Sept. 30, 1932



General view, Corner Stone in front at
middle - looking N. Sept. 30, 1932



The local office of Sullivan & Fried Gen-
eral Contractors. looking S.E. Sept. 30, 1932



Unloading cut Queenston Limestone;
used for trimming. This stone is for the
course just above cement foundation. looking N.
Sept. 30, 1932



The Front of the Administration Building—note cut stone trimmings, looking N. Oct. 7, 1932

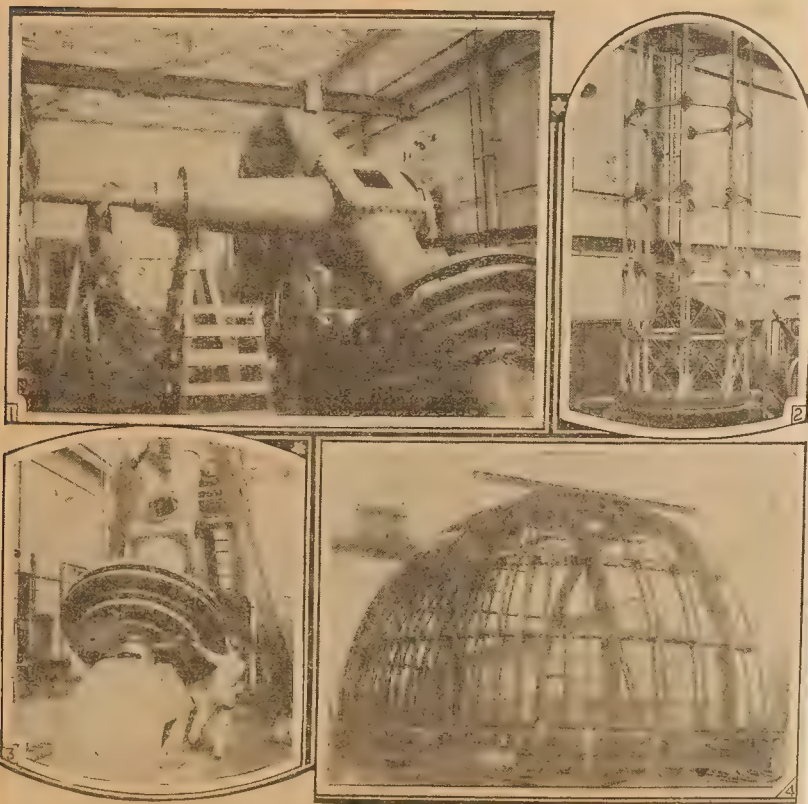


Looking N.E. note the Corner Stone behind derrick. Oct. 7, 1932



Drilling to determine the nature of the earth at the site for the Great Dome, looking S.E. Oct. 7, 1932

ASSEMBLE TELESCOPE FOR GREAT DUNLAP OBSERVATORY



Construction proceeds apace in England on the gigantic telescope for the David Dunlap Observatory, where the University of Toronto will study the stars, near Richmond Hill. The scenes above were photographed in the shops of the Sir Howard Grubb, Parsons & Co. at Newcastle-on-Tyne, which is building the equipment. In No. 1 is shown the massive mounting for the instrument. The horizontal shaft is the "declination axis," permitting observation from the horizon to the zenith. The observing tube will be attached at its right end, with a counterweight at the left. The inflex tube itself is shown in No. 2. Thirty feet high, it will contain the seventy-four-inch reflecting mirror in a casting fitted to its base. In No. 3 the

telescope is shown in its position, ready to be sent to the site. No. 4 shows the huge steel dome, twice the size of the largest bandstand at the Exhibition Grounds. The crossbar at the top carries trucks for the shutter. Below are wheels for rolling the dome to face any part of the sky. Size of the parts is brought out by figures of engineers in the pictures. When all the pieces have been tried in place, the dome and telescope will be dismantled a gain, and sent to Toronto part by part. After the telescope is erected citizens will have a weekly opportunity of taking a peep through it, according to the plan of Dr. C. A. Chant, Professor of Astrophysics, announced yesterday.

Huge Telescope Will Be Built for Toronto U.

TORONTO, Oct. 8. Best known to the world of science and to hymen as the home of insulin, the University of Toronto will shortly have new claims to fame in the field of astronomy. Work has started on a new observatory in the Richmond Hill section of this city which, when finished next year, will house one of the largest telescopes in the world. The huge, sixty-one inch instrument to be finished in England within the year will be the centre of the "David Dunlap observatory" costing over \$500,000 when completed.

The new telescope will be of the reflecting type, with a clear aperture of seventy-four inches. The disk of glass, seventy-six inches in diameter and a foot thick, weighs 5,000 pounds.

*The Northern Mail.
The Pas. Man. Oct. 8, 1932*

Telescope Project Nears Completion

In conjunction with the gift of the second largest telescope in the world, donated to this university, three smaller telescopes are to be installed, Dr. Chant, head of the astronomy department informed *The Varsity*. The Dunlop memorial telescope is to be housed in a separate observatory and the three smaller ones in a single building with three domes. This building will also serve as an administrative building and will contain the observatory offices.

Completion of the project is expected within the year. The architects drawings exhibit buildings of beautiful design, a notable addition to the university property. Completion of the project is expected within the year. Following the customary procedure, the observatories will be open for inspection to the public for one evening.

Dr. Chant has promised more detailed information at a latter date including technical details and sketches.

This group of four telescopes will assist immeasurably in the work of the department of astronomy and will lend a distinction to the University of Toronto accorded to none other in the world.

Toronto Globe . Out. Oct. 5, 1932

*The Varsity Undergraduate
Newspaper Toronto Oct. 7, 1932*



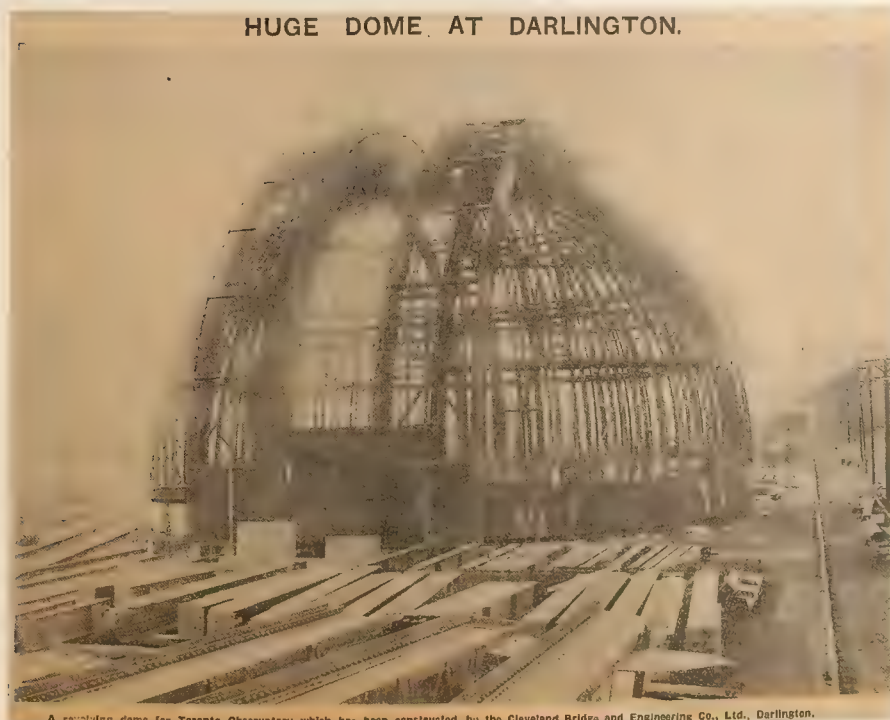
Toronto 61-ft. Dome. View taken during construction at Darlington, June 1932



Two views during the construction of the 61-foot Dome at Darlington, Eng. The steel girders in front are for the floor of the building. August 1932



*Circular Steel Building erected at Darlington before being shipped to
Toronto April - 1933*



From the Darlington "Northern Despatch" August 29, 1932



Looking E. showing Front of Building
Front Door behind derrick - Oct. 14, 1932



Looking N.W. The shanty at left front contains the steam hoist. Oct. 14, 1932



Looking S.E. showing Front and part of N. end. Oct. 14, 1932



Looking N.E., Corner Stone is $1\frac{3}{8}$ in from left edge. Oct. 14, 1932



C.A.C. in S.W. Corner office
Oct. 14, 1932



The Front Door - Oct. 14



The Front Door - Oct. 21, 1932



Looking S.E. showing N.W. corner
Oct. 21, 1932



Front and S. end of Building-
Oct. 21, 1932



S.W. Corner-corner stone
is dark.- Oct. 21, 1932



Front of Building, looking N.E.
Oct. 28, 1932



North end and Front of Building
Oct. 28, 1932



Looking N.W. toward Richmond Hill
from Building Oct. 21, 1932



Looking S.W. from second floor of Building.
Oct. 21, 1932



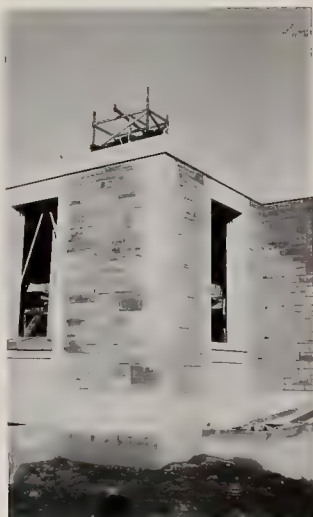
South End & part of rear. Oct. 28, 1932



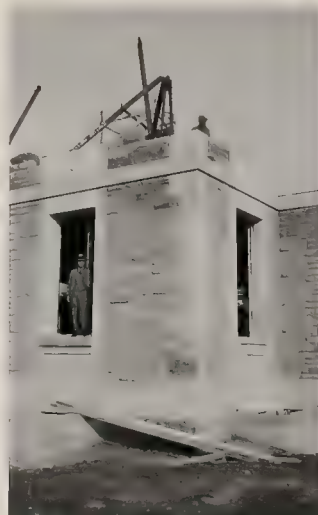
Preparing to pour upper floor Oct. 28, 1932



Stone-cutters Wilson (left) and Boyle. Oct. 28



S.W. corner - note corner stone. Oct. 28.



S.W. Corner - note upper portion. Nov. 4



Northwest portion of Building Nov. 4, 1932



Northerly part of Front Nov. 4, 1932



Gear for Operating the Shutter of the 61-foot Dome (in shop) the hand-wheel can be used if the electric power is off - September, 1932



*Gear for Operating the shutters, in place in Dome (at Darlington) Oct. 1932
View from Motor end.*



No 850

Reducing Gear

Motor

Gear for Operating the Wind-screen of the Dome (in shop) Sept. 1932



No 854

Gear for Operating the Wind-screen, showing Spindle in place in Dome Oct. , 1932



858
No 858

Gear for focussing the Cassegrain Mirror - The mirror is moved back or forth by a motor controlled from the bottom of the lattice tube - Sep. 1932

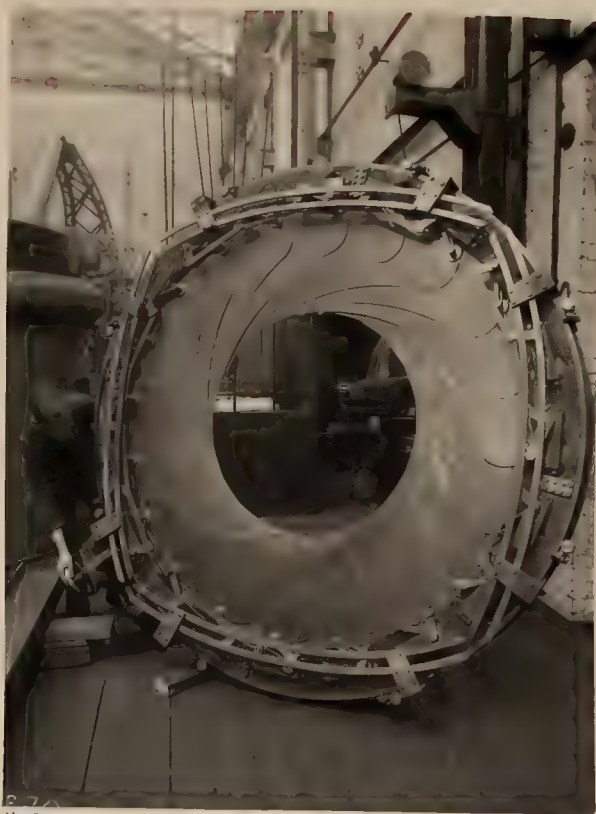


No 861

View of Shutters of the great 61-foot dome, from inside the dome, during construction. October 1932



N-801



No. 870.



No. 871

Three views of the iris Diaphragm, (which is within the centre-piece of the tube of the telescope.

(see page 86.)

869. The diaphragm wide open

870. Half closed.

871. At minimum aperture.

Nov. 1932



No 868.

View in the shop at Newcastle. The lattice tube is in the right foreground. In the middle the polar axis is seen mounted, with the declination axis inserted & with the declination counterweight in place. The centre-piece is on the floor to the right and in front of it is.... the iris diaphragm - In the background is the Greenwich 36-inch reflector, and in the foreground is the mounting for the twin 16-inch astrographic refractors, for Leyden Observatory Nov. 32



Looking N.W. Hoisting-engine house at night. Nov. 4, 1932



The upper floor, looking N. some of the floor poured; preparing another portion (at left) Nov. 4, 1932



The men mixing the cement at rear of building - looking S.E. Nov 4, 1932



*Facing S.W. Showing Front and North
End - Nov. 12, 1932*



*Facing N.W. Showing Rear and South
End - Nov. 12, 1932*



*Facing S.W. Showing Rear and North
End - Nov. 12, 1932*



*Building the inner Wall of the Second
Story - looking N. Nov. 12, 1932*



*South-west Corner and
Front. Nov. 12, 1932*



Front Door - Nov. 12, 1932



*The Front Door - Nov. 22, 1932
(Mrs. Chant in door)*



*The S.W. Corner and
the Front - Nov. 22, 1932*



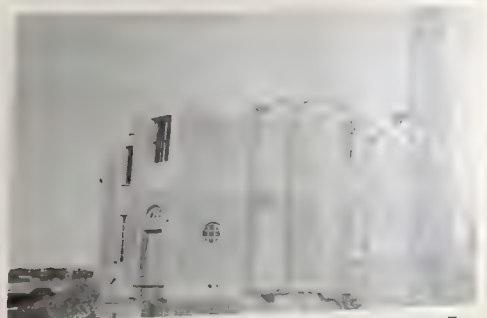
*The North End and the Front
Nov. 22, 1932*



*The sign in the Lane - Nov. 26, 1932
(Mr. F. Jno. Bell's car; a bitter cold day)*



*Preparing the forms for the Cement
Roof. (Front to left, looking N.) - Nov. 26, 1932*



*The south end and the rear of the
building - Nov. 26, 1932*



The Front Door and Portico - Nov. 30, 1932



The S.W. Corner and Portico - Nov. 30, 1932



*The S. part of the rear.
Nov. 30 (men pouring cement roof)*



The Rear of the Building. Nov. 30, 1932



*Pouring the Cement Roof. looking S.
(Front at right) - Nov. 30, 1932*



*Pouring the Cement Roof-looking N.
(Front at left) Nov. 30, 1932*



Front Door - Dec. 22, 1932



Portico - Dec. 22, 1932



*S.W. corner and Portico
Nov. 22, 1932*



South End - Dec. 22, 1932



*View of North End and Front.
Dec. 22, 1932*



*View of South End and Rear
Dec. 22, 1932*

Photographs taken about Dec. 27, 1932



Mrs. D.M.D. at Portico



Mr. and Mrs. D.M.D.



Front of Building



Another view of Front





No 888.

Shutters Closed *Jan. 1933*
Views of the two Domes 18 feet in diameter, being the north and south domes for the Administration Building



No 889.

Shutters Open *Jan. 1933*
The two domes in the shop at Walker Gate, Newcastle



Administration Building, Front and South end Jan. 16, 1933



Administration Building Rear and North end Jan. 16, 1933

University of Toronto
TORONTO 5, CANADA

DEPARTMENT OF ASTRONOMY

January 21, 1933

Dear Mrs. Dunlap: —

As I told you over the 'phone this morning I have had discussions with Mr. Bell about the big disc of glass, and I prepared a memorandum for him giving my views on the present condition of affairs. He prepared a memorandum for the firm in England and sent me a copy. Herewith I enclose copies of the two documents. They summarize the negotiations up to the present. I hope to have something definite to report on your return from Bermuda.

Very sincerely yours

C. A. (hant)

P.S. I am sending similar copies
to Mr. Holden.

C. A. f.

University of Toronto
TORONTO 5, CANADA
DEPARTMENT OF ASTRONOMY



Mrs. D. A. Dunlap
93 Highland Ave.
Toronto 5.

Photographs taken February 16, 1933



View facing East. Feb. 16, 1933



View facing Southeast. Feb. 16, 1933



View facing Northeast Feb. 16, 1933

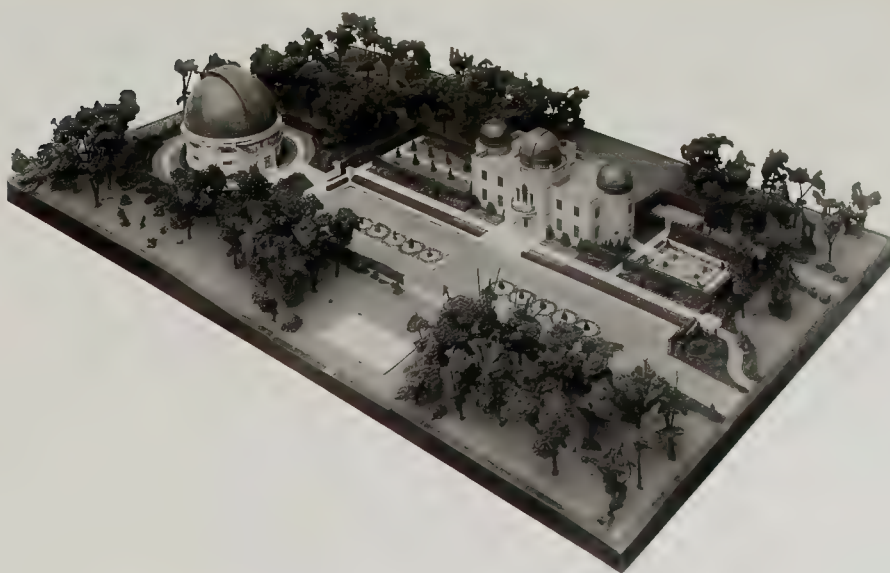


Sir Frederic Stupart



C.A.C.

two views of Portico - Feb. 16, 1933



Model of the Observatory constructed in the offices of Mathers and Haldenby and shown at the Exhibition of the Ontario Association of Architects, February 1933.



View of north end, and rear of building showing cases containing the three domes. Mar. 1, 1933



March 1, 1933

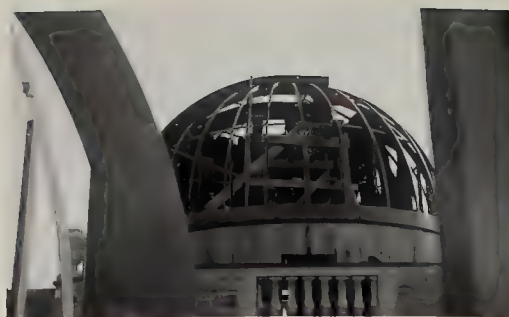


March 1, 1933

Two views showing cases containing the three Domes as they arrived from England



March 16
Central Dome. The Tarpaulin protects
the men from the Wind



March. 16
South Dome, as seen from Central Dome



Front of Building, showing Domes, Mar. 16



Another view of the Domes



North Dome, showing
cement wall and Dome
partly erected on it, Mar. 16



Preparations made to excavate for
pier for 74-inch Telescope, Mar. 16, 1933

Photographs taken March 22, 1933



Dropping a heavy weight
on the frozen earth to break
it for the excavation for
the great pier.



The bucket of the steam shovel
descending for a load of earth



The loaded bucket rising - looking N.E.



The bucket in the air - looking N.W.



The ironwork of the Domes erected



The Library on Mar. 22

Photographs taken April 4, 1933



Covering the S. dome with
"agasote" and copper



Another view of S. dome
C.R. Chalker (inspector, left)
E.W. Haldenby (right)



Reinforcing iron at bottom
of excavation



Ready to pour the cement



Bringing in water from
Richmond Hill through
the mud



View of steam shovel, and materials for cement
and excavated earth, looking N. April 4, 1933

Photographs taken April 21, 1933



Two views, showing the North and South Domes almost completed and the centre one well forward



The North dome : the cement wall yet to be covered with copper



C.A.C. beside one of the stone urns, the north dome in background



*Looking southeast
Two views showing progress on the cement pier for
the 74-inch Telescope*



Looking Northeast

Four upper photographs taken April 27, 1933



View showing domes completely covered



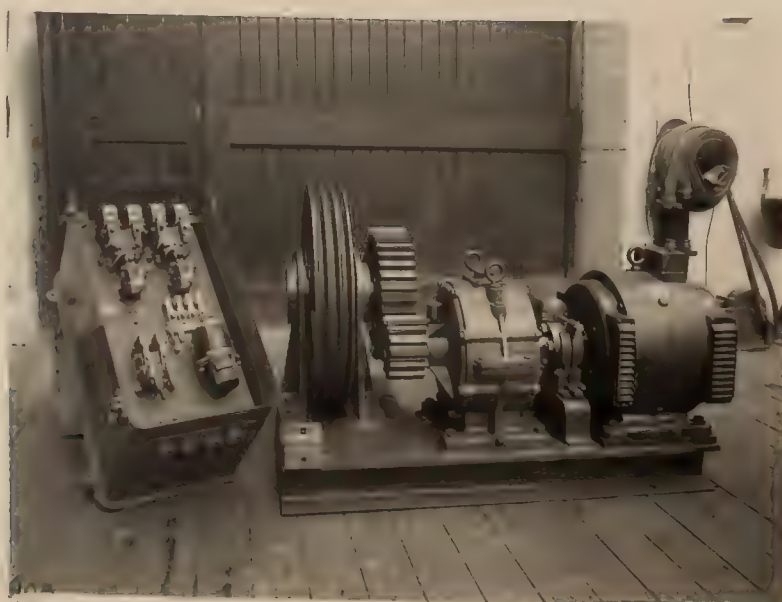
Steel rods on left, are reinforcements for elevator pit, south part of pier up to surface of ground S.E. view



Showing W. side of N. part of pier. looking N.E.



North end and part of front of Building



No 408 Control box

Reducing gear

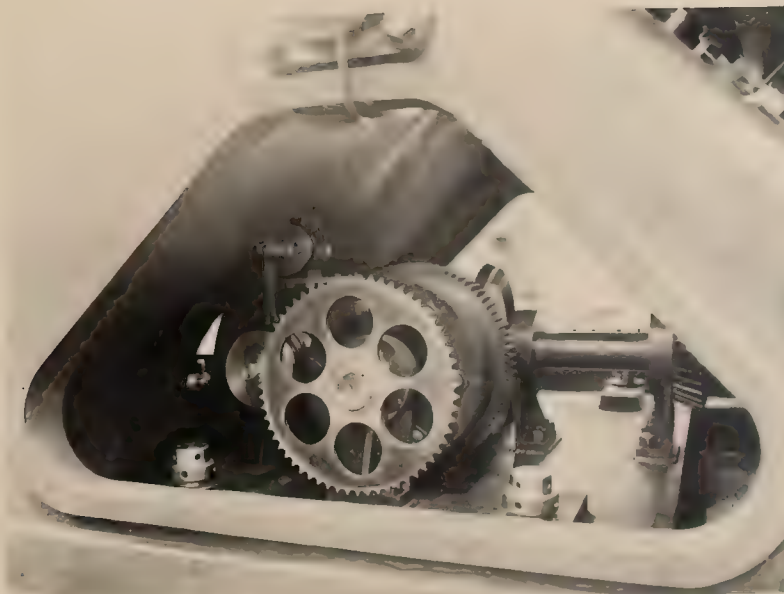
Motor

Gear for turning the 61-foot dome - (in England, April 1933)



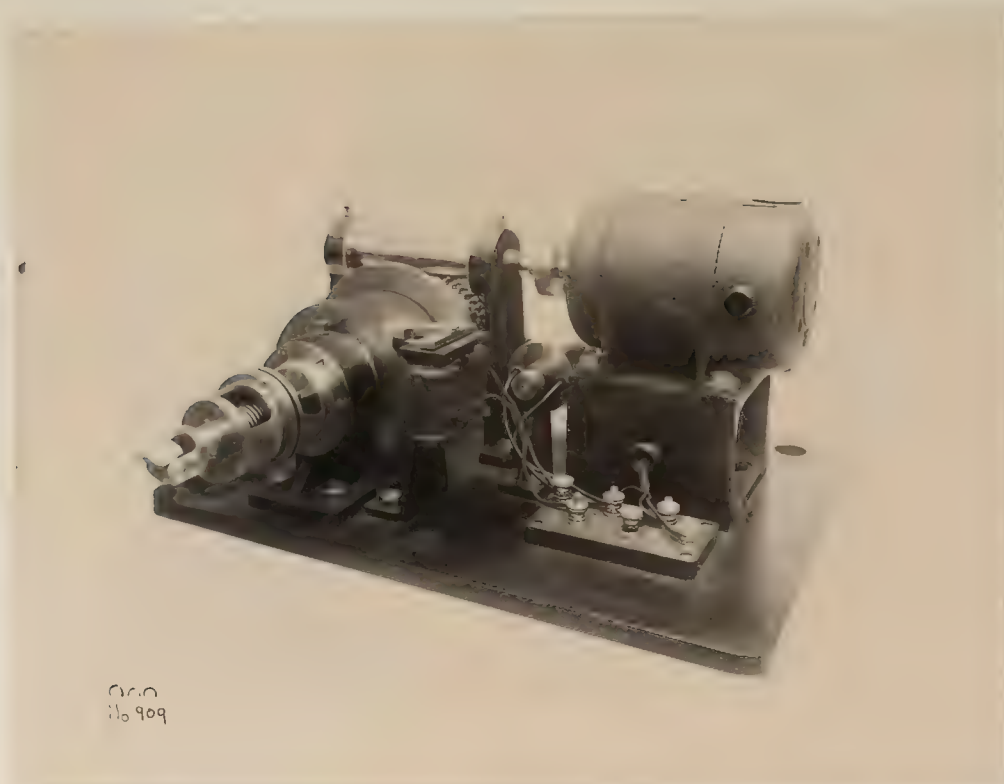
No 915

Gear for quick motion in right ascension, View from motor end. April 1933

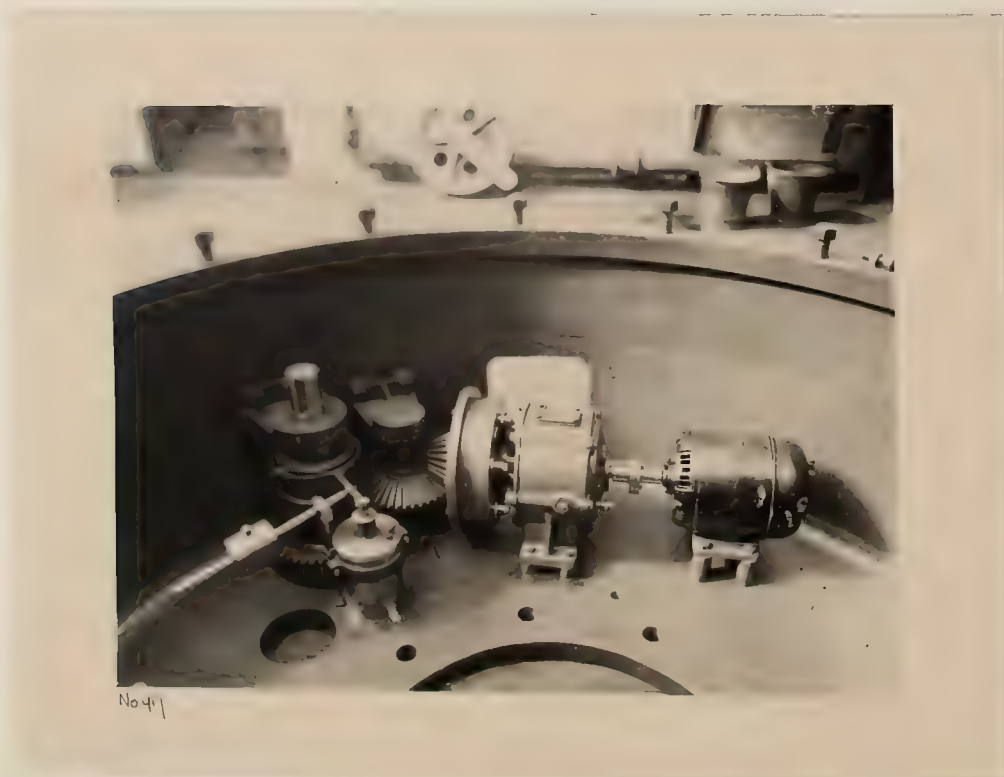


No 916

Gear for quick motion in right ascension. View from sprocket wheel and showing solenoid - April 1933



Gear for two speed slow motion of telescope in declination - April 1933



Gear for quick motion in declination, assembled inside the counterweight - April 1933

Photographs taken May 10, 1933



Looking Southeast



Looking Northeast

The Administration Building



*Workmen covering with copper
the wall of the central Dome.
(from S. Dome)*



*Showing progress on the great
pier, looking S.E. Part of
wooden form removed.*



Looking Northeast

The circular wall is part of the form for the cement wall of the building - Progress on the Great Pier and the Wall of the Building, May 10, 1933



Looking Northwest

Photographs taken May 16, 1933



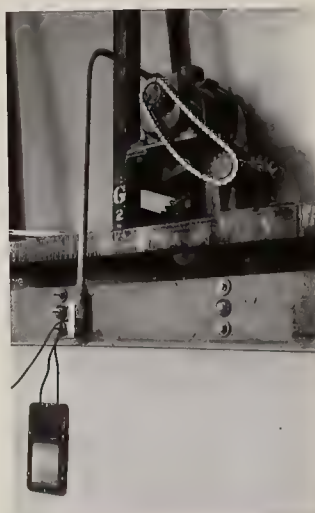
*Looking Northeast
Showing earth filled in around the great pier and progress made on the cement foundation for the wall of the Building*



*Looking Northeast
Showing progress made on the cement foundation for the wall of the Building*



*Looking Southeast
The circular wooden form is for the cement foundation of the building*



The electric motor for rotating the central dome.



Showing Progress on the Portico



The Southeast corner of the Building May 16, 1933

Photographs taken May 22, 1933



Looking N.W. Steam shovel at left is filling in around the pier; that in front is being used for grading.



Looking S.E. Showing the pier rising above the ground, about 20 feet in this view.



Looking N.E. Circular foundation for building in front.



Steam shovel being used for the grading around the Administration Building.



Another view, facing East.



View of Building facing N.E. May 22, 1933

Photographs taken May 29, 1933



Facing N.W. From second story window of Administration Building.



Facing N.W. View from parapet of Administration Building



Facing N.E. The structure at left is for hoisting and pouring the cement.



Facing N. slightly E.



Showing walk to South Dome



No 921

View showing the polar axis of the 74 inch
Telescope with declination axis in place,
and centrepiece ready to be hoisted into
position

May 1933

The centrepiece in place
Newcastle-on-Tyne, May 1933



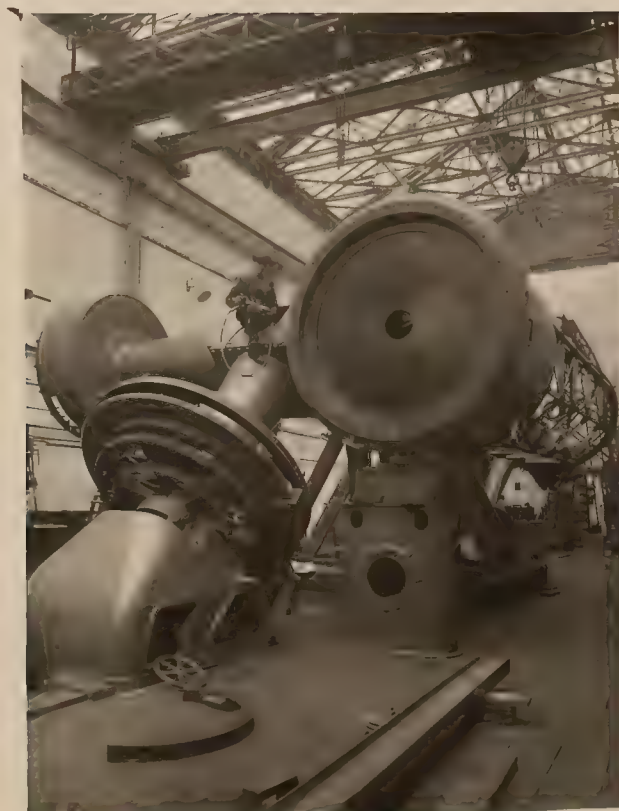
No 422



No 925

The 74-inch Telescope
View taken near the upper end of the
tube.
May, 1933

View from back of centrepiece. The
mirror cell is removed to show the
iris diaphragm.
Newcastle-on-Tyne. May 1933



No 925



Looking N.E. Progress on pier
June 1, 1933 J.A. Pearce



Looking N.E. Progress on pier June 1, 1933



At the Front Door June 1, 1933
C.A.C. Mrs. C.A.C. J.A. Pearce



The Pier on June 6, 1933



Sir Robert Falconer, A.D. Le Pan
June 6, 1933. Robert Falconer (son)



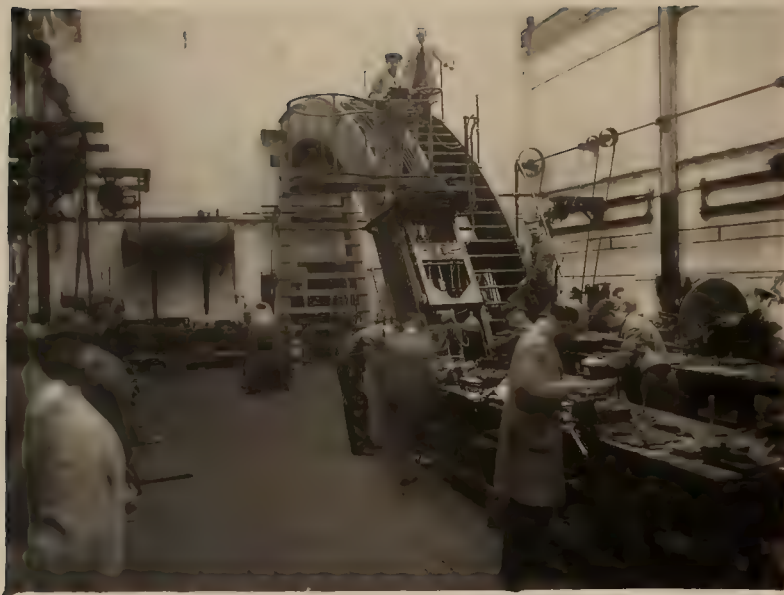
A.D. Le Pan ↑ C.A.C.
June 6, 1933 Sir Robert.



The Pier on June 13, 1933



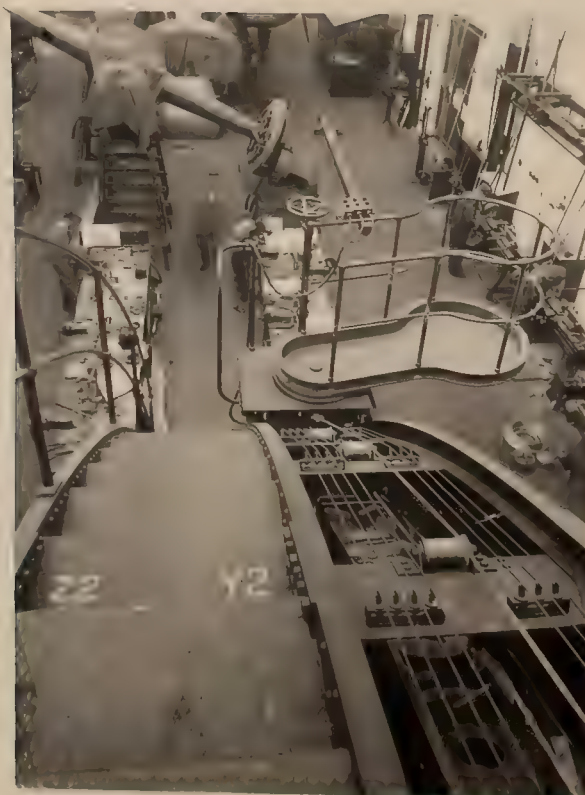
View of Pier from second story window of Administration Building
June 13, 1933



No 450

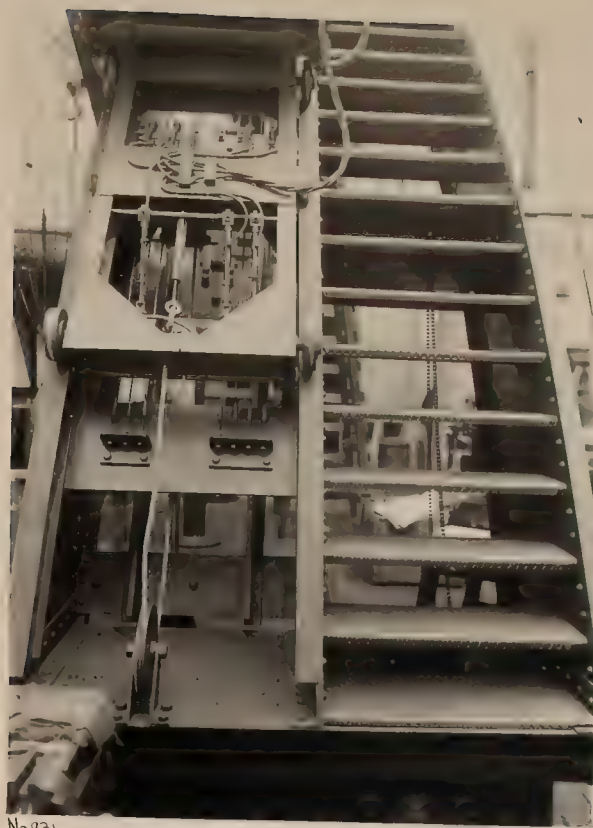
This bridge is to enable the observer to place himself near the upper end of the telescope tube.

View of the moving bridge for the 61-foot dome - This bridge may be moved to left or right and the observer on the carriage may be drawn up or let down. At Newcastle, May 1933



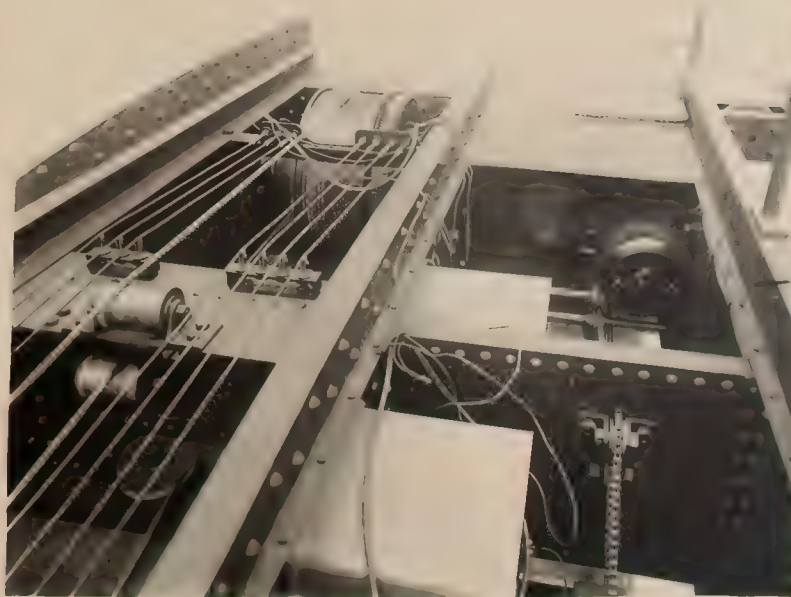
No 450.

View from upper end of bridge, looking down it. June 1933.



No. 431.

View of lower part of Bridge , below the carriage June, 1933



No. 432.

View of upper part of bridge, above the carriage . The drum and cable draws up and lowers the bridge . The electric wires are for the various motors June. 1933

Toronto Wireless of 1899 Sent Message Across Room

Prof. Chant Recalls First
Canadian Demonstration
at University.

SENT FIRST WIRELESS.



PROF. C. A. CHANT
68 years old yesterday, who directed
operations when the first wireless
message was sent in Can-
ada, June 1, 1899.

On the eve of his 68th birthday which he observed yesterday, Professor Clarence A. Chant laughed merrily when he was reminded of an important experiment in 1899, during which the first wireless telegraph message was transmitted and received under his direction.

That message was carried a distance of only a few feet from one side of a lecture room to the other, but it received headlines in newspapers and the publicity of the assembled assembly of the Canadian Institute.

It was in a lecture room of University College and under the direction of C. A. Chant, then a lecturer in physics and now professor of astronomy, that "wireless" made its debut in Canada. The first word sent spelled "university," and a few days later the first wireless telegraph message was sent from the University of Toronto to the University of Ottawa.

In the issue of The Mail and Empire of Nov. 25, 1899, Prof. Chant made a report on the wireless telegraph and of the future of the telegraph and of the future of the telegraph and of the future of the telegraph.

"Telegraphing without wires," as the "chief" of the report, as it is stated that on the previous 11th day Prof. Chant's assistant, as been completed and the first message transmitted and received across the "cableless" distance of 100 feet in the west wing of University College.

The apparatus which carried the message was the first of its kind in Canada. The messages were taken on a "cableless" device and the first word spelled out was "university," the first word.

In a report, appearing in the Mail and Empire of Nov. 25, 1899, Prof. Chant's report on the wireless telegraph is reprinted. It states in part: "Mr. Chant expressed doubts as to whether it would be possible to send messages in the way of a rock the ocean although as he pointed out it is necessary to speak with much caution of anything relating to the wireless telegraph."

"Yes," he added, "we've gone a long way since then. I don't know how far we've gone, but I think the apparatus is now a great deal better than it was then. And I believe that now we can send the first message sent in Canada."

Mail and Empire
June, 1933



E.W. Haldenby, R.L. Blackburn
Mrs. Dunlap, Mrs. Blackburn
Mrs. W.R. Hodge, June, 1933

DUNLAP OBSERVATORY NEARING COMPLETION

Work on Telescope is
Progressing in Two
Hemispheres.

TO USE NEW GLASS

Steel Building, Built in
England, Dissembled
and Shipped Here.

While workmen of two continents are engaged on the intricate work of producing the separate units which will some day join to form one of the most powerful and delicate reflecting telescopes in the world, the future home of the giant instrument is nearing completion just outside Toronto.

Only the finishing touches are needed to complete the administration building at the David Dunlap Observatory on North Yonge Street. Already the huge concrete pier, a masterpiece of concrete construction, extending from 30 feet underground to 50 feet above, awaits the gigantic telescope. Within a month, the circular steel building which will house the entire assembly, will be nearing completion.

This building alone, and its circular revolving dome, weighs 180 tons. The unit was built and assembled in England. It was then dissembled, packed and shipped and will arrive at a Toronto dock before July 1. It will then be trucked to the site and re-erected and will be completed in August. It will then encompass the lunette concrete pier on which the delicate telescope will repose.

Work on the latter is progressing in two hemispheres. In England the construction of the telescope is being pressed. After experiments in England to produce glass for the large mirror, observatory officials learned of research in the United States on pyrex glass for such mirrors. As a result, plans made three years ago when the order for the telescope was given for the mirror's production from a type of crown glass were abandoned.

Pyrex glass expands much less under rise in temperature than any other form of glass and is very desirable for such mirrors. The mirror to the big telescope must be ground from a sheet of glass 76 inches in diameter and 12 inches thick, and work is now progressing on this sheet of glass at the Corning Glass Works, Corning, N.Y. It will then be shipped to England for grinding and polishing and then will be shipped to Toronto for fitting in the telescope mounting, which by Summer of 1934 will be ready to receive it.

Some idea of the size of the mirror is evidenced from the fact that it will weigh two and one-half tons. The telescope itself will weigh 58 tons when mounted on the pier, yet the entire instrument will be balanced and adjusted to the accuracy of a chronometer. The mounting for it will be shipped from England in August, accompanied by a representative of the manufacturers who will superintend the task of erecting and adjusting it.

In the small dome of the observatory the 19-inch telescope is already installed but not yet mounted. The refracting telescope for the central dome and the astronomical cameras for the north dome will be supplied later.

Mail and Empire
June 19, 1933

WORK CONTINUES ON OBSERVATORY

Steel Building to House
Telescope Comes From
England

LARGE PIER ALSO READY

Considerable progress has been made recently on the David Dunlap Observatory. The stout administration building is almost completed. On the exterior some iron railings have yet to be put in place and a little painting has to be done. The grounds about have been graded. Within the building, a small amount of carpentering work and painting remains. The sixteen-inch refracting telescope, which will occupy the south dome of the building, is within the dome, but not yet mounted. The refracting telescope for the central dome and the astronomical cameras for the north dome will be supplied later.

Building and Dome by Sea.

The cement foundation for the six-story circular steel building, which will house the great reflecting telescope, is now ready. The building itself and its revolving dome were constructed and temporarily erected in England. They were then taken to pieces and put on a steamship, which will deliver them at the dock in Toronto before July 1. They weigh 180 tons. They will be taken by truck to the site and erected. The building will be completed in August.

The immense pier for the telescope is also ready. It extends thirty feet into the earth and rises about fifty feet above the ground. It is a masterpiece of concrete construction. The mounting of the telescope is ready and will be shipped from England in August. It will be erected on its pier in September. Although the telescope weighs fifty-six tons, it is constructed with the accuracy of a chronometer, and to mount and adjust it demands special skill and care. A representative of the makers will be on hand to direct operations.

Pyrex Disc Is Ordered.

Unfortunately the large mirror, which is the critical portion of the instrument, is not ready. The delay in producing it is due to the difficulty in obtaining a disc of glass for it. This disc has to be 76 inches in diameter and 12 inches thick, and its weight is nearly two and a half tons. When the order for the telescope was given three years ago it was expected that the mirror would be made from a kind of crown glass, but while the unsuccessful efforts to produce the disc were being made in England experiments in casting large discs from pyrex glass were carried out. This glass expands with rise of temperature much less than ordinary glass and it is much more desirable for a telescope mirror. When it was learned that a disc could be obtained from pyrex glass, the firm which is making the telescope, being anxious to supply an instrument of the very highest quality, ordered a pyrex disc. It is being produced at the Corning Glass Works, Corning, N.Y. and in September will be shipped to England, there to be ground and polished. If no special difficulties are encountered the mirror may be ready in the summer of 1934. When completed it will be shipped to Toronto and will be placed in the telescope mounting, and the instrument will begin its career of astronomical research.

Toronto Globe
Statement by C.A.C. June 20, 1933



Two views of the Great Pier
Looking N.W. from roof of administration Bldg.



Looking N.E. June 26, 1933

DUNLAP TELESCOPE MIRROR CAST TO-DAY

Donor Present—Grinding and Polishing May Take a Year

Corning, N. Y., June 21.—Casting of the disc of glass for the mirror of the great telescope of the David Dunlap Observatory at Toronto was made at Corning Glass Works here to-day.

The disc is made from a special type of pyrex glass and is efficient of expansion about 1-4 that of ordinary glass and is therefore especially suitable for a telescope.

The disc is 76 inches in diameter, more than 12 inches thick and weighs nearly 2 1/2 tons. It was cast at six a.m. and at 10 the cover of the mould was removed showing the immense block. It was then placed in the annealing oven where it will remain for four months. Then it will be shipped to Newcastle-on-Tyne, Eng., to be ground and polished to form the mirror. This process may take a year.

The operation was witnessed by a party from Toronto which included Mrs. D. A. Dunlap, donor of the great observatory, Prof. and Mrs. C. A. Chant and Prof. R. K. Young.

Telescope Mirror Cast For Toronto Observatory

(Canadian Press Despatch)

Corning, N. Y., June 21.—Casting of the disc for the mirror of the great telescope of the David Dunlap Observatory at Toronto was made at the Corning Glass Works here today from a special type of pyrex glass having a coefficient of expansion about one-fourth that of ordinary glass. It is especially suitable for a telescope mirror. The disc is 76 inches in diameter, more than 12 inches thick, and weighs nearly 2 1/2 tons. After being cast it was placed in the annealing oven, where it will remain for four months. Then it will be shipped to Newcastle-on-Tyne, England, to be ground and polished to form the mirror. This process may take a year. The operation was witnessed by a party from Toronto which included Mrs. D. A. Dunlap, donor of the great observatory, Professor and Mrs. C. A. Chant and Professor R. K. Young.

THE BARON STEUBEN CORNING N. Y.



The Globe, June 22, 1933

The hotel where we stayed.

The Star, June 21, 1933

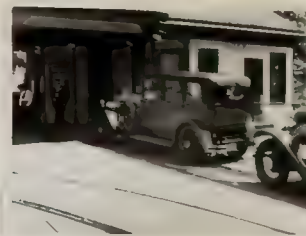
On the way to Corning N.Y. to see the pouring of the Glass Disc for the Mirror. June 20, 1933



On the Lewiston Bridge Can. End.



a little farther on the Bridge



on the Bridge, U.S. End.



Mrs. Chant & Mrs. Dunlap Your car
Mrs. Dunlap & Mrs. Chant
About to leave Dansville



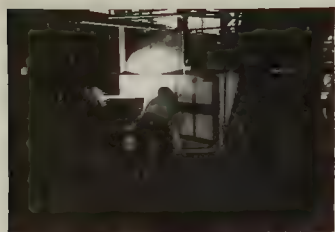
Main Street, Dansville
Looking N.

The above photographs were made from negatives taken with a Leica Camera.



Payne (chauffeur)
The main Street, Dansville, Looking S.

Views at the Pouring of the Disc for the Mirror



Taking a ladle of glass to the mould.



Thrusting the ladle into the mould.



Bringing back the empty ladle

The above pictures were taken about 7 a.m.



Pushing carriage (with mould) away after bee-hive cover of mould had been removed 1.30 p.m.



View of the glass disc in the mould 1.30 p.m.



Mrs. Clant F. Cameron Mrs. Dunlap Mrs. Hostetter



O.A. Sage P.H. Mitchell F.J. Bell Hostetter



O.A. Sage R.K. Young P.H. Mitchell F.J. Bell



R.K. Young F.J. Bell P.H. Mitchell

Views at the Breakfast table, Corning Glass Works
7.30 a.m. June 21, 1933



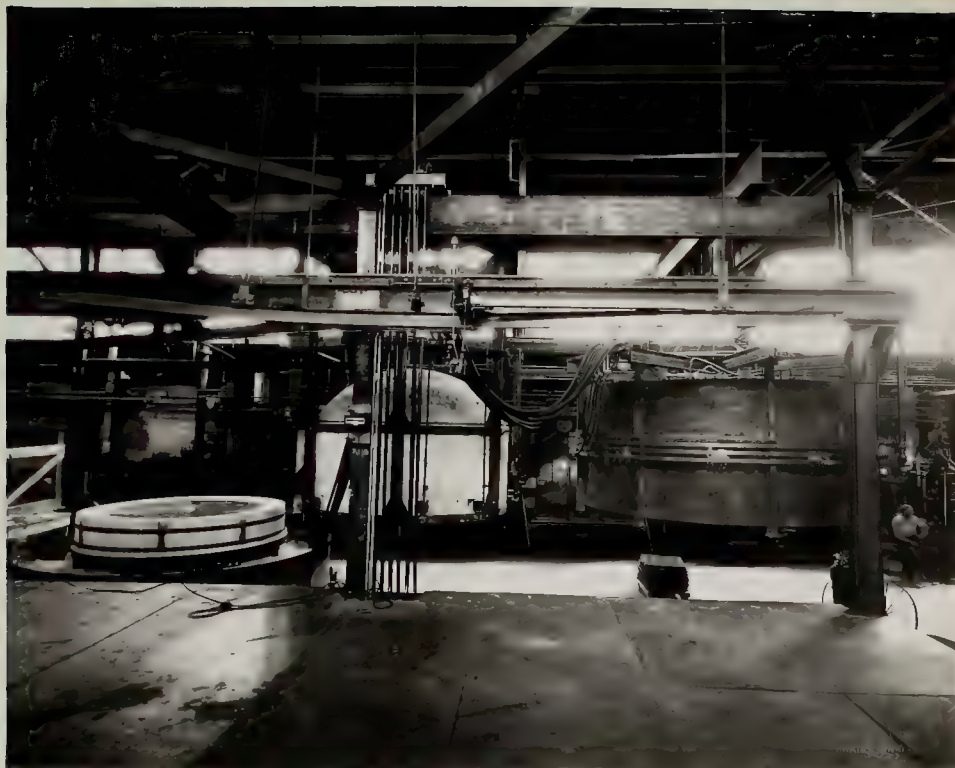
*View showing the interior
of the mould for the disc
of glass, and also the bee-
hive covering.*

*The core projecting
upward is to produce a
round hole in the disc.*

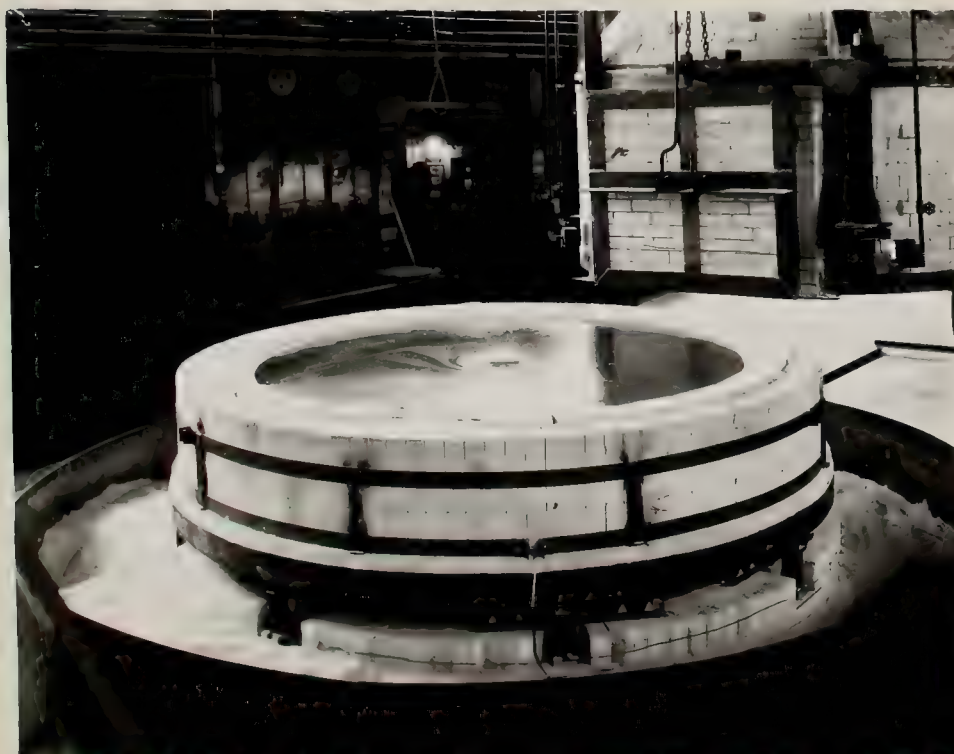
*View showing work-
men inserting a
ladle full of glass
and pouring it into
the mould.*



↑
R.H. Young



The 76 $\frac{1}{2}$ inch disc A Mould Cover An Annealing Oven
 View in the Corning Glass Works



View of the Disc, 76 $\frac{1}{2}$ inches in diameter and 13 inches thick, before being placed in
 the Annealing Oven.

Tracking the Stars!



Here is the giant telescope of the Dunlap Observatory, photographed during the course of construction at Newcastle-on-Tyne, England. While English workmen were fashioning this largest telescope in the Empire, Toronto workmen were constructing the base on which it will be placed at the Dunlap Observatory, North Yonge Street, shown at the right.



ADMINISTRATION BUILDING OF THE DUNLAP OBSERVATORY, NORTH YONGE STREET



THIS STONE WAS LAID BY DAVID MOFFAT DUNLAP SEPTEMBER 10TH 1932



Top—Professor C. A. Chant, who will be in charge of the observatory. Centre—The corner stone of the observatory, which was established as a memorial to the late David A. Dunlap. Below—The dome-topped observatory for away through the trees as seen by Yonge Street motorists—(Photo of Prof. Chant by Krueh.)

LARGEST telescope in Empire will bring the Heavens to Dunlap Observatory on North Yonge Street—Seen from Highway.

By KENNETH CRAGG.

Three rust-red, copper-sheduled domes, topping a stone building which stands on the highest level of a mound rising east of Yonge street and a short distance north of the Langstaff Jail Farm, give the highway motorist some conception of the wonder that is to be the David Dunlap Observatory, the latest major addition to the University of Toronto.

They give a conception only, for nothing but the foundation of the feature part of the observatory is visible at the present time. To the uninitiated, the concrete piers, in time to bear a telescope that is larger than any other with the exception of the giant instrument in the Mt. Wilson Observatory, California, might be anything from the base of a new-fangled hat case to the beginnings of a Hydro substation.

Within a few days a little tramp steamer will dock at the foot of York street. Out of her hold will be lifted the most essential part of this observatory, the great 48-ton telescope, largest telescope in the British Empire and second largest in the world.

It will be transported by motor truck, along with the parts of the 180-ton dome structure, from the quay side to the Markham Township observatory site and will be mounted on the concrete base. Months later—some time next year—the two and one-quarter ton mirror with its 74-inch concave face will be bolted to the base of the telescope. The observatory will then be ready for use and a long dream of Professor Clarence A. Chant, eminent astronomer and head of the Department of Astronomy, University of Toronto, will be realized.

Perhaps not fully realized, for deep down in that avant-garde heart is a hope, strongly rooted, that this observatory will make distinct contributions to science and that they are made during the period that it is under his skilled and devoted direction, all the better.

ready for use, some one will be on duty by the telescope every hour of a clear night. Every clear night the great shutters on the top of the dome will be opened wide and through the sixteen-foot opening, the skies will be raked by the telescope. Observations will be made and photographs will be taken.

The labors of the little group of scientists stationed there will not be halted by winter cold if conditions for observations are favorable. It can be cold there—ask the workmen engaged on the Administration Building who were up on that mound, exposed to every wind that blows, during the past winter. On clear winter nights when the wind will howl around and through the dome, muffled figures, with electric heating elements in their special clothing, will stand watch in the eerie light as the Northern Lights sweep up in great tongues of pink tinted fire.

There is in this story of the fulfillment of the professor's dreams, the element that makes "human interest" news. For the observatory is not only being erected in Dr. Chant's native township, but the place where this famous York County native son was born, some six-and-a-half miles away across country at Unionville, is visible from the roof of the Administration Building.

There are few stories of the star gazing of Clarence Chant as a youth. He had a close kinship with nature and the "forerunners of the angels" which nightly hung over his head doubtlessly impressed him and went him with good heart, up from Markham High School, through the University of Toronto and Harvard and later back to Toronto where he joined the University staff in 1911.

But beyond personal ambitions, there rests in the minds of Professor Chant and his colleagues, Professor R. K. Young, the determination to direct the observatory in such a manner that it will be a worthy memorial to the man whose name it bears.

Keep Constant Watch.

Once the equipment is in its place and

That is how discoveries are made, some times by tracking through the countless miles of space to a definite objective and sometimes through countless camera shots with the camera lenses fixed in time exposures on the telescope's mirror.

Professor Chant states there is a practical reason for all of this vigilance. The equipment costs too much to let it remain idle, an unnecessary hour and besides, while a large amount of the department's teaching work will be lifted from the half-planned observatory, located in the angle formed by University College and Hart House, to the Dunlap Observatory, that latter institution's chief function is astronomical research and study.

The late David Dunlap would have it so. He was a keen student of the heavens and astronomy and geology were his favorite studies. Several years after his death in 1924 Mrs. Dunlap became interested in the project and years before the formal announcement of the project was made by Professor Chant in December, 1929, definite assurances had been given that the memorial would be erected by Mrs. Dunlap and her son, Dr. Moffat Dunlap. In fact, the order for the telescope and its equipment was made in May, 1930, with the Sir Howard Grubb Parsons and Co., Newcastle-on-Tyne, more than 1400 miles prior to the announcement.

The location was fixed late in December, 1930. For some time it was believed that it should be on the Hill, a site of 100 acres, but it was found that the site was not suitable. It was then decided to build just south of the Hill and a quarter of a mile to the east of the highway. The site was then found in the development believed they had found

at least not it was determined that the railway running midway between the observatory and Yonge street, would cause serious vibration. It was ascertained that the observatory, 110, 75 feet above the highway, should have a foundation of ground more than 175 feet deep. Approximately 175 acres were secured for park land development, to give the institution an adequate setting.

Full appreciation of that setting is largely the result of a study made of the country from the roof of the Administration Building. It is a study of the ground, the hills, the valleys, the fields, the woods, the rolling valley land and the distant city.

Building Care Essential.

The circular concrete base, surrounding the telescope foundation is essential. No foundation was found, and the entire concrete work, from the solid foundation to the ground level, in the construction of the foundation and in the foundations of the Administration Building, care which has been given in any other type of structural work, kept the contractors longer than a business would anticipate. The Administration Building, however, was started in August. The interior is not finished.

Great difficulty was experienced in getting the material over the temporary platform was from Yonge street. Trucks were loaded and untimed horses, back-breaking labor was required to lift them up to the platform. The difficulty was overcome by the use of a crane, which was used to lift the material up to the platform.

The telescope is now on its way, three years after the order was given. But the

mirror is not with it. The original plans called for a 74-inch mirror, but it was found that a 74-inch mirror would be too large for the telescope. A 70-inch mirror was ordered instead. The mirror is now being ground at the University of Toronto. It will be mounted on the telescope when it is ready.

Professor Young was appointed to the University of Toronto in 1928 and moved there for his work. In 1921 he joined the University of Toronto and brought with him the newly completed 48-inch telescope, now in the observatory. It was the first time that a telescope of this size had been used in the Mt. Wilson Observatory.

Professor Young was appointed to the University of Toronto in 1928 and moved there for his work. In 1921 he joined the University of Toronto and brought with him the newly completed 48-inch telescope, now in the observatory. It was the first time that a telescope of this size had been used in the Mt. Wilson Observatory.

Telescope Is Gifted.

The telescope is now on its way, three years after the order was given. But the

the basement of Baldwin House, St. George and College streets, and after long weeks of labor, the telescope was completed. It had the distinction of having the largest pyrex mirror of any telescope up to that time and of being the second largest telescope in Canada.

This instrument, which will fill one point in comparative greatness when the new telescope is completed, is already established in the south dome of the new observatory, although it will not be ready for use until autumn.

When the Dunlap telescope is completely installed its makers claim that the surface of the moon can be made to appear no more than 50 miles from the eye of the observer. An object the size of Hart House, they say, would stand out clearly while the mounting and major deep-etched areas on the moon's surface would loom out in heavily defined lines.

While the mirror is, undoubtedly, the heart of the telescope, the mounting is of supreme importance. The ends of the main mounting, called the polar axis, rest in bearings that are fastened to the concrete pier already in place. Near the end of the polar axis that rests on the lower pier, a cross arm is fixed at right angles that is known as the declination axis. On one end of the declination axis is bolted the telescope and on the other end is bolted mechanism which also serves as a counter weight to offset the weight of the telescope.

The purpose of the mounting is not only to support the optical parts, but to enable the observer to train the telescope on any celestial object and to hold it in the field of view for examination or photography. This is a difficult matter for an astronomical telescope because the earth's rotation makes all stars that objects appear to rise in the east, pass over the sky and set in the west, and to follow this, it becomes necessary to move the telescope by clockwork or electric motor to follow the stars.

The instruments are balanced so evenly on either polar or declination axis, that a comparatively slight pressure will move them. By the employment of the two axes, at right angles to each other, the telescope can be pointed on any point in the sky.

Beyond the mirror there appears to be little more in the telescope. The reflecting element and type that is commonly known in the first case, the rays of light are received and focused by a concave mirror or system of mirrors, and in the other, the light rays are gathered by a lens. Telescopes of the first class are called "reflectors" and second, "refractors."

Face Is Renewable.

The "reflector" type, light from a star passes down the open tube to the large concave mirror. It is reflected back from the silvered surface to a converging focus point at or near the forward end of the tube and is then intercepted by a small flat mirror placed at a short distance inside the focus and directed to the eye piece or camera placed there to receive it.

It is an old story that the big mirror is made from a section of the great concave face of the telescope. Light rays of a distant star focus on the inner surface of the telescope. The concave face is focused on the inner surface of the telescope.

More, besides that that broadly defined plane gives an accurate conception of the operation of a telescope. But it does not take in anything of the innumerable pieces of apparatus that are used in the process. It does not even take in the electric motor that turns the dome around on its track so that the opening in the top can be placed in any quarter desired by the operator.

It is a place for the electrically operated elevator that is used every time the mirror has its face washed. Silver tarnish quickly and telescope mirrors must shine at their best at all times. If discoveries are to be made. And it isn't like lifting down the bathroom mirror when this 5,000 pound telescope glass is taken out of the base of the telescope. At the Dunlap Observatory it will be used out of the base by the elevator. The old silver on its face will be removed and a new coating applied, right on the premises.

More of the right on the premises work will be undertaken in the near future when officials will start making photographic test copies out of lenses and odds and ends, for use in the north dome at the Administration Building. So far, nothing but plans are in prospect for the centre dome. Some day, they hope to have a 10-inch lens telescope installed there.

At present there are no immediate prospects for the completion of that plan. But time and the collection of odd bits of glass and so on, will bring it to a realization. The department of the Department of Astronomy, University of Toronto. Friends of the department can't forget that 12-inch telescope in the south dome.

201 Hamilton Ave.

June 28th 1933.

Dear Mr. Dunlap:

Yours truly, W. H. Holden

Just left coming back 20 - Mrs. Dunlop

and her - in the 5 - 10 - 15 - 20 - 30 - 40 - 50 - 60 - 70 - 80 - 90 - 100 - 110 - 120 - 130 - 140 - 150 - 160 - 170 - 180 - 190 - 200 - 210 - 220 - 230 - 240 - 250 - 260 - 270 - 280 - 290 - 300 - 310 - 320 - 330 - 340 - 350 - 360 - 370 - 380 - 390 - 400 - 410 - 420 - 430 - 440 - 450 - 460 - 470 - 480 - 490 - 500 - 510 - 520 - 530 - 540 - 550 - 560 - 570 - 580 - 590 - 600 - 610 - 620 - 630 - 640 - 650 - 660 - 670 - 680 - 690 - 700 - 710 - 720 - 730 - 740 - 750 - 760 - 770 - 780 - 790 - 800 - 810 - 820 - 830 - 840 - 850 - 860 - 870 - 880 - 890 - 900 - 910 - 920 - 930 - 940 - 950 - 960 - 970 - 980 - 990 - 1000

ref. to a most interesting evening seeing the structure of Dunlop Farm. It has an imitation of the house standing to the rear and then have both been on long but it

Note from C.A.C.

I went up to the Observatory Monday and according to arrangement Mr. Holden came up to pick me up and take me home. He was rather late in account of the thunderstorm and did not arrive until 5:45. He had with him his brother and Phyllis and Don. We looked around a little and he was favorably impressed with everything.

On the way home we discussed the water supply and also the entrance to the grounds. I told him I thought superintendent Le Pan had pretty fully investigated the former. He agreed and said he would instruct Le P. to proceed with the project.

As to the road, he expressed willingness to give a 99-foot roadway through his property and he wished me to get the Superintendent's office and Mr. Young (in my absence) to explain.



The Great Pier

June 19, 1933 (Leica Camera)

NATURE

JULY 22, 1933

Annual Meeting of the Royal Society of Canada

THE Royal Society of Canada held its annual meeting on May 18 20, at Queen's University, Kingston.

In his presidential address to Section 3 (Mathematical, Chemical and Physical Sciences), Prof. C. A. Chant dealt with the inception and construction progress of the new 74-m. telescope at the University of Toronto. After paying tribute to the generosity of the donor, Mrs. David Dunlop,

Prof. Chant told how the project is being carried out. The administration building is nearly completed and the great dome will be erected during the summer. The mirror is to be made of pyrex glass of especially low expansion coefficient, which is expected to be of great advantage. Delivery for grinding is promised by the makers in September.



July 11, 1933.

NEW OBSERVATORY IS BEING BROUGHT CLOSER TO COMPLETION



Most modernly equipped and to house one of the world's largest telescopes, the David Dunlap Observatory, situated at an elevation of 750 feet about half a mile southeast of Richmond Hill, is nearing the final stages of construction. Above are

pictured, at left, the massive cement base on which the big telescope will be erected, and, at right, the administration building, with its three domes brought across the Atlantic for installation. The observatory is a gift to the University of Toronto by Mrs. D. A. Dunlap in memory of her husband.

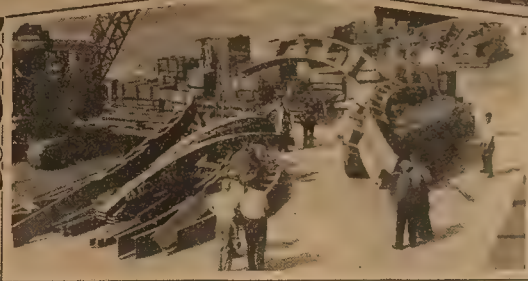
The Globe. July 25, 1933

TORONTO, WEDNESDAY, AUGUST 2, 1933.

SHIPMENT OF STEEL FOR DUNLAP OBSERVATORY ARRIVES IN CITY



The Globe
Aug. 2, 1933



A rather strange-looking collection of structural steel parts are piled on the docks at the Toronto Terminal Warehouse, long curved pieces, shorter curved pieces, straight sections, tapered beams, even a set of latticed steel stairs, the weight of the whole unique consignment being estimated at well over 200 tons. All pieces are marked "Made in England," and addressed to the Dunlap Observatory, Richmond Hill, and the arrival of this shipment means active construction work on the giant telescope and observatory which is being given to the University of Toronto by Mrs. D. A. Dunlap. Illustrations above by The Globe photographer show scenes on the dock as a derrick, aided by a crew of husky workmen and numerous trucks, started the job of transporting the steel. One of the many curved pieces being placed on a truck is seen at the upper left; the right-hand picture shows workmen handling another of the lengthy beams; while below is a general view of the 1,000-ton pile, an idea of the size of the consignment being gathered from the fact that everything's size in the photograph except the derrick is part and parcel of the shipment. Giant telescopes are intricate things.



*Photograph taken at
Toronto Waterfront*

Photographs by University Photographer



The flat boxes contain the Sheet Metal for covering the wall of the building and the "Agasote" for covering the revolving dome. About Aug. 2, 1933



Material for the Great Dome, on the dock at Toronto. About, Aug. 2, 1933

*Views of the Erection of the Circular Building and Re-
volving Dome. Aug. 8 and 16, 1933*



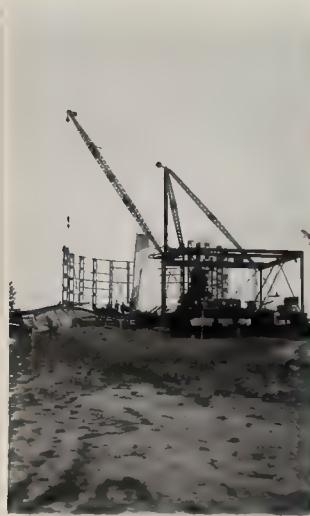
Looking N.E. Aug. 8



Looking N.W. Aug. 8



*Looking N.W. (from second)
story in Administration
Building. Aug. 8*



Looking North. Aug. 8



Travelling Crane. Aug. 8



*The wall erected, the base of the dome (on wheels)
in place ready to receive the framework. Aug. 16, 1933*



*The first curved member of
the dome being lifted to its
place. Aug. 16, 1933*

Photographs, August 16, 21, 25, 1933



*Portion of shutter in centre.
Aug. 16*



Aug. 16



*Sinking test hole for Water.
(100 ft. E. of Administration
Building went down 242 ft. un-
successful. Aug. 16*



Aug. 21



Aug. 25.



F.J. Bell, C.A. Chant Aug. 25.



F.J. Bell, H.P. Martin, Aug. 25, 1933

University of Toronto
TORONTO 5, CANADA

DEPARTMENT OF ASTRONOMY

Lothome Bay,
August 6, 1933

Dear Mrs. Dunlop, —

After a long wait I got the enlargements from my little negatives, and they are not quite as good as I thought they were. I enclose five of them and will show you the rest later.

I wrote Mr. Holden quite a long letter last week, and probably he has taken up with you some of the matters I discussed. They referred chiefly to the entrance from the south and the furnishings of the Administration Buildings, Maxwell and Shortt of the Superintendent's Office, and Young and I examined sketches of alternative routes and discussed them on the ground. We were unanimous in our conclusion that a road truly N.-S. from the centre of the great dome, and parallel to the front of the Administration Building was the best. Young and I had rather favored swinging to the east near the old pump and then



201 Madison Ave.,

August 20, 1933

Dear Mrs. Dunlop, —

Enclosed herewith is a copy of a letter to Mr. Holden, in which some account of the work on the Observatory is given. As you are quite as much interested I typed it so that you could have a copy.

Matters are proceeding rapidly and I shall be "on the job" all the time.

Also enclosed are some more enlargements from the Corning film. I had these made through J. F. Hartz Co., the agents for the Leica camera, and I think they are rather better than the previous ones. When I get out to see you we can discuss them together. I send also a print from a negative taken on Wednesday last.

I hope you are quite well; I know you are busy.

Very sincerely yours

C. A. Grant

P.S. I have had my first lesson in driving a car.

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is about 14 tons, with its
ro cranes, i.e., the one
I will lift it far up in
utter opening in the
was proposed to leave out
be brought within; but
can sent out from New-
method. I believe the
was expected, and hope

is for the dome, Sullivan
in the smaller domes, with
other tenders also.

the entrance has been sent
I wished to have some of
believe they did not have
fly. You will probably get
can get the sideroad
by which the municipality,

two floors of the dome and
if the telescope for the
Dr. Young's reflector
pleted. The latest news
he mould about the mid-lo

yours

To keep you informed
C.A.B.

August 28, 1933

Dear Mr. Holden,

The work on the big dome is proceeding rapidly and satisfactorily. Practically all the heavier structural part is done with the exception of the shutters. A multitude of bolts have yet to be put in, especially in the metal covering of the walls. The building looks fine. The gallery around it and the windows with their shutters add variety to the picture.

The Dominion Bridge Co. have sent in a tender for the erection of the telescope, which is now timed to arrive about the last week of September. They offer to do the work at cost plus 10 per cent., with a guarantee not to exceed \$2000. Mr. Bonus told the firm that this limit seemed rather high. They replied that the work is of such a special nature and the instrument is of such great value that they felt they should have that amount, but they said we should be charged according to cost. The firm had proposed to erect a second derrick like the "stiff-leg" one now in use, and to employ the two at the same time to handle the polar axis, which, I learn, with attachments will weigh 16-17 tons; but now they intend to install a derrick with an arm 92 feet long and able to carry the entire weight. They own the derrick. The erection of it, or of a second smaller one, is a considerable part of the cost. The big one is to be mounted on the ground, not on a platform, or is the one now in use. The ferral tender will go forward to you at once. The Superintendent's Office and I think it wise to accept it. Mr. Bonus says the firm is worthy of trust, and that from a progress report on the big dome the cost of erecting it will be well within the limit set.

Mullivan and Fried have not yet sent an estimate on the covering of the big dome. There is still some consideration of the best way to plaster on the inner coat of asbestos, but the estimate should be ready in a few days. The dome should be covered before the telescope is put in place.

The digging of the well is giving some anxiety. At a depth of about 15 feet boulder clay was struck, and it lasted until about 242 feet when shale was reached. After a few feet in this the digger, Mr. Davis, said he thought it unwise to go farther. He said that water obtained in a stratum of shale was nearly always salty and unusable. Professor Coleman could not be reached but his successor, F. Moore, was in town. He was taken up on Friday. He examined the samples brought up from different depths, and agreed with Davis that it would be useless to go deeper. Davis was instructed to pull up the 170 feet of casing which he had put down.

The question then was where next to sink a test hole. Davis suggested between the railway and the creek on Stephens's property, but was told that was out of the range of possibility. Exploration east near Bayview did not seem promising. It would seem that all the wells which were on the Observatory property, as well as those at the houses just north of us, are shallow and receive only the water which filters through the layer above the boulder clay. The boggy land on Bayview probably gets its water in the same way. There is little there now.

the creek flowing through into a much larger current a considerable depth with the best place to try it is about 75 feet below. Davis agreed that this try there is being sent in touch with your Mr. r is being sent as usual. you. I do hope we shall see to pump will be much

VanNostrand's men were

Y yours

C.A. Chant

UNIVERSITY OF TORONTO
TORONTO 5
DEPARTMENT OF ASTRONOMY

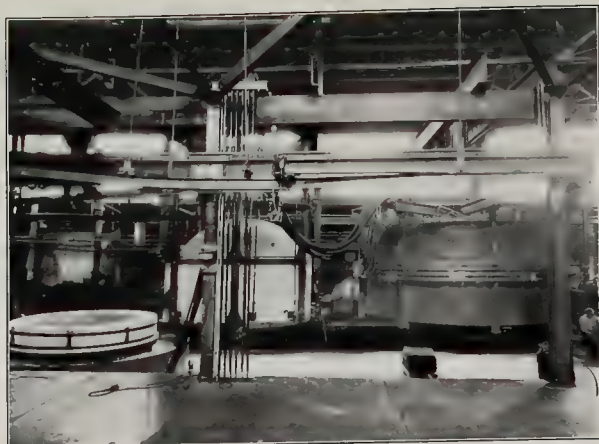


Mrs. D. A. Dunlop,

Don-Wald Farm,

Todmorden P.O.,

Ont.



IN THE GLASS WORKS
of the David Dunlap Observatory, at the University of Toronto, Canada.

284 Progress on the David Dunlap Observatory

in England. When finished the disc will have a diameter of 76 inches, a thickness of 12 inches at the edge and 11 at the centre.

I hesitate to predict when the mirror will be ready, but hope that there will be no unforeseen causes for delay and that it will be ready to be put in its mounting in one year.



Aug. 30, 1933



Dr. L. S. Bell, Winnipeg, Mrs. Dunlap,
Miss Cassie Bell, New York, Mrs. R. W.
Hopper, R. W. Hopper.



Mrs. Dunlap, Dr. L. S. Bell, C.A.C.,
Miss Cassie Bell.

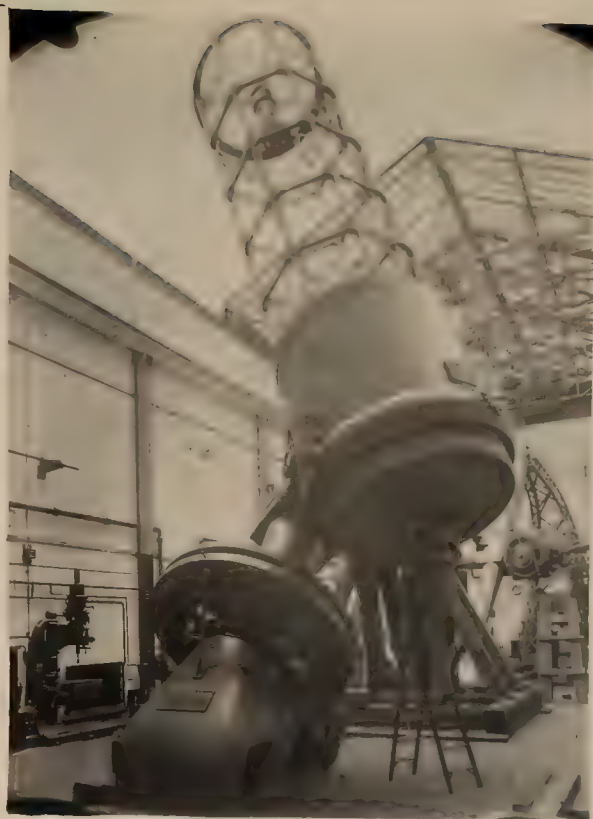


R. W. Hopper, Dr. L. S. Bell, Miss
Cassie Bell, Mrs. Dunlap.



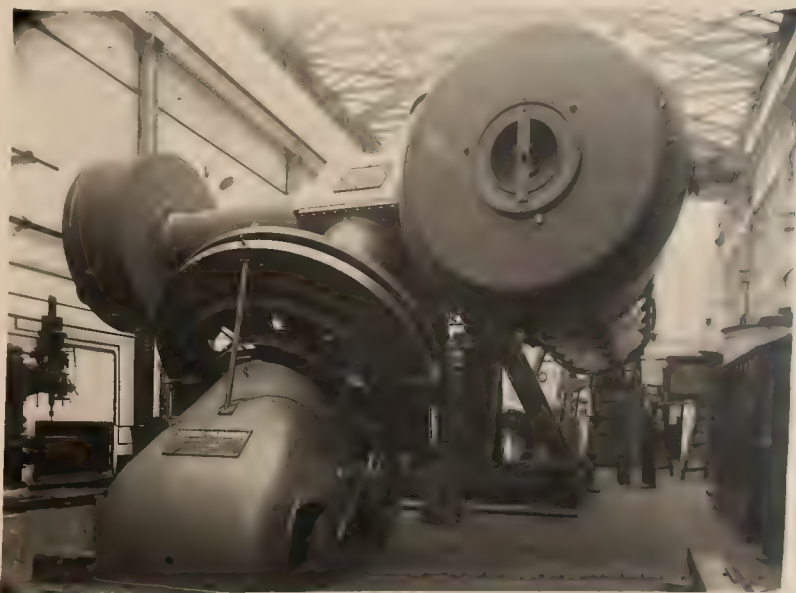
C.A.C., Mrs. Dunlap

Photographs at Observatory House
Sept. 3, 1933



No 940.

View of Telescope in Shops - Looking N.W. July 1933



No 943.

The completed Telescope - Tube horizontal, pointing North - July 1933.



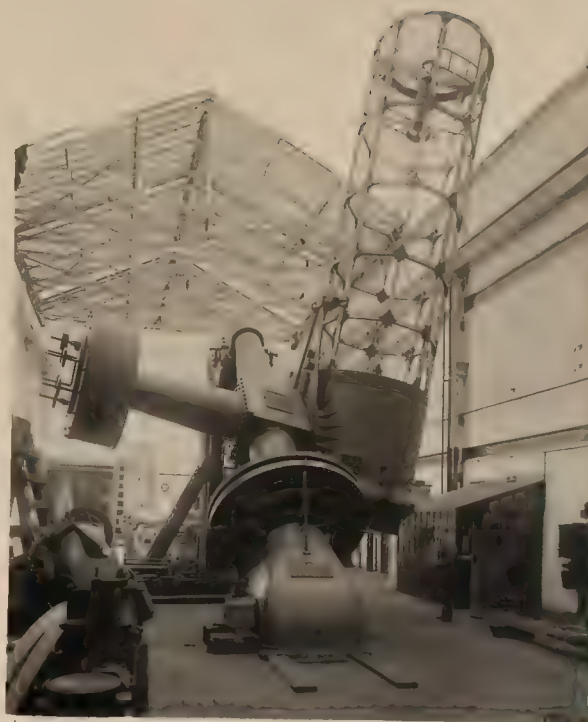
No 120

View of interior of the tube, from inside the centre piece ~ in the distance is the 36-inch telescope for Greenwich. July 1933



No 121

View of the Telescope, Looking N.W. ~ The tube is on the east side of the polar axis and is nearly vertical. Aug. 1933



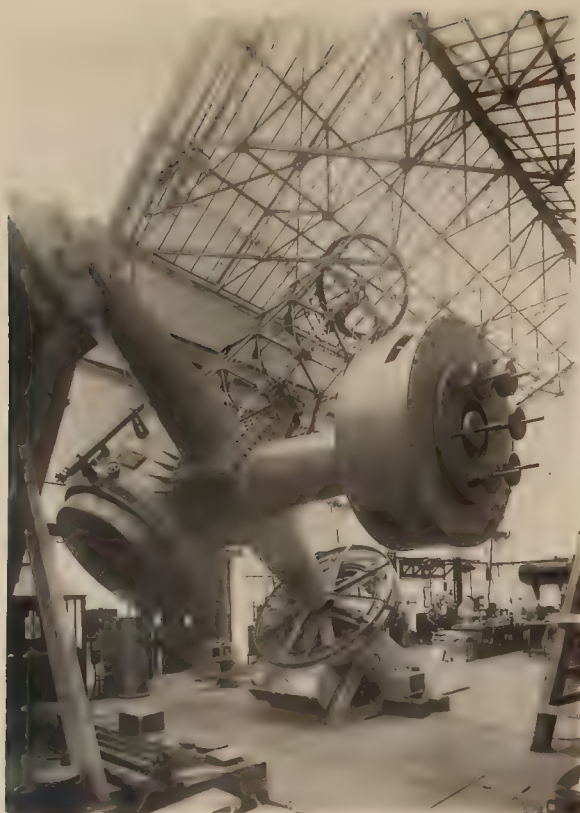
No 956.

View of Telescope, looking nearly North - The tube is on the east side of the polar axis which is turned slightly to the east. August 1933



No 757

The 74-inch Telescope and the 18-inch reflector for Dundee - Both are pointed approximately towards the pole of the sky August - 1933



No 455

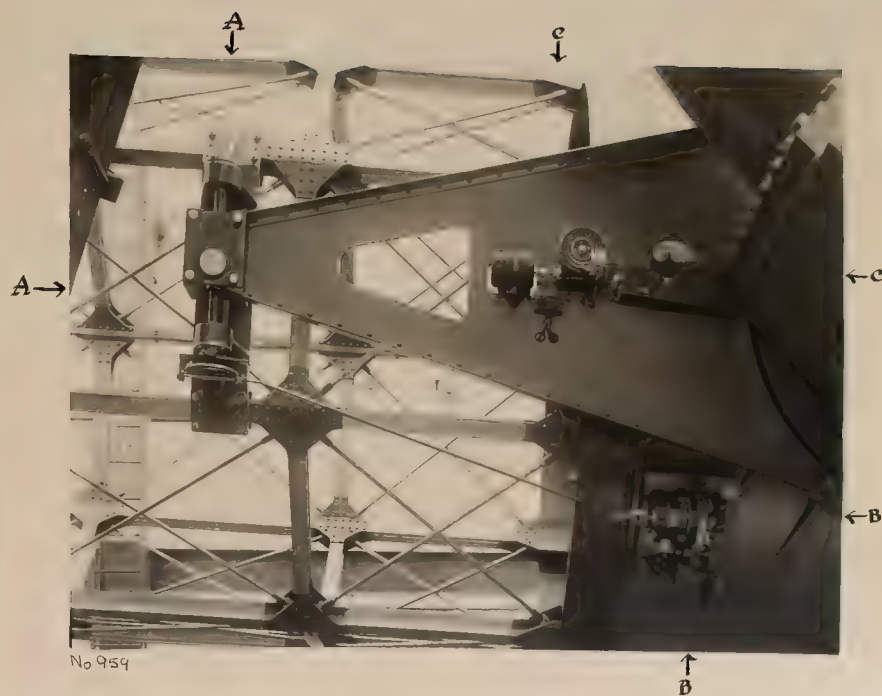
View of the Telescope, looking south-east.

The tube is on the east side of the polar axis and pointing approximately towards the celestial equator. August 1933



No 1-8

The driving Clock of the Telescope. August 1933



Close-up view of the Declination Clamp (A,A) actuated by the motor (B,B), also Gear for slow motion (C.C), Aug. 1933

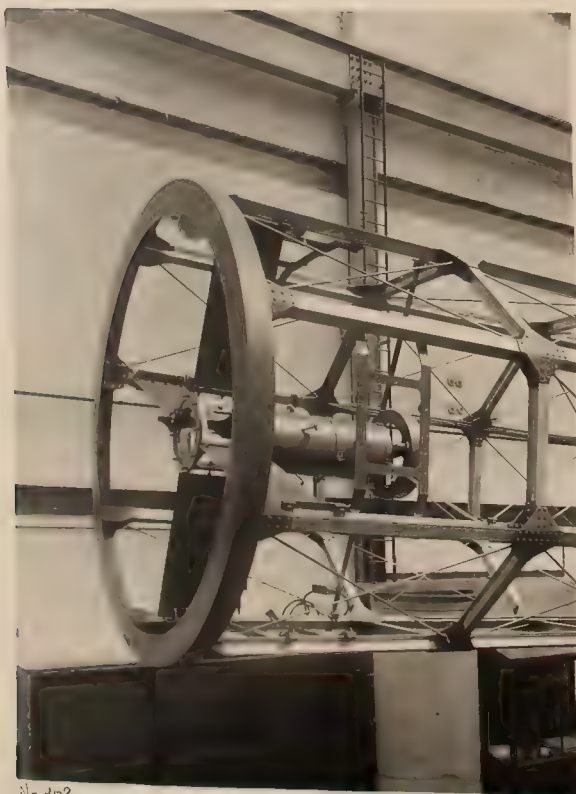


View of the iris Diaphragm, taken from inside the tube. Diaphragm nearly closed Aug. 1933



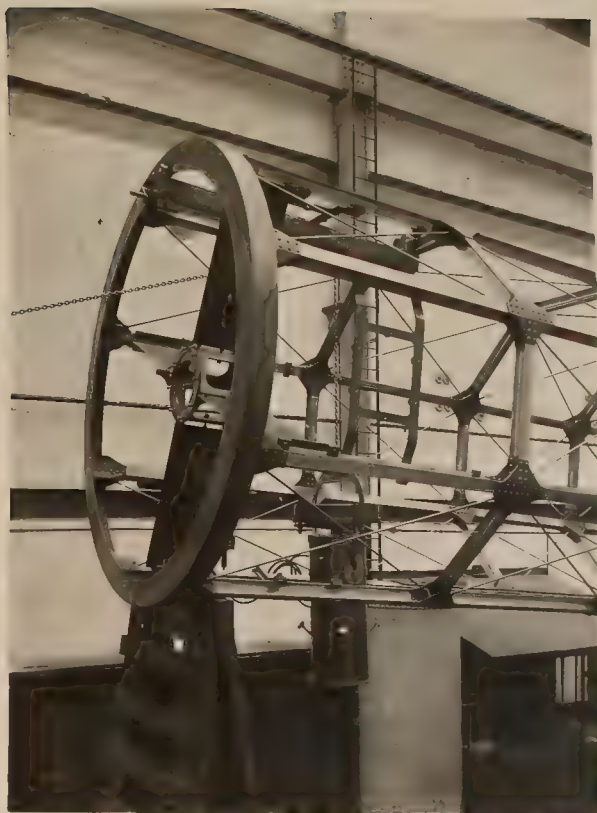
No 101

View looking down the tube from outside it. Aug. 1933



No 102

The holder of the Cassegrain Mirror, and gear for focusing it, in place at upper end of the tube. Aug. 1933



No. 463.

Cassegrain holder and focusing gear being removed from the top of the tube. Aug. 1933



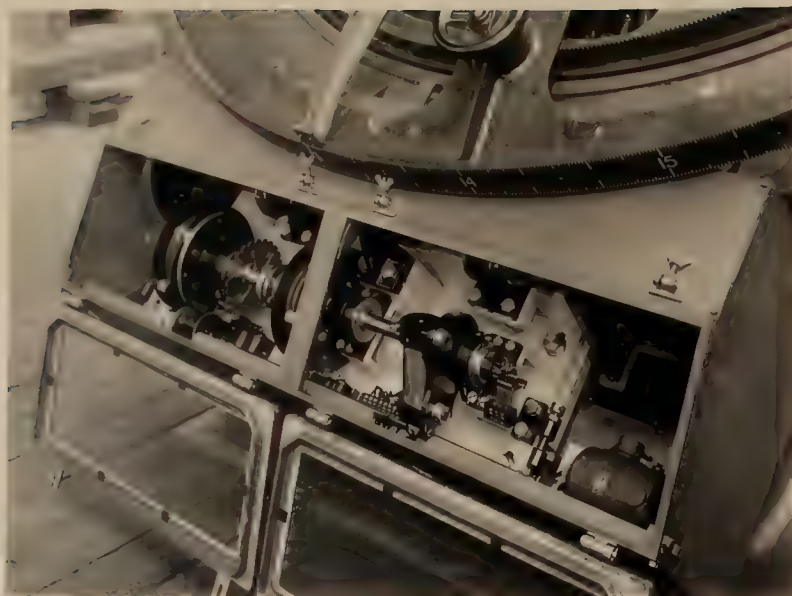
No. 464

Holder of Newtonian (flat) mirror in place at the top of the tube. Aug. 1933



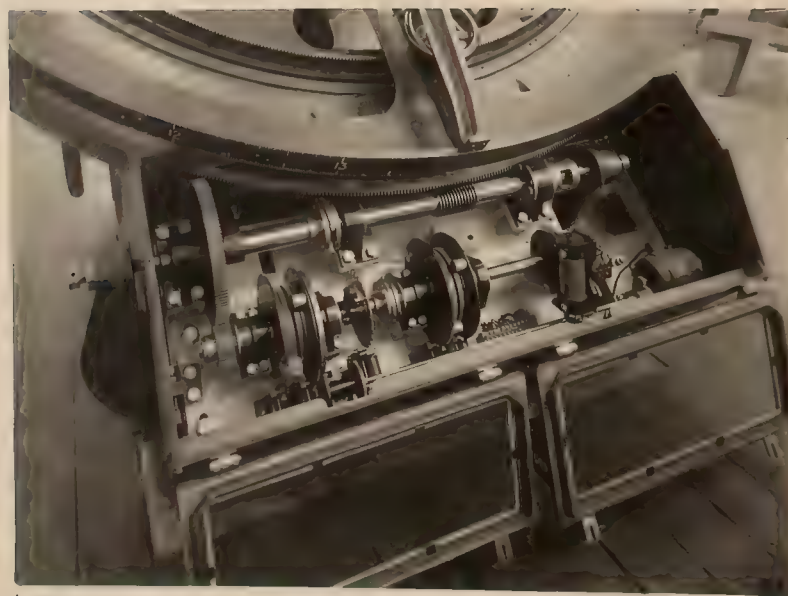
No. 465

Newtonian holder being removed from the top of the tube. Aug. 1933



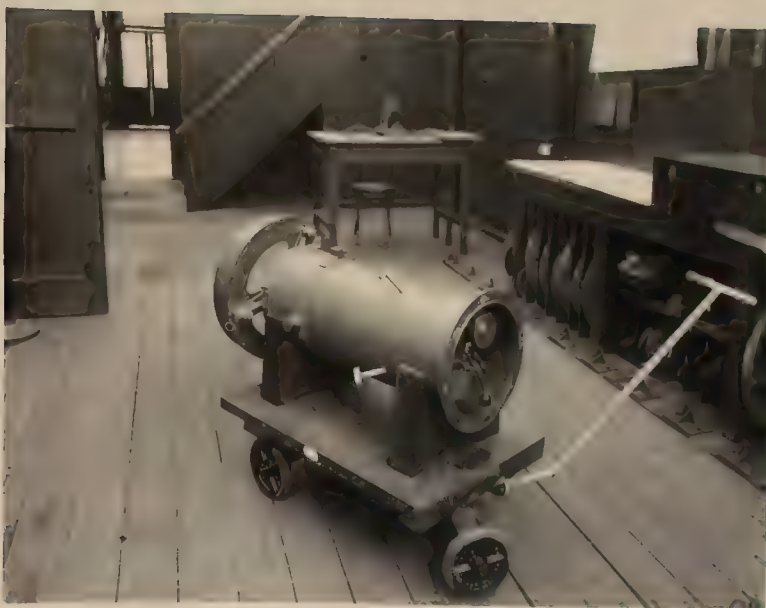
No 960

Gear Plate of the Great Telescope , View from Right. August, 1933



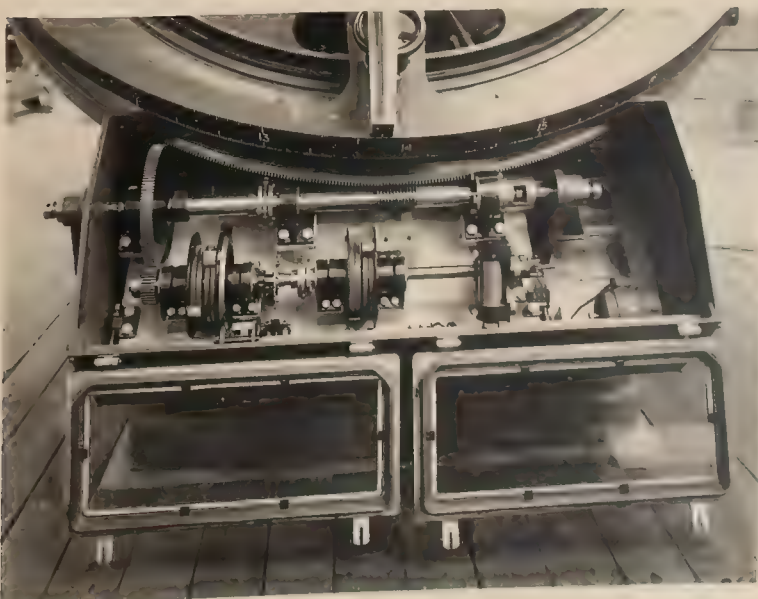
No 90.

Gear plate, View from Left. note the driving worm and the great worm wheel or driving circle (see pages 24.....) Aug. 1933



No. 981

The holder of the Cassegrain Mirror, with its Focussing Gear (within the tube) When not on the Telescope, it is kept on its truck (see pages 110, 111.) Aug. 1933



No 408

Sear plate, Front View. Aug. 1933



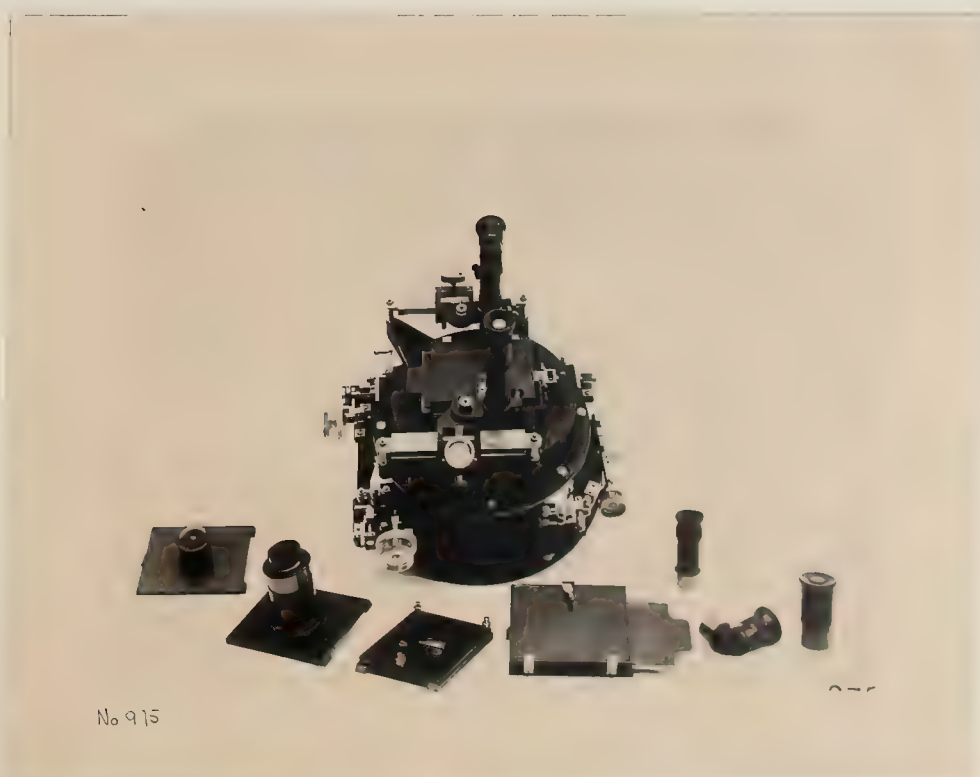
Newtonian Breechpiece, with large Plate Holder. Aug. 1933



Newtonian Breechpiece, with small Plate Holder and Guider-Eyepiece. The breechpiece is attached to the telescope tube near its upper end. It is used when direct photographs of a celestial object are being taken. Aug. 1933



Newtonian Breechpiece, with small Plate Holder and Guide Eyepiece. The finder Telescope is in the foreground. Aug. 1933



Newtonian Breechpiece, with all Plate Holders, Eyepieces and other attachments. Aug. 1933

Photographs taken September 7, 1933



Taken from the parapet of the Administration Building



*The above Photographs show the second Shutter being lifted into place.
This was the last work for the hoist shown in the pictures and it was dismantled
a few days later. September 7, 1933*



Work proceeding on the shutters. The old crane still in place. Sept. 12, 1933



The new 20-ton crane (at right) and the caterpillar crane. October 3



The Administration Building. Oct. 3



Showing progress in covering the dome with "Agasote" Oct. 3



Looking almost North. Oct. 7, 1933



Looking almost West. Oct. 7, 1933

These two views show progress in covering the dome.

From "Nature", London Eng.

OCTOBER 14, 1933

NATURE

589

590

N A T

The 74-inch Reflecting Telescope for the University of Toronto

SINCE November 1930, a 74-in. reflecting telescope, the largest in the British Empire and the second largest in the world, has been under construction for the University of Toronto at the works of Sir Howard Grubb, Parsons and Co., Newcastle-upon-Tyne, a subsidiary company of Messrs. C. A. Parsons and Co., Ltd. The observatory will be situated in a park of 177 acres on Richmond Hill, eight hundred feet above sea-level, a few miles to the north of Toronto and, in addition, to the 74-in. reflector and its dome, will comprise a block of administration buildings on which three smaller domes will be mounted. It is being built as a memorial to the late David A. Dunlap, the funds being provided by his widow and son, and, on completion, will be presented to the University of Toronto and administered by the Department of Astronomy.

The telescope is of the reflecting type with a clear aperture of 74 in. and is arranged so that it can be used either as a Cassegrain or Newtonian (Fig. 1). The mounting is of the modified English or "Campiate" type, in which the tube is placed on one side of the polar axis and the counterpoise on the opposite side.

THE MOUNTING

The polar axis is 22 ft. long, built up of two tapered tubular steel castings with forged steel pivots shrunk into the ends and held to a central tubular steel box. It runs in self-aligning radial ball bearings, with ball thrust bearing at the lower end housed in a cast iron base casting. The weight of the axis is nine tons. The bearing housing at the upper or North end is fitted with adjustments for the true alignment of the axis.

The driving circle, or wormwheel, is mounted on ball bearings on the lower pivot of the axis and can be clamped rigidly to the quick motion gear wheel, which is keyed to the axis, by a motor operated clamping gear. The circle is a steel casting and the bronze rim, in which the teeth are cut, is slightly shrunk on and fixed with a number of delta metal screws. The pitch diameter is 8 ft., cut into 996 teeth of 8 mm. circular pitch. The wheel was copied from a 4 ft. diameter master circle graduated on silver.

The sidereal circle is 8 ft. diameter and is strung

on the boss of the driving circle. It can easily be rotated by any one of six handwheels attached to it and fitted with pinions gearing into a toothed ring attached to the driving circle. It is constructed of fabricated steel with a delta metal rim 3 in. wide graduated on both edges to 1 minute of time divisions. The lower set of divisions

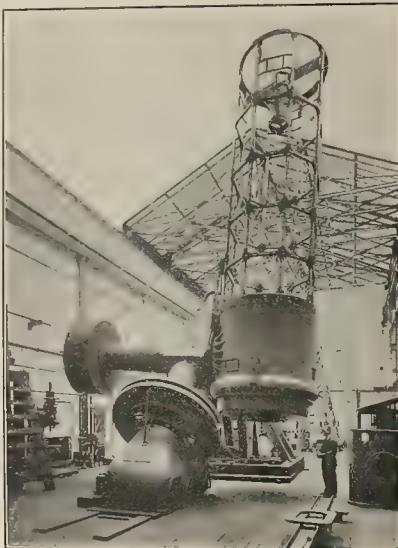


FIG. 1. The Toronto 74-in. reflector. View from south-east.

read against indexes fixed to the guard of the driving circle and mark sidereal time, while the upper set of divisions read against indexes attached to the polar axis and mark right ascension.

An hour angle circle is fixed to the quick motion gear wheel. The declination axis passes through the cubical centre section of the polar axis at right angles, supported at the outer end in a tubular tapered steel casting attached to the cubical portion of the polar axis.

The declination axis itself is of forged steel 13 ft. long, weighing 3½ tons and formed with a flange

at the inner end 3 ft. 5 in. diameter, to which the tube is attached. It is mounted in ball radial bearings, with a double thrust bearing at the small, or outer, end, to which is keyed the quick motion gear wheel and also the declination circle 6 ft. 3 in. diameter, graduated to 1° of arc. The gear wheel and circle are housed inside a drum-shaped casting attached to the end of the tubular support carrying the axis, this casting forming the counterpoise for the tube. The declination circle drives two drums 12 in. diameter, geared up 72:1, and arranged so that there is no backlash. These drums are graduated to 60 divisions, each division indicating 5 minutes of arc. The tube comprises three sections,

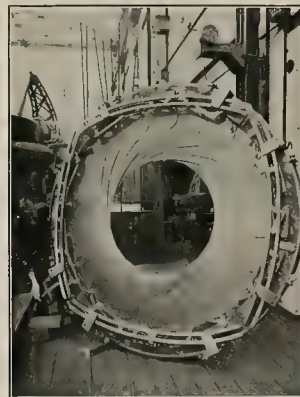


FIG. 2. The ins. day diagram. Half-aperture.

the lower portion a cell in which the main mirror is mounted, the central section which is bolted to the flange of the declination axis, and the upper, or skeleton, section for carrying the Cassegrain or Newtonian mirrors and the photographic breech-piece.

The main mirror cell is a ribbed steel casting fitted with nine circular pads for the back support of the mirror. The pads are mounted in groups of three on spherical seatings on three triangles, which are themselves mounted on ball-ended screws, by which the mirror can be 'squared'. The nine pads are so disposed that each bears its proper proportion of the load.

For the lateral support of the mirror, 18 weighted levers are disposed round the inside of the cell mounted on universal joints, the short end of the levers engaging in holes in brackets riveted to a flexible band loosely clamped round the mirror.

Description of the Telescope supplied by the makers.

This band has blocks fixed at intervals on its inner edge, loosely fitting into a groove in the edge of the mirror to keep it central. A large wormwheel is fitted to the back of the cell for carrying the spectrograph.

The centre section of the tube is a steel casting 7 ft. diameter weighing 5½ tons, and formed with a large boss on one side for attachment to the flange of the declination axis. Just above the lower flange, to which the cell is attached, the casting is swelled out to 8 ft. 7 in. diameter to accommodate an iris diaphragm (Fig. 2).

The construction of this diaphragm is generally similar to those used in cameras, but in order to prevent sag, due to the considerable weight of the leaves, and to keep the aperture central when the tube is in the horizontal position, it has been found necessary to guide the moving ends of the leaves between radius bars. The range of the diaphragm is from 12 to 74 inches aperture and is operated by a handwheel.

The upper, or skeleton, portion of the tube is octagonal in section, the main and cross members of which are constructed of 3-in. duralumin I beams, connected by steel gusset plates. Diagonal tension rods of duralumin, screwed with right and left hand threads, are fitted in each bay and tightened up to a predetermined tension, so that they are always in tension, whatever the position of the tube.

A fabricated steel box of square section, with circular flanges, is suspended in the centre of the upper end of the tube on four spring steel strips placed edgewise, to which the mountings for either the Cassegrain or Newtonian mirrors can be attached. Gear is supplied for handling these mountings and interchanging them conveniently and safely. The Newtonian mirror is arranged so that the reflected beam can be directed to any one of four positions round the sides of the tube, where frames are fixed to which the photographic breechpieces can be attached. This breech-piece comprises a focusing gear and plate holder with two guiding microscopes mounted on cross slides operated by micrometer screws. It is fitted also with a rotary motion to correct for rotation of the field. The plate holders take plates 3½ in. × 4½ in. and are interchangeable with a knife-edge focusing plate and with adaptors for oculars.

The driving clock comprises a heavy crossed arm governor driven indirectly by a ½ horse power direct-current motor through a differential gear box. A weight suspended from a chain passing over a pulley and attached to the outside of the differential gear box gives a constant driving force to the governor. The gear box carries an arm with a contact at its outer end passing over a number of studs arranged concentrically and connected to resistances in the field circuit of the motor. When the motor is running at a correct speed, the weight is kept floating, but if its speed alters, the gear box rotates together with the contact arm, which immediately corrects the speed of the motor by altering the field resistance. The governor runs at

50 r.p.m. and is directly coupled to a worm gearing into a wormwheel on a jack shaft making one revolution in 24 seconds. The jack shaft is mounted parallel to the shaft on which the driving worm is cut and geared to it by a pair of accurately cut spur gears. Thus there are only one pair of spur gears and one worm wheel between the clock governor and the driving screw.

A 'Grubb' type of electrical control operated by one of the observatory clocks, is incorporated on the jack shaft.

The right ascension and declination quick motions are operated by 1½ and ¾ horse-power reversible motors respectively and are arranged to rotate the axes at the rate of one revolution in eight minutes. The drives are taken through dog clutches operated by solenoids in parallel with the motor circuits, so that the motors and reduction gears are automatically disconnected from the telescope when not in use. A friction clutch is also incorporated and the switch gear controlling the motors so arranged that when the motors are switched off the dog clutches remain engaged for a few seconds, allowing the friction clutches to slip and bring the telescope to rest.

The declination clamp is placed between the tube and the polar axis, and consists of a fabricated steel arm about 7 ft. long mounted on a 'V' ring 4-ft. diameter attached to the side of the polar axis. This arm can be rigidly clamped to the 'V' ring by a toggle gear operated by a small electric motor. The outer end of the arm carries a nut mounted on a link motion and engages with a screw mounted in bearings attached to a bracket which is firmly clamped to the side of the tube. A motor operated two speed gear is connected to this screw for giving the setting and guiding motion in declination. This motor is controlled by two separate reversing switches. For the setting motion a dog clutch operated by a solenoid in parallel with the motor connects the gear to the slow motion screw and moves the tube in declination through 15 minutes of arc in one minute of time. For the guiding motion an electromagnet in parallel with the motor brings a differential gear into action, giving a rate of motion of 30 seconds of arc in one minute of time. Similar rates of motion are provided for setting and guiding in right ascension.

Three finders are provided, one of 4½ in. aperture with eyepiece mounted on cross slides, at the lower end of the tube, one of 2½ in. and one of 2 in. aperture at the upper end.

OPTICAL PARTS

The main parabolic mirror, of pyrex glass, will have a focal length of 30 ft. The Cassegrain and Newtonian mirrors, of hard crown, are of 19 and 20 in. diameter respectively, the former being designed to give an equivalent focal length of 111 ft. (*F/18*).

The total weight of the telescope is about 50 tons. The stellar spectrograph is being constructed by Messrs. Adam Hilger, Ltd. It is of the single

prism type with 2½ in. collimator and two cameras of 12½ in. and 25 in. focus respectively.

THE DOME

The steel dome is 61 ft. outside diameter, with parallel opening 15 ft. wide extending from the horizontal to 7 ft. beyond the zenith. Two parallel moving shutters running on rails at the top and bottom of the dome close the opening and are operated simultaneously by means of wire ropes connected to a motor-operated gear. An emergency hand gear is also provided. Two motor-operated wind screens of sail cloth are mounted in the opening, one rising from the bottom, the other descending from the top.

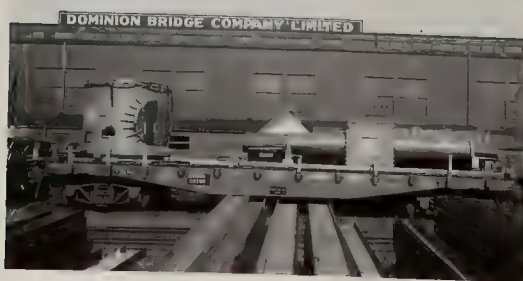
The dome, which weighs about 80 tons, is carried on 24 casted rollers of 27 in. diameter mounted in self-aligning ball bearings and running on a flat bottomed rail. Sixteen pairs of lateral roller bearings on the inner and outer edge of the rail keep the dome in position. Two segmental platforms, the lower one at the base of the opening, the upper one at the back at a 16-ft. higher level, are fixed inside the dome. Rails are bolted on the parallel straight ends of these platforms on which runs a bridge in the form of a semi arc. This bridge is 5 ft. 6 in. wide, and divided into two portions, the right hand side forming a stairway, the left being a track on which a truck carrying the Newtonian observing platform runs. The upper portion of the truck is automatically kept horizontal as it travels up the curve of the arch, by means of a lever, the outer end of which travels on a contoured rail. The truck is operated by means of a motor driven winding drum fixed at the top of the bridge, where is also fixed the gear for giving it the cross traverse motion. The rail on which the dome runs is mounted on a strong annular girder on the top of 24 steel pillars, 21 ft. high. The pillars form a circular building, sheathed inside and out with steel sheeting.

The entrance is in the south side through a steel porch with two pairs of doors. The observing floor is 13 ft. above ground level with doorway leading on to the top of the porch. From each side of the porch a stairway gives access to a gallery running right round the building 23 ft. above the ground. A similar gallery on the inside of the building gives access to the lower segmental platform in the dome and so on to the Newtonian observing platform.

A lift is provided inside the circular building for removing the cell and in an inner room when it requires to be removed. The 7½ horse-power motor and gear for rotating the dome will be fixed at ground level in the building.

Rotation is effected by means of an endless steel rope passing almost round an annular chumel ring carried on brackets fixed to the base of the dome over two tangent pulleys and down to the turning gear, a tension pulley being provided for keeping the rope tight.

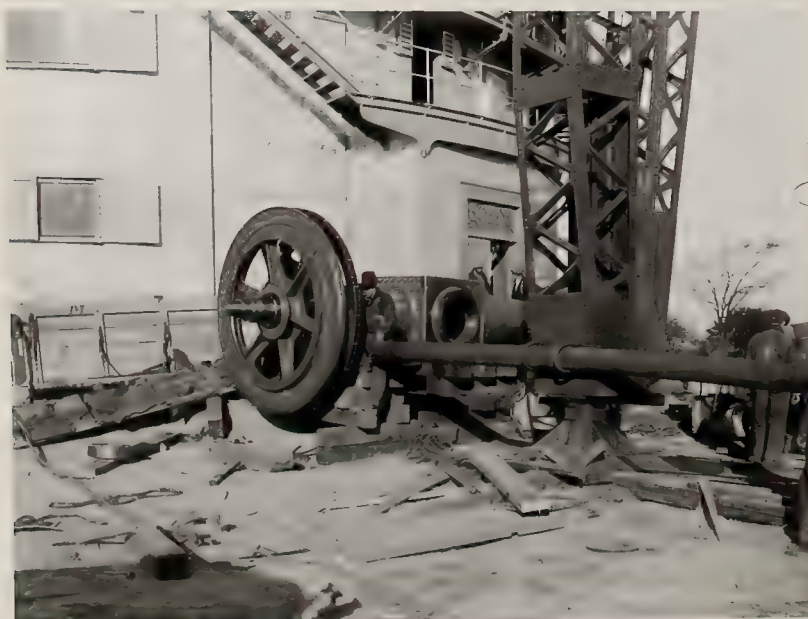
The dome and building was despatched in June and the telescope has recently been shipped.



↑ tube
↑ centre piece
Monday Oct. 16, 1933
↑ centre piece
Polar ↑ Axis
 Three views of the Telescope as it arrived from Montreal, at the Dominion Bridge Works, Shaw St. Toronto.



Oct. 16, 1933

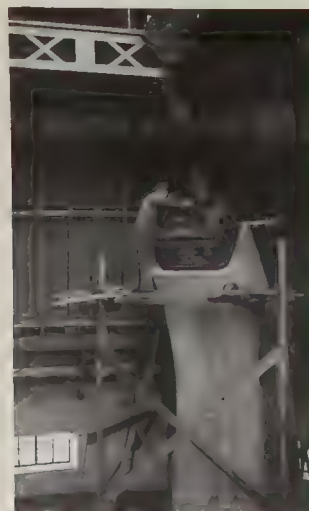


Polar Axis being assembled, at the Base of the 20-ton crane. Friday, Oct. 20, 1933

Views taken Friday, October 20, 1933



Preparing the Housing of the lower bearing on the South Pier.



Preparing the Housing of the upper bearing on the North pier.

Within the Great Dome



1. The Sidereal Circle is in place and the Driving Circle or worm wheel is being put on.



Shortreed ↑ Chant
Young ↑ Sinclair (England)
2. Putting in place the large setting gear wheel.

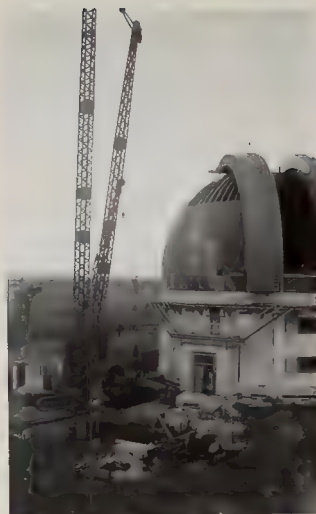
Three Views showing the Assembling of the Polar Axis



3. Putting on the Ball Bearing on which the lower end of the axis turns.



General View of Parts of the Telescope as delivered by truck to site.



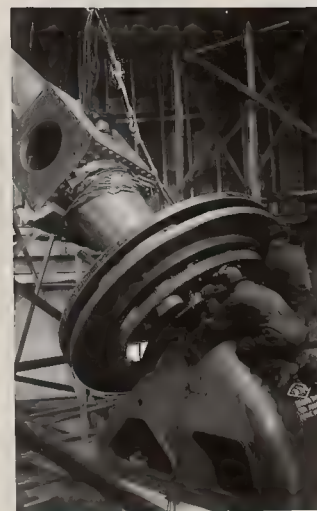
*Three Views taken on Saturday, Oct. 21, 1933, showing
Preparations for hoisting the 16-ton Polar Axis*



1. The Axis high in the Air



2. Just entering the slit in the Dome



*3. Workmen guiding the lower end
into its position*

*Three Views taken on the morning of Monday, Oct. 23, 1933
Showing the Hoisting of the Polar Axis. The whole
operation occupied less than one hour.*

*Three Views showing the Hoisting of the Declination Axis
on the afternoon of Monday Oct. 23, 1933*



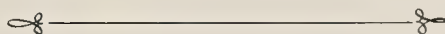
1. The Axis high in the Air



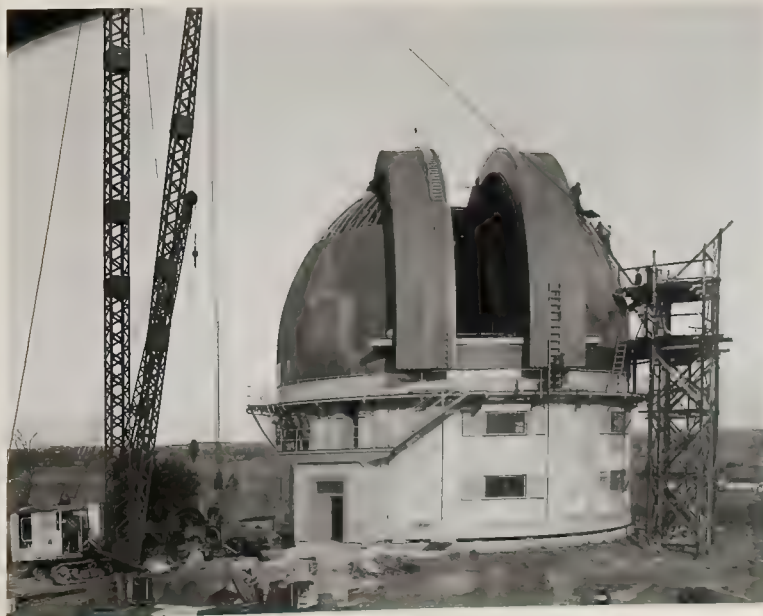
2. About to enter the Dome.



*3. The Axis being put through
the Cube of the Polar Axis.*

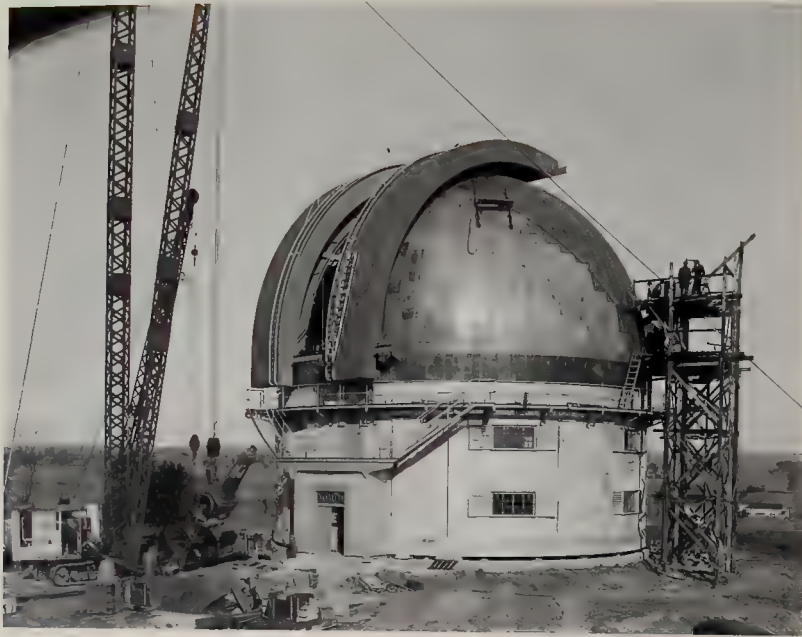


*Following are Seven Photographs of larger size, showing the Hoisting
of the Polar and Declination Axes.*

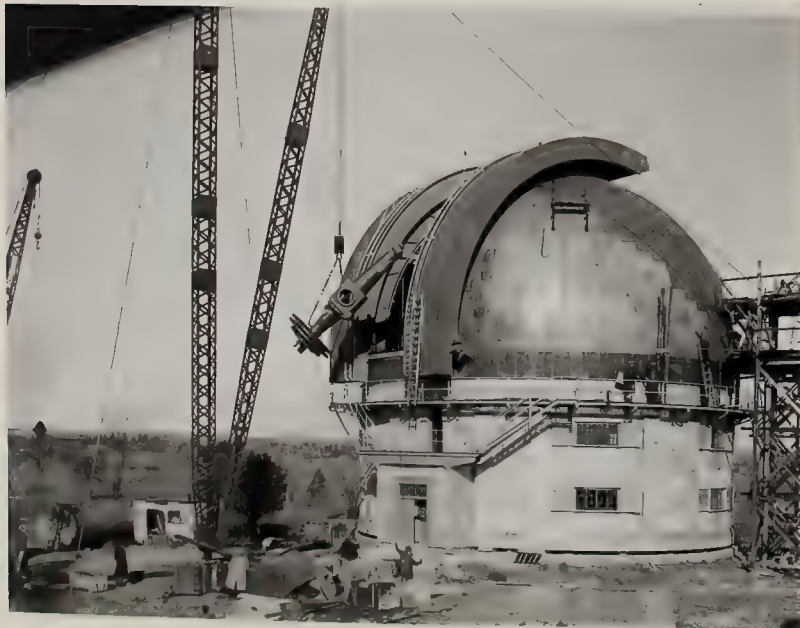


*1. The Polar Axis on the Ground. Adjusting it for Hoisting.
(The Polar Axis weighs 16~tons) Morning of October 23, 1933*

Morning of Monday, Oct. 23, 1933

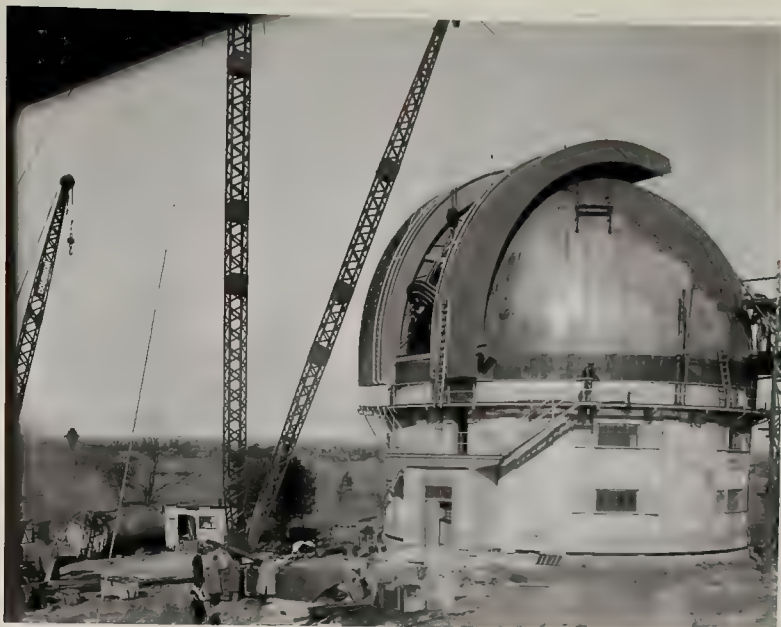


2. The Polar Axis just leaving the Ground.



3. The Polar Axis about to enter the Dome.

Monday October, 23, 1933

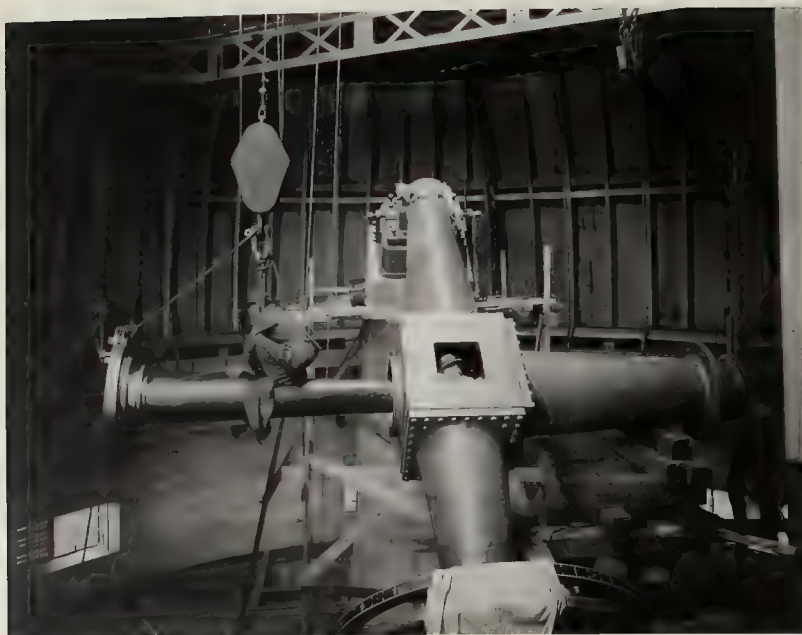


4. The Polar Axis almost within the Dome, Morning of Oct. 23, 1933

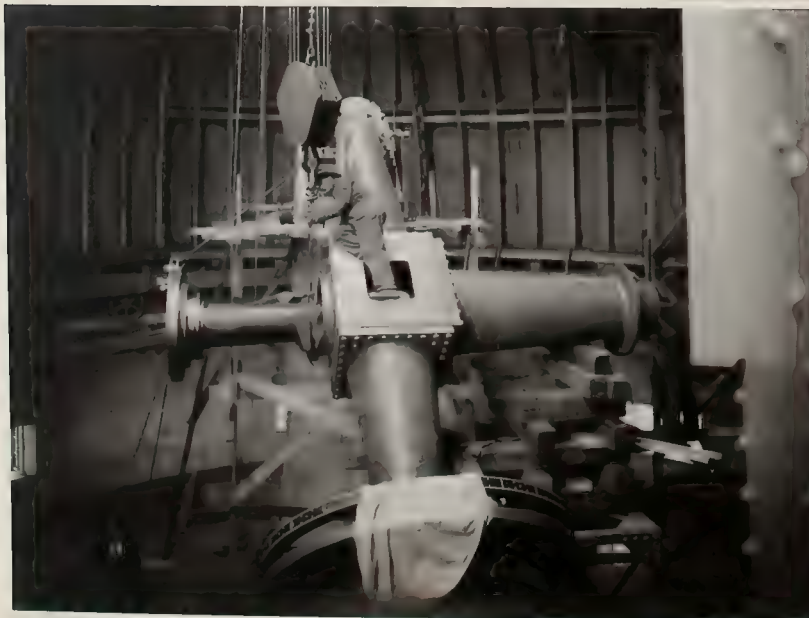


5. The Declination Axis entering the Dome. Morning of Oct. 23, 1933.

Afternoon of Monday, Oct. 23, 1933



6. Guiding the Declination Axis into position.



7. The Declination Axis moved farther into place.

Afternoon of Oct. 23, 1933



Photograph by A. Van

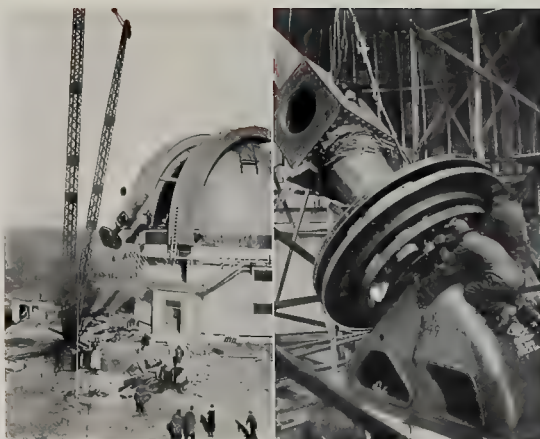


Photograph by A. Van

Two Views of the Observatory, looking Northeast, from a point just North of the lane and near the Railway. Taken at the end of October, 1933.

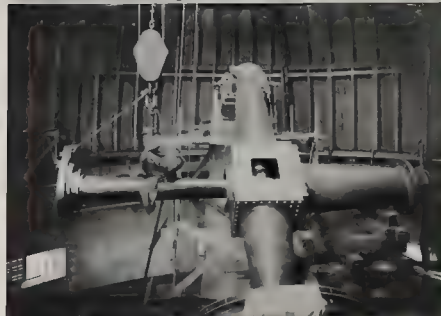
The lower picture was taken with an infra-red plate.

Journal of the Royal Astronomical Society of Canada, 1933



ERECTING THE 74-INCH TELESCOPE OF THE DAVID DUNLAP OBSERVATORY
Left—Hoisting the polar axis. *Right*—Guiding the lower end into place.

PLATE IX



ERECTING THE 74-INCH TELESCOPE
Left—Hoisting the polar axis. *Right*—Guiding this axis through the polar axis.
Journal of the Royal Astronomical Society of Canada, 1933

THE JOURNAL OF THE ROYAL ASTRONOMICAL SOCIETY OF CANADA

Vol. XXVII, No. 10 DECEMBER, 1933 Whole No. 229

THE 74-INCH TELESCOPE OF THE DAVID DUNLAP OBSERVATORY

By C. A. CHANT

WITH PLATES IX AND X

THE photographs reproduced in Plates IX and X illustrate progress in the installation of the 74-inch telescope of the David Dunlap Observatory.

The circular building with its revolving dome arrived at the end of July. It was supplied by the firm which constructed the telescope, namely, Sir Howard Grubb, Parsons & Co. of Newcastle-on-Tyne. The structural part was produced by the Cleveland Bridge Co., of Darlington, but the intricate gear for revolving the dome, moving the shutters and wind-screens, &c., was made in Newcastle. As shipped it weighed 210 short tons. It was loaded on the steamship *Ilustor* in England and brought directly to the Toronto harbour. From the dock it was taken by truck to the site fifteen miles away. The building had been erected in England and was speedily re-erected by the Dominion Bridge Co. of Toronto.

The diameter of the building is 64 feet and the walls are two feet thick. They are sheathed inside and outside with heavy galvanized iron. To bind the sheets together some 32,000 bolts were used. The revolving dome is sheathed with "agasote" a preparation of papier maché and a good heat-insulator. The outer covering is one-half inch thick and the inner one three-eighths of an inch. Over all is a covering of copper for protection from the weather. The covering of the dome was done by Sullivan and Fried, who erected the Administration Building with the domes on it.

401

402 The 74-inch Telescope of the David Dunlap Observatory

The telescope—all but the large mirror—was shipped from Newcastle by the steamer *Carnross*. At Montreal it was transferred to freight cars which brought it to Toronto. Then it was carted to the site and erected by the Dominion Bridge Co. As one piece of the telescope, the polar axis assembly weighed 16½ tons a special 20-ton crane with a boom 92 feet long was erected to handle it. First, the housings of the bearings were carefully put in place on the north and south piers, and the axis was prepared. The driving and other wheels were placed on the south end; then when the large crane had raised the axis from the ground a second crane lifted one end until the axis was inclined to the horizontal at an angle of about 14° (the latitude of the site). The lower end was tied to the hook of the large crane so that the inclination of the axis should be maintained. Then the large crane raised the axis high in the air, passed it through the 15-foot slit of the dome and gently lowered it to the pier. The whole operation from leaving the ground until in its final position occupied less than an hour.

The left picture in Plate IX shows the axis in mid-air and the right picture shows the workmen guiding the lower bearing into its place. Then the declination sleeve weighing 3 tons was raised into position and bolted to the central portion of the polar axis. This was done on the morning of October 23.

In the afternoon the declination axis, weighing 3½ tons, was raised and put in place. In the upper picture of Plate X the axis is seen passing into the dome, while the lower picture shows it being guided through the polar axis and the declination sleeve.

At the present date (November 27) the large switchboard with its relays for controlling the fifteen motors and their circuits is being erected on the ground floor. The ground and observing floors, which will be of cement, will soon be put in place.

The great disc of Pyrex glass, which was poured at the Corning Glass Works on June 21, after three months in the annealing oven was removed from its mould. By means of polarized light it was examined for internal strains and it was found to be exceptionally good. On October 5 it was shipped from Corning, N. Y., to Montreal and left that port on October 15 on the same ship which brought over the telescope. The disc is now at the optical works in Newcastle where it is being ground into shape.

DEATH OF LADY PARSONS

Lady Parsons died on October 16, aged 74 years, at her beautiful country home at Ray, about twenty miles from Newcastle-on-Tyne. She was the widow of Sir Charles Parsons, the inventor of the steam turbine, and known to astronomers as a son of the Earl of Rosse, who completed a 6-foot reflecting telescope in 1845, and also as the real head of the firm of Sir Howard Grubb, Parsons and Co., which has in recent years constructed a number of large telescopes, including the 74-inch reflector for the David Dunlap Observatory of the University of Toronto. Sir Charles died in February, 1931, and a sketch of his life appeared in this JOURNAL, volume 25, p. 185 (1931).

Lady Parsons' maiden name was Katharine Bethell, of Rose Park, Yorkshire, and she married Charles Algernon Parsons in 1883. Her intellectual ability and mechanical taste allowed her to appreciate her husband's great achievements. A son, Major A. G. Parsons, was killed in action in 1917 while serving with the Royal Artillery in France. A daughter, Miss Rachael Parsons, survives. She has the distinction of being one of the three women members of the Institute of Naval Architects.

Lady Parsons was buried beside her husband in the little village churchyard at Kirkwhelpington. The account of Sir Charles' life referred to above contained two plates, one being a photograph of Sir Charles and Lady Parsons taken at Ray by the present writer. A copy of the sketch was sent to Lady Parsons and I take the liberty of printing her reply below.

Ray Demesne
Kirkwhelpington, Northumberland
July 4th, 1931

Dear Mr. Chant,

I received your letter of June 1st, and thank you so much for sending me a copy of the Obituary you wrote on Sir Charles. I think it is an excellent account of the main features of his life and work, and you have dealt with it all in a very appreciative and sympathetic manner. I am very pleased indeed to have it.

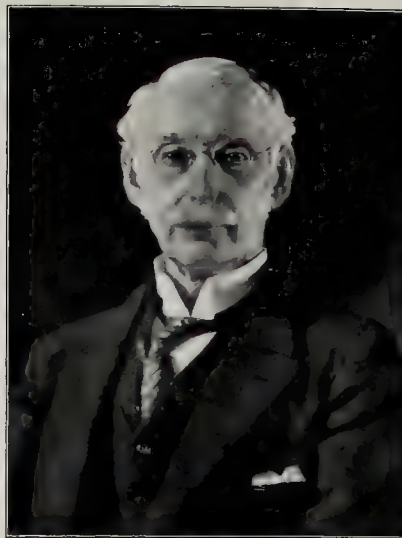
I had forgotten the photograph. It is very good in spite of the misty weather.

I hope very much that I shall meet you all again some day.

With many thanks

Sincerely yours

KATHARINE PARSONS



HIS SIR CHARLES ALGERNON PARSONS, K.C.B., F.R.S., O.M.
(Born July 13, 1854, died February 11, 1931)
(From a Portrait by Orpen)

Reprinted from the Journal of the Royal Astronomical Society of Canada
May-June, 1931

SIR CHARLES PARSONS

By C. A. CHANT

(With Plates XII, XIII)

THE recent death of Sir Charles Parsons, while on a pleasure cruise to the West Indies, came as a shock to his many friends and admirers. Though he had reached a somewhat advanced age, his activity of mind and body gave promise of further years of achievement.

Charles Algernon Parsons was born July 13, 1854, the fourth and youngest son of the third Earl of Rosse, who is famous in astronomical history as the constructor of the six-foot reflecting telescope of Birr Castle, Parsonstown, Ireland. Lord Rosse was born in 1800 and died in 1867. He graduated at Oxford University with first class honours in mathematics, and in later life he was a profound student of economics. While a youth he showed a special taste for mechanics and developed a remarkable manual dexterity.

Having leisure and means Lord Rosse looked about for the best way to employ his mechanical skill. Now Sir William Herschel died in 1822, and Lord Rosse admired exceedingly his skill in building reflecting telescopes as well as in using them. Herschel, it is said, had two hundred failures before he succeeded in making a satisfactory mirror; but ultimately he was able to produce one four feet in diameter, with a focal length of forty feet. Lord Rosse knew well that to construct such large astronomical instruments would try his mechanical ability to the limit, but he deliberately decided to make it the chief business of his life.

The material then used for telescope mirrors was an alloy of copper (two parts) and tin (one part). It is extremely hard and brittle, quite unlike either of its two constituents. It is very difficult to mould and the manipulation of the casting, after it has been made, demands experience and judgment; but it takes a fine polish and does not tarnish quickly.

After much experimenting Lord Rosse, in 1842, undertook to produce a mirror six feet in diameter and five inches thick—far larger than any attempted up to that time. He met many difficulties and had many failures but he would not give up, and the great telescope was completed in 1845. It was a scientific wonder of the world and remained unequalled in size until about fifteen years ago.*

In his father's workshop Charles Parsons learned to love mechanical pursuits. It is on record that in 1868 he and two brothers built a four horse-power steam "horseless carriage," which developed a speed of ten miles per hour; but it figured in an accident to a pedestrian and the father locked it up.

Sir Charles received his early education from the able young astronomers employed by his father to observe with the six-foot reflector—among them being Dr. G. Johnstone Stoney and Sir Robert Ball—and then, after a brief period at Dublin University, he went to Cambridge. In three years he graduated (in 1876) with high honours in mathematics (eleventh wrangler). In the same year he became an apprentice in the Armstrong ordnance works at Elswick, where he remained four years, devoting all his spare time to the study of steam. He next joined Messrs. Kitson, of Leeds, as experimentalist. This company manufactured a rotary engine which he had invented. Here he remained two years, until 1883.

Journal of the Royal Astronomical Society of Canada
Dec. 1933

PLATE XIII



SIR CHARLES AND LADY PARSONS
(Taken at Ray by the present writer)

Then he became a member of the firm of Clark, Chapman, Parsons and Co. His work on the rotary engine had directed his attention to the steam turbine. Many had attempted to solve the numerous problems involved in it and had failed, but this did not deter him—perhaps attracted him. He refrained from examining any of the previous patents, but attacked the general problem in his own way—from analogy of the water turbine. On April 23, 1884, he applied for two patents, and the history of the modern steam turbine began. The first patent covered a method of constructing a dynamo to run at the tremendous speed of 18,000 r.p.m.; the second was for a steam turbine on the same shaft, to drive it. Considering the rudimentary condition of electrical theory at that time, the construction of such a generator was an audacious project, while the turbine was entirely new. In less than a year a 4 k.w. dynamo and turbine unit was completed—the first successful turbine ever built.

During the next five years his firm constructed about three hundred turbo-generators of this type, the largest having a capacity of 75 k.w. During all this time Sir Charles laboured strenuously to improve the machines. Becoming somewhat dissatisfied, in 1889 he left the firm, but the latter, as the result of an arbitration, was allowed to retain the patents. The price asked for them was about £98,000, which Sir Charles would not give. Undaunted, in the same year he established the firm of Charles A. Parsons and Co. at Newcastle-on-Tyne, and resolutely set to work to produce a turbine on a different principle. In the original type the steam moved parallel to the axis; in the new one, radially outwards. In three years, in 1892, he produced a larger and more economical machine than the original one, though he still considered the parallel-flow method preferable. However, his old firm was not able to utilize the primary patents and Parsons got them back in 1894 for about £2,000. This gave him great satisfaction.

It was in this year Sir Charles began the application of the turbine to ships. A multitude of new and vexatious difficulties were encountered, arising chiefly from the excessive speed of rotation of the turbines. But one by one they were mastered, and in 1897, he made a grand demonstration of the use of turbines for marine propulsion. He equipped a little vessel, the *Turbinia*, one hundred feet long and nine feet wide, with engines of about 2,000 h.p., and at the naval review on the occasion of the Diamond Jubilee of Queen Victoria it raced up and down the line of warships at the unprecedented speed of nearly forty miles an hour. This performance drew universal attention to the marine turbine and its use in ships grew rapidly. In 1907 the *Mauretania*, of 40,000 tons displacement, with the turbine engines of 68,000 h.p., was completed, just ten years after the *Turbinia*, of 40 tons displacement. At present a great Cunard ship is being built which will have engines of 200,000 h.p.

As early as the nineties Lord Kelvin made the statement that the Parsons turbine was the most important development in steam engineering since the days of James Watt. The years since then have abundantly confirmed the accuracy of this remark.

Sir Charles Parsons had many other interests. One of these was the artificial production of diamonds. In various ways he subjected carbon to enormous pressures and temperatures, but without success. He would seldom give up a method without pushing it to the limit, and it is a pretty safe prediction that no one else will succeed where he failed. On these notable experiments about £30,000 was spent.

From his father Sir Charles inherited an interest in optical work. As early as 1887 he took up the production of mirrors for searchlights. As usual, he improved methods of manufacture, and devised a process for making mirrors up to seven feet in diameter, true to a desired shape. Nearly all the parabolic searchlight reflectors in Great Britain are from the Parsons works at Newcastle. He also made reflectors parabolic in one plane and hyperbolic or elliptical in the plane at right angles. Finding that there were many casualties among these mirrors he undertook an investigation into the glass from which they were made, which led to the production of a special heat-resisting boro-silicate glass, much more suitable for the purpose.

In January, 1921, he secured a large interest in the optical firm of Ross, Limited and became chairman of the Company. Here again his original mind and engineering genius suggested new methods of grinding which led to improvements in manufacture, with increased efficiency and a nearer approach to perfection in the finished products.

It is well known that at the outbreak of the great war Britain lagged far behind in the production of optical glass. This art, in which she was supreme a few decades ago, had been neglected, and the researches of Abbé and Zeiss in Germany had given that country almost a monopoly of this material. During the war the Derby Crown Glass Works had been established and it had been successful in producing glass of the highest quality; but after the war the urgency was gone and the continuance of the works was problematical. Sir Charles, of course, realized the importance of optical glass in the production of optical instruments; but no doubt his devotion to his country made him view the loss of such an industry with great misgiving. It is also believed that he was attracted by the possibility of making large lenses for telescopes. At any rate in July, 1921, he acquired the Derby Company, and its name was later changed to the Parsons Optical Glass Co. Once more his originality and experimental ability soon brought about many changes. He evolved a method of producing large discs and promptly accepted an order from Sir Howard Grubb and Sons for the flint disc for the 27-inch objective for Johannesburg. He then developed his own methods for much larger discs for lenses. He introduced improvements in the furnace in order to avoid breakage of the melting pot, he devised methods of stirring, thereby obtaining greater homogeneity; he overcame difficulties in handling the melted material and in annealing. In this way he

succeeded in producing a pair of discs for a telescope objective 42 inches in diameter—slightly larger than the lenses of the Yerkes refractor. Some work has been done on these discs in the expectation of using them in a refractor for the Russian Soviet government.

The name of Grubb is highly honoured among the makers of astronomical instruments. Thomas Grubb, born in Kilkenny, in 1800, early in the century established a shop near Dublin. Among his products was the original equipment for nearly forty magnetic stations, one of which was established in Toronto in 1840, and has been in operation ever since, though removed to Agincourt in 1898. The Earl of Rosse often consulted Grubb in the construction of his telescopes; indeed they were friends with kindred tastes. Grubb's most notable production was the famous four-foot reflector at Melbourne, Australia. Thomas Grubb retired from active work in 1868 and was succeeded by his son Howard, who extended the business, making many instruments, large and small. In 1918 the works were moved to St. Albans, England, in connection with the manufacture of periscopes, as the firm had made a specialty of those instruments. Many makers of scientific

instruments suffered in the aftermath of the war, amongst them the firm of Sir Howard Grubb and Sons. Sir Charles Parsons was interested, not only in the work of the firm, but also in its members and their family history, and it is pleasing to learn that in February, 1925, he acquired the business and the firm name became Sir Howard Grubb, Parsons and Co. Large and well-equipped shops were erected beside the turbine-electric works at Newcastle, and the firm was ready to construct astronomical equipment of the largest size. Since then some large instruments have been produced, including a 36-inch reflector for the Royal Observatory, Edinburgh; also a 40-inch reflector and a 24-inch refractor, with their revolving domes, for the new Stockholm Observatory. The order for the 74-inch reflector for the David Dunlap Observatory, to be erected near Toronto, was given in May, 1930, and it is to be regretted that Sir Charles did not live to see its completion. He was accustomed to the construction of mighty machines, some of them much larger than this instrument, but I believe he was thrilled with the prospect of producing a telescope excelling his father's masterpiece of 1845.

During his lifetime Sir Charles Parsons received many honorary degrees and medals. To enumerate them would take too much space, but three distinctions may be mentioned—F.R.S. in 1898, K.C.B. in 1911, and O.M. (Order of Merit) in 1927.

In 1883 Sir Charles married Katherine, daughter of W. B. Bethell, of Yorkshire. They had two children, a son, who was killed in 1918, in the war, and a daughter.

Sir Charles and Lady Parsons made many journeys together, including one to South Africa in 1929. Early in February last they started on a cruise to the West Indies. When some days out Sir Charles became indisposed, but nothing serious was suspected. On the 10th he seemed rather better and was able to sit on deck, smoking and chatting with his friends; but next day he became worse and quietly passed away at 8.30 p.m. on board ship, in the harbour of Kingston, Jamaica. The body was removed and taken back to England, where it was laid to rest with simple ceremony in the churchyard of Kirkwhelpington, on his estate about twenty-five miles north-west of Newcastle. Very appropriately a memorial service was also held in Westminster Abbey.

In closing my sketch of this great man I will quote a few remarks made by some who knew him much longer than I did.

DR R. T. GLAZEBROOK, in *Nature*.—

No one meeting Parsons casually would have recognized in the gentle, modest man, somewhat quiet and hesitating in speech and manner, one of the world's great benefactors. In public he said little but interest him in a problem, ask his advice on some knotty point of scientific or engineering practice, give him, perhaps, a little time for quiet thought, and your problem was solved, or if solution was not at once to be found, you were set on a track promising to lead to the desired end. Moreover, if the attainment of that end seemed of importance you secured for the rest of your journey the support and assistance of a most wise counsellor and, what is more, a most kind friend.

J. HAZELKES, Managing Director, Ross, Limited. —

He was unusual in his ability to lay his fingers on the weak spot of any contrivance which had been much thought of by its originator, and here his great charm of manner was in evidence, so that no man hesitated to lay his ideas before him, and all were always the richer for the advice and help unstintingly given. Sir Charles had ready almost at once not one solution only, but several, leading to embarrassment as to which of them to choose.

A SCIENTIFIC FRIEND in *The Yorkshire Post*. —

No one can ever know the range of his help to those who had fallen on hard days, but it was inevitable that from time to time instances should become known. Most striking of all in this connection was his solicitude for those who in pain or sorrow were nearing their end. He would spare no trouble to cheer their days in some thoughtful way. On receiving the Order of Merit he wrote, "I think the congratulations of one's old friends give far more pleasure than any honour" no formal phrase, but sterling truth from him.

University of Toronto

CORRECTIONS TO SKETCH OF SIR CHARLES PARSONS

The present writer wishes to make three small corrections to his sketch of Sir Charles Parsons which appeared earlier in this volume (p. 185). He tried to be accurate in every detail but was not able to refer to original sources and had to depend on authorities which were not absolutely without error.

Sir Charles was the sixth son of the third Earl of Rosse. The second and third sons died in early boyhood, and of the four who grew to manhood he was the youngest. He left Cambridge University in 1877. In referring to the steam carriage constructed by Charles and his next elder brother it was stated that there was an accident to a pedestrian. This was in 1869. As a matter of fact, Lady Bangor, a cousin of the family, who was riding on it, fell off and was instantly killed.

The above corrections are made from an admirable obituary notice in the *Proceedings of the Royal Society*, prepared by "J.A.E." (Sir James Ewing), who is thoroughly acquainted with the scientific work of Parsons and was a personal friend since 1891.

C.A.C.

The Dunlap Observatory

DURING the past year we have casually noticed the new Dunlap Observatory in the course of construction just south of Richmond Hill.

We have been particularly interested in its advancement owing to the fact that our Headmaster laid the corner-stone and because its donor, Mrs. D. A. Dunlap, has always been so sincere a friend to St. Andrew's College.

The buildings are advantageously situated on a hill-top from where they command an unequalled view over the surrounding country, and are easily accessible from the highway.

It was a bleak day I had chosen for my expedition and as I struck from a little road across to the buildings, the wind that swept unchecked over the open country tugged at my coat and almost buffeted me from my feet. After hastily snapping a picture, I made for the huge dome which I knew contained the gigantic telescope. It is enclosed by strips of shining copper broken only by the telescope-opening running from the centre to the circular walls, which are wholly of steel for the purpose of grounding any charges received in an electrical storm. These greyish walls, relieved of their monotony by a few large windows, are encircled by a small platform, upon which a heated student might stroll if fatigued with his enthusiastic observing.

Hearing the hammering and riveting from within I eagerly entered the building.

The greatness of everything amazed me—perhaps it was the modernness of it all, or was it the men working high up on the huge metal tubing above? I did not know.

Immediately a man approached me; it was plain to tell from his broad accent that he came from the north of England. I found out later that he had been especially sent by the manufacturers of the telescope to superintend its assembly. I told him my reason for being there and he kindly offered to show me round.

I shall exclude the first few blundering questions I fired at him and let you profit by my extreme embarrassment and discomfort.

The telescope was built at Newcastle-on-Tyne, England, and has been assembled by a Canadian bridging concern. The size of a reflecting telescope depends entirely upon the dimensions of the main mirror which is enclosed in the bottom of the great machine.

This particular mirror, for which a special mould had to be built, is still in the process of setting; when completed it will weigh two and a

half tons and be seventy-four inches in diameter. When the telescope itself is finished it will weigh fifty tons and rank as the second largest in the world.

There are no lenses in a reflecting telescope, although some of us supposed there were; the actual observing is done, as the name implies, through the reflection of large mirrors. In the top are two of these; the



Newtonian for photographing, and the Casagrain, which would replace it for spectrographing (that is examining the different colours of light).

The telescope is counterbalanced by large drums which secure the most amazing effect of all, namely, the ease with which it is controlled. On the mere pressing of a button a powerful motor turns at one hundred and ten revolutions per minute and the telescope is quickly swung into any required position within the dome.

"What is that steel framework for?" I asked, proud of my discovery.

"When the big mirror is installed," explained the superintendent, "it has to be re-silvered about twice a year. That is an electric lift used for the removing of the glass, which otherwise could not be done without a great deal of trouble."

We climbed a narrow flight of steps to a steel walk encircling the interior of the dome. From here I could see easily a long sort of tower swung out over us. This was the observer's platform equipped with one small wheel which could swing the big girder in the desired direction. The actual observing is done from here after it has been "sighted" by several smaller finders secured to its sides.

From the little platform high in the air I obtained a splendid view of both the unfinished floors beneath. In all the observatories the domes are revolving, this one on large wheels running with a track, is driven by an engine installed on the ground floor.

On descending to the little walk again, we passed engines for closing the massive dome.

It is interesting to know that the temperature remains always the same as outside, because of the expansion and contraction of the glass.

Becoming confused with so much technical knowledge, after a vague hint, my guide led me across to the Administration building in which the business and studies of the Observatory are conducted.

The building is of stone and most imposing with its three copper domes which lend to it the air of a Persian temple. The lawn in front, I was told, is to be graded and planted with shrubs and flowers.

The renovated home of Dr. Chant, of the department of Astrophysics at the University of Toronto, is conveniently situated near the highway and within a few minutes walk from the Observatory.

On entering the building I was struck with its brightness. From the long hall branch the offices and class-rooms; the latter accommodate fifty students.

The front hallway is beautifully finished in Italian marble and upon the floor are engraved the points of a true, coloured compass.

The inscription, to the memory of the late David Alexander Dunlap, on highly polished marble with gold engraving, catches the eye immediately upon entering.

A magnificent stair leads up to the second floor, and a smaller flight of steps to the three domes surmounting the building. Here the students are to do their own observing through the three telescopes, a large refractor, a small reflector and a celostat.

Having become so engrossed in my surroundings, I had not noticed how quickly the time had passed, and now it was growing dark. The

wind had stopped blowing—the clouds had been driven from the sky. As I opened the door, the last faint glow of the red sun fell upon the inscription:

"THIS OBSERVATORY
WAS PRESENTED TO THE UNIVERSITY OF TORONTO
BY
JESSIE DONALDA DUNLAP
AS A MEMORIAL
TO HER HUSBAND
DAVID ALEXANDER DUNLAP
1933."

THOMPSON,
Form V.

Nephew of
Ernest Thompson
Seton.



The Farm House and its Transformation.



Woodshed Kitchen

March 1, 1933

View from North.



View from South.

April 27, 1933



March 1, 1933
View from North-west



March 1, 1933
View from West



April 27, 1933
View from South-west



Brick oven
Tearing down Kitchen. Woodshed already removed

May 29, 1933

The Farm House being changed



*July 11, 1933
View from North*



*August 16, 1933
View from South-east*



View from South-west August 16, 1933



*Janitors Apt. ↑ Sept. 7, 1933
View from South-east*



*The Director's Residence
"Observatory House"*

October 22, 1934



October 24th
1932

Mrs. D. A. Dunlap,
93 Highlands Avenue,
TORONTO, Ontario.

Dear Mrs. Dunlap :

We are sending to you under separate cover, four airplane views of the property near Richmond Hill, showing the building of the new observatory from different angles.

We also enclose two views taken from ground level, in September, and suggest that the next picture might now be taken. Will you kindly let us know if you wish this done.

Yours very truly,

ROUS AND MANN, LIMITED,

John D. McChaul

JDM:BJ
ENCL.

ADDRESS ALL COMMUNICATIONS TO THE FIRM WITH ATTENTION OF MR. J.D. McChaul



ROUS and MANN Limited

PRINTED ADVERTISING • 172 SIMCOE ST., TORONTO, CAN.

November 26th
1934

Prof. C. A. Chant,
Dept. of Astronomy,
University of Toronto.

Dear Sir :

The air pictures of the Observatory, about which you enquired, were taken on October 2nd, 1932.

Yours truly,

ROUS AND MANN LIMITED,

John D. McChaul

JDM:BJ

ADDRESS ALL COMMUNICATIONS TO THE FIRM WITH THE ATTENTION OF THE INTERESTED PERSON



ROUS and MANN Limited

PRINTED ADVERTISING • 172 SIMCOE ST., TORONTO, CAN.

December 24th
1934

Mrs. D. A. Dunlap,
93 Highlands Avenue,
TORONTO, Ontario.

Dear Mrs. Dunlap :

I am sending you herewith the mounted sheets for the Great Record book. The titling is not completed as we have been busy dry-mounting the pictures in, but it will give some idea of the scope of the book.

As soon as you are finished with these sheets I will get them back and finish the lettering. I understand that Professor Chant has still some pages to come.

Wishing you the Compliments of the Season, I am,

Yours very truly,

Arthur H. Hobson

AHR:BJ
ENCL.

A. H. Hobson

ADDRESS ALL COMMUNICATIONS TO THE FIRM WITH THE ATTENTION OF THE INTERESTED PERSON



The Hon. Lady Parsons

a concave mirror placed at the back end of the tube. The image formed in each case is viewed through a magnifying lens, or combination of lenses, called the eyepiece.

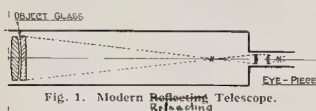


Fig. 1. Modern Reflecting Telescope.

Taking the Refractor first (Fig. 1), of which the ordinary "held glass" is an example, the usual form of achromatic object glass consists of two lenses, the outer component a double convex lens of crown glass, and the inner lens of flint glass, the surface adjacent to the crown lens being concave, while the fourth surface is generally a convex of long radius. All four surfaces have spherical curves. It is essential that the optical glass, from which the lenses are made, should be of the highest quality, free from veins and striae, homogeneous or of equal density throughout and perfectly annealed.

As these conditions are very difficult of attainment in large pieces of glass, there appears to be a fairly definite limit to the size of the refracting telescope, the largest so far constructed being that at the Lick Observatory, U.S.A., having an objective 36 inches in diameter, built in 1888, and the Yerkes telescope of the Chicago University, completed in 1897, which has an objective 40 inches in diameter. The production of a satisfactory disc of crown glass for the 36-inch Lick objective was only attained after nineteen failures.

Another reason which places a limit to the useful size of the refractor is that the larger the lens the thicker it is necessary to make the glass, and a stage is reached where any gain in light due to the increased diameter is counterbalanced by the loss to transmission through the greater thickness of material.

Now, turning to the Reflecting Telescope, the main mirror which collects the light must necessarily be of parabolic section, in order that all the rays from a star which to all intents and purposes are parallel, should be brought to a common or prime focus. This focus is formed away at the mouth of the tube, in the bottom of which the main mirror is fixed and as

obviously the image formed could not be viewed there, it is necessary to introduce a second mirror in the cone of rays between the main mirror and the prime focus, in order to divert the image to a more convenient position. Various ways of doing this have been devised; thus, in the Gregorian telescope (Fig. 2) the

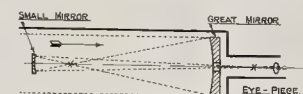


Fig. 2. Gregory's Reflecting Telescope.

main mirror has a hole in its centre, in which is fitted the eyepiece, which receives the rays from the main mirror after their projection on to a smaller mirror placed in an inverted position in the centre of the telescope tube. In the Herschelian reflector (Fig. 3) the main

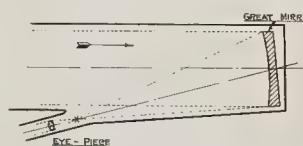


Fig. 3. Herschel's Reflecting Telescope.

reflecting mirror is tilted, causing the reflected rays to be projected to one side of the tube and near its mouth, at which point an eyepiece is inserted. Other forms are the Newtonian and Cassegrain, which are the only types now constructed.

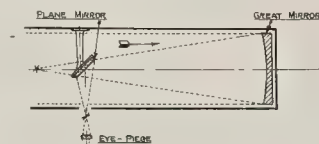


Fig. 4. Newton's Reflecting Telescope.

In the Newtonian form (Fig. 4), devised by Newton about 1668, a plane mirror is fixed in the cone of rays at an angle of 45 degrees to

ON TELESCOPES

By C. Young, F.R.A.S.

MANY of us have paid our penny to have a look through a telescope at the seaside just to read the name of a ship, and some of us are the happy possessors of a pair of field glasses.

There is a certain fascination in being able to see more than one can with the unaided eye.

A telescope does not magnify in the sense that a microscope does, it does not make an object appear larger than it actually is. What it does is to bring it apparently nearer. The pupil of the human eye is about $\frac{1}{4}$ th of an inch in diameter and can separate or distinguish between two adjacent objects whose angular distance apart is about 30 seconds of arc. With

the aid of a 6-inch diameter telescope, the angular distance is reduced to one second, with a 12-inch to a $\frac{1}{2}$ second and so on.

The late Sir David Gill, an Aberdonian, and H.M. Astronomer at the Cape of Good Hope Observatory, was once asked what he could see with his big telescope, a 24-inch refractor, and he replied that he would be able to pick out a threepenny piece if it was 84 miles away.

There are two distinct classes of telescopes; the Refractor, in which light is gathered together into a focus by refraction, or bending, through a lens or object glass placed at the front end of the tube, and the Reflector, in which the same end is attained by reflection from the surface of

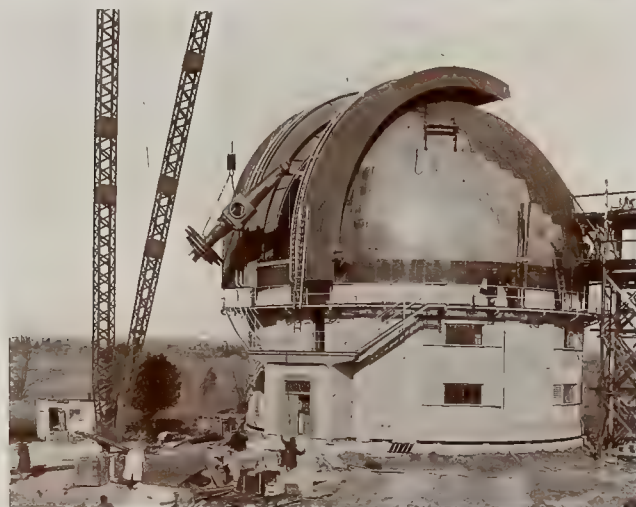


Fig. 6.—The 12 ton polar axis of the 74 inch Reflecting Telescope for Toronto University being hoisted into the 61 ft. dome on Richmond Hill, Toronto.

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the optical axis, in such a position that the focus is formed a few inches outside the periphery of the tube without any alteration of focal length. In the Cassegrain (Fig. 5), a form suggested by Cassegrain, a French sculptor, in



Fig. 5.—Cassegrain's Reflecting Telescope.

1672, the 45 degrees plane mirror is replaced by a mirror with a convex surface of hyperbolic section, coaxial with the main mirror, so that the cone of rays is reflected back and the focus formed a foot or so behind the main mirror, which in this case is provided with a central hole as in the Gregorian telescope.

This has the effect of increasing the focal length of the main mirror, usually by three or four times, according to the position and curvature of the second mirror, with a corresponding increase in magnification. Most of the existing large reflecting telescopes can be used either as Newtonians or as Cassegrains.

The limit to the size of the reflecting telescope, has, so far, been the difficulty in making a large enough disc of suitable material for the main mirror. All astronomical mirrors are silvered on the front surface.

The parabolic mirror must not only be extremely accurate and symmetrical in form, but it must remain so in whatever position it is placed. For this reason it must not only be made thick, usually about one-sixth of its

diameter, but must be evenly supported at the back, usually by pads mounted on a system of triangles.

Previous to about 1870, astronomical mirrors were made of speculum metal. As this metal only reflects about 60 per cent. of the incident white light, the loss of light due to the two mirrors was considerable.

The last big telescope with a speculum metal mirror was the 48-inch reflector built in 1867 by Sir Howard Grubb's father, Thomas Grubb, for the Melbourne Observatory, Australia. Since that time all large astronomical mirrors have been made of glass with a thin film of silver deposited on the front surface. Glass can be worked much more accurately than speculum metal, and a freshly silvered surface will reflect about 95 per cent. of the incident light. It is usually necessary to resilver the surface two or three times a year owing to the loss of light due to tarnishing of the silver film. This film is only about 1,200,000 of an inch in thickness and will not stand repolishing many times.

The largest instrument of this type so far built is the 100-inch reflector at Mount Wilson Observatory, California, and the second largest, the 72-inch reflector of the Dominion Observatory, Victoria British Columbia.

Both of these instruments have been completed since the war. The 74-inch reflector (Fig. 6), recently built at the Optical Works, for Toronto University will take second place, and Canada will thus possess two of the largest existing telescopes.

Mount Wilson Observatory, California, is planning to build a 200 inch reflector, but it is likely to be several years before this project comes to fruition.

SIR HOWARD GRUBB, PARSONS & CO.

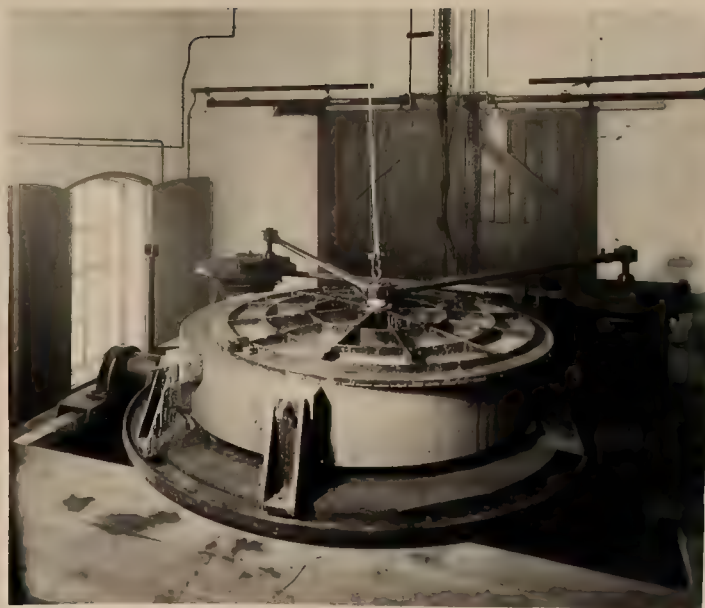
Since our last issue, the 61 ft. Dome and 74-inch Reflecting Telescope Mounting have been shipped to Toronto and are now being erected on site.

The Prime Mirror Disc for this Telescope has recently been received from the Corning Glass Works. The grinding, polishing and figuring of this mirror will probably take 12 months.

The 18-inch Reflector for the Mill's Observatory, Dundee, and the 30-inch Reflector for the Royal Observatory, Greenwich, have been delivered and erected.

The shops are at present engaged on—
(1) Two Refractor Telescopes for Victoria University
(2) Solar Telescope for Oxford University Observatory
(3) Mirror Reflector for 24 ft. Wind Tunnel for D.R.O.
(4) Royal Astronomical Society, London, etc.

Two Views of the Main Mirror



No 996.

Grinding the Mirror. The disc rotates and the Tool moves over it. Newcastle, Dec. 1933



No. 998

The grinding Tool raised; the rough grinding of this surface completed. The disc is $76\frac{3}{4}$ in. in diam., 12 inches thick. Newcastle, Dec. 1933

The Reflecting Telescope for the David Dunlap Observatory

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Paper to be presented at the General Professional Meeting of The Engineering Institute of Canada, Montreal, Que.
February 8th and 9th, 1934.

SUMMARY.—After explaining the principles on which reflecting telescopes are based, the author mentions the preliminary considerations which led to the placing of the contract for this large telescope. The general arrangement of the instrument and the various parts of its mounting and accessories are described, including the tube, the declination axis, the polar axis and driving gear and the driving clock. A brief description of the building and dome follows.

The principle of the telescope was discovered about the year 1608. The first famous telescope of Galileo was one and five-eighths inches in diameter and about twenty inches long. At that time the making of a telescope would not have been considered an engineering problem. The principles of the telescope are the same now as then but the increase in size from an inch or so to several feet makes all the difference. There is probably no instrument which necessitates the combination of technical and engineering skill to such an extent if the final product is to be satisfactory. A description of the 74-inch telescope of the David Dunlap Observatory, Toronto, will be of interest to members of The Institute.

The principles upon which a reflecting telescope is based are illustrated in Figs. 2a and 2b. In a the Newtonian form is shown, so-called after Sir Isaac Newton, its inventor. The light from a distant source is imagined coming from the right in a sensibly parallel beam which strikes the mirror A. The front surface of A is made hollow, being a paraboloid of revolution. This type of surface reflects the rays of all colours to the same point D, called the prime focus. In the Newtonian type of telescope, a plane mirror is placed at B in the path of the converging beam and inclined at 45 degrees to the axis, so that the beam is bent through a right angle and the focus is near C, where it may be examined with a magnifying glass or eyepiece.

In b we have the Cassegrain form of the telescope. The light which is reflected from the silvered glass mirror A in this case falls on a convex mirror B before it reaches the focus, and if the convex mirror is a portion of an hyperboloid of revolution, all the rays will be reflected back to a point near C. It is necessary in this case to have a hole in the centre of the mirror A to allow the light to reach the focus C.

These schemes look so simple that it is not at first apparent why it should be a difficult problem to make a telescope several feet in diameter. In the first place, however, the mirror A must be truly parabolic in shape with departures of the order of one-millionth of an inch. It must be very rigid to resist flexure and consequently heavy. The best material from which to make a mirror yet discovered is glass. In the instrument now described the mirror will consist of a solid disc about thirteen inches thick weighing about 5,000 pounds and made of pyrex. The advantages of pyrex are that it has a small temperature coefficient of thermal expansion and is a stable glass. The disc has been cast at the Corning Glass Works Corning, N.Y., and is now in England at the optical shops of Sir Howard Grubb, Parsons and Company, at Newcastle-on-Tyne, where it will be ground and finished.

In the second place the telescope must be capable of being pointed towards any object without loss of time and must be driven by some mechanism to follow the object continuously as it moves across the sky. It must not only keep the object in the field of view but must hold the image stationary. A drift of the image of a thousandth of an inch during the course of an exposure when the telescope is being used for photography would seriously affect the definition of the picture. The engineering difficulties confronting the telescope maker in thus keeping the tube so accurately pointed are about equivalent to keeping a gun pointed on a target an inch in diameter at a range of twenty miles while the target is moving at the rate of five feet per second.

The mechanical parts which support the mirrors are usually spoken of as the "mounting." The whole instrument must be sheltered by a building which can be opened to permit a view in any direction and provide means of convenient access for the observer to any part of the telescope.

The positions of objects in the sky are located with reference to imaginary circles drawn among the stars very similar to the circles of longitude and latitude on the surface of the earth. They are however given different names when drawn in the sky. Those which correspond to the circles of latitude on the earth are termed declination circles and that declination circle corresponding to the equator on the earth is termed the celestial equator. The angular distance of a star north or south of the celestial equator is called declination. The circles corresponding to longitude circles on the earth are termed right ascension circles and the corresponding angular co-ordinate is called right ascension. Right ascension is numbered from the zero right ascension circle and increases toward the east. Thus the position of any star in the sky is designated by its declination and right ascension in a manner very similar to the location of points in the earth's surface by latitude and longitude. Due to the rotation of the earth the sky appears to rotate about an axis, so that there are two points in the sky at which the earth's axis produced cuts the celestial sphere, and about which the stars seem to revolve. Of course, in the northern hemisphere only the north celestial pole can be seen and indeed its elevation above our horizon will be equal to the observer's latitude. Due to the rotation of the earth also the right ascension circles drawn among the stars come in turn to the meridian and that right ascension circle which coincides with the meridian is called the sidereal time. If any star on its right ascension circle is not on the meridian it is said to have an hour angle measured east or west in time from the meridian, the sky apparently rotating 15 degrees per hour.

The technique of telescope making has evolved gradually. Experience has had to point the way for advances and, as in other lines of engineering production, the changes are suggested by the weaknesses of former models. There are comparatively few firms that have sufficiently large machinery to handle the massive castings which support the optical parts, and, when we combine with this the fact that the number of firms which have had experience in telescope building is very small also, we see that the project could be tendered for by a very restricted few. In 1927, when Mrs. D. A. Dunlap expressed her willingness to provide the means to construct a large telescope, tentative specifications were drawn up and in June of 1928 sent to four firms, the Carl Zeiss Company of Germany, the Sir Howard Grubb, Parsons and Company of England, the Warner and Swasey Company of Cleveland, and J. W. Fecker of Pittsburgh. The preliminary specifications stipulated the general form of the mounting but left

considerable latitude in detail. The Warner and Swasey Company did not submit any tender and the design of the Carl Zeiss firm was not attractive as it was of a very radical nature. There was not much difference in the designs of the other two firms but after due consideration it was decided to accept the tender of the Sir Howard Grubb, Parsons and Company, England. This was a very fortunate choice because the decrease of the pound sterling

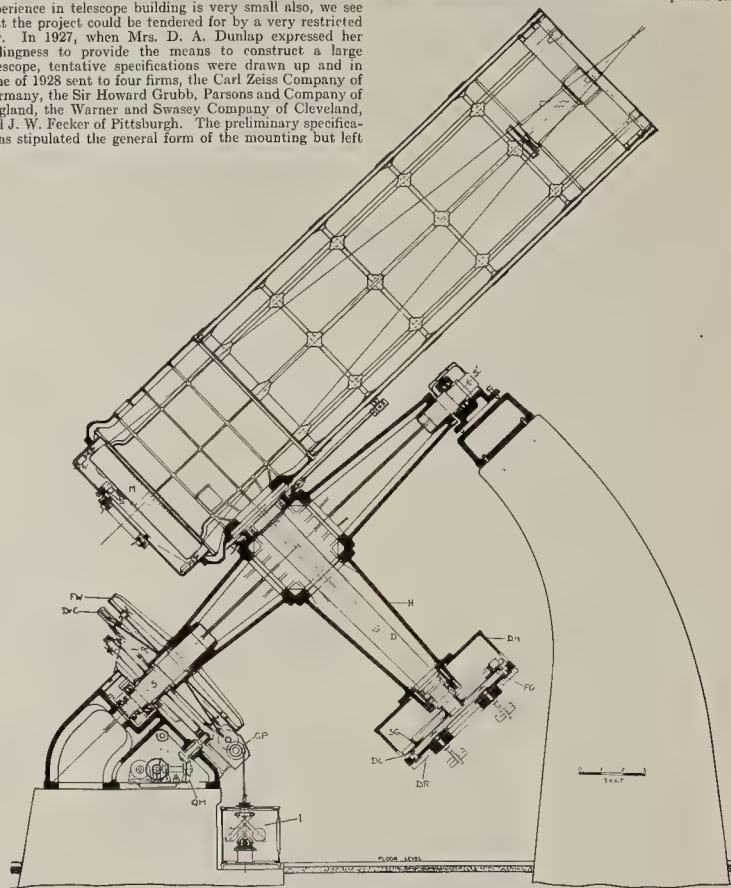


Fig. 1—General Arrangement of 74-inch Reflecting Telescope.

and the inflation of the American dollar later made the cost of the telescope very much less than it would have been had the contract been let in the United States.

The general form of the mounting is shown in Fig. 1 and a photograph as it appeared in the workshops prior to shipment is shown in Fig. 3. The essential features of any mounting are the tube which supports the optical parts and two axes of rotation at right angles to each other. The

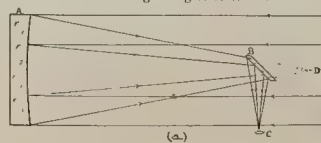


Fig. 2—Principle of the Reflecting Telescope.

polar axis (SS' in Fig. 1) must point to the celestial pole and the declination axis D to which the tube is attached provides another degree of freedom so that the tube may be pointed to any part of the sky. The polar axis is rotated by clockwork I in such a way that the telescope will continue to follow an object as it moves from east to west across the sky in its diurnal rotation. This is the equatorial form of mounting.

There are many forms of the equatorial mounting, which differ in the way the tube is attached to the declination axis, the latter attached to the polar axis, and the way in which the polar axis is supported. The kind of mounting selected for the 74-inch telescope is usually called the English type. As can be seen from Fig. 1, the polar axis is supported at both ends by two piers and the tube is carried on one end of the declination axis. In very heavy telescopes the polar axis is sometimes built in the shape of a large O and the declination axis supported by bearings at each end of the short diameter. This design is adopted in the mounting for the 100-inch telescope at Mount Wilson, California. In other telescopes the declination axis is attached to the polar axis beyond the two bearings of the latter. The 40-inch telescope of the Yerkes Observatory, Chicago, is an example of this type of mounting. In other telescopes again the polar axis is built in the form of a massive fork with two prongs and the tube of the telescope is mounted between the arms of the fork as in the 60-inch telescope of the Mount Wilson Observatory. Each type has its peculiar advantages in regard to strength, accessibility, and freedom of motion. The form adopted for Toronto has been very successful in the 72-inch telescope at Victoria, B.C.

When the contract was awarded in 1930 the main outlines only of the instrument were specified and Sir Howard Grubb, Parsons and Company began work on the detailed drawings. It seemed advisable toward the spring of 1930 for the author to spend some time in England going over the plans of the instrument with the manufacturer, and F. Jno. Bell, M.E.I.C., the Canadian representative of the firm, arranged to have the author visit Newcastle while the essential features were being drafted. This enabled much more rapid progress to be made than would have been possible if all the details had to be settled by correspondence.

On the author's arrival in England early in July 1930 the drawings were found well underway. Fortunately the firm had recently completed the 36-inch reflecting telescope for the Edinburgh Observatory and were engaged on a 40-inch telescope for the Stockholm Observatory. As there are many parts very similar in all telescopes, it was easy, from the parts of the 40-inch, many of which were in the workshop, to form a picture of the proposed new telescope. The completed drawings of the Edinburgh telescope were also a great help. In addition, the author had had a long experience with the 72-inch reflecting telescope at Victoria, B.C., and since the proposed mounting was very similar to this, conferences with Mr. Young, the manager of the optical works, enabled rapid progress to be made on the final design. When the author left towards the end of August no work on the actual construction had been done but the plans were so far advanced that construction was started in the fall.

THE TUBE

The tube of the telescope which immediately supports all the optical parts is made in three sections. These three sections are called the mirror cell, the centre piece, and the skeleton, and are shown in Fig. 1. There are several features that the tube must possess. It must be as light as possible but must carry the 5,000-pound mirror at the bottom and the small mirrors at the top with a minimum amount of flexure. It must permit a free circulation of air so that light rays, in their passage through the tube to the

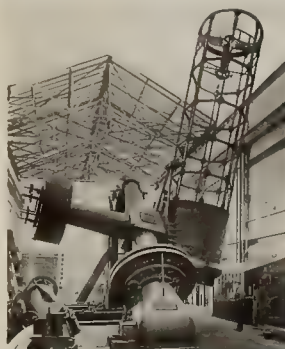


Fig. 3—Telescope with Tube Nearly Vertical.

main mirror and back, travel in a homogeneous medium. It must be capable of adjustment for alignment so that the optic axis is truly at right angles to the face to which the declination axis is attached. In order to reduce vibration to a minimum it should be mounted as close to the polar axis as possible.

The mirror cell consists of a steel-ribbed casting, weighing about 2,100 pounds. The mirror is supported on the bottom by a system of discs 7 inches in diameter in three sets of three each. These supports are shown in Fig. 1 at S. Each disc is mounted on a universal joint and the web-work which supports each set of three discs can be moved in or out to set the mirror with its face perpendicular to the optic axis. The disposition of the various discs is such that each carries its share of weight. Very careful attention was also given to the edge support. It is



Fig. 4—Truing Faces of Centre Piece of Tube.

surrounded by arcs of a strong steel rim which fits into a groove in the disc. Heavy weights shown at W are attached to this rim by universal joints and are also attached by brackets attached to the interior wall of the mirror cell.

When the tube is vertical there is no thrust on the edge of the mirror but as the tube assumes a horizontal position the weights act as levers with the brackets on the mirror cell acting as fulcrums and support the resolved component of the weight of the mirror so that in all positions the mirror is, as it were, floated. This might seem an elaborate system of support but the mirror is so sensitive to any distortional stresses that there is danger of the disc being warped if the stresses are not as uniform as possible in all parts. The mirror cell is bolted by a flange to the central casting.

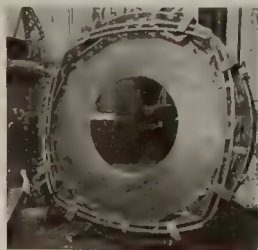


Fig. 5—Iris Diaphragm at Half Aperture.

The centre piece of the tube is also of steel and is heavily ribbed, weighing 4½ tons. A boss is cast on one side where it bolts to the declination axis. Instead of making a perfectly flat contact between the flange on the declination axis and the centre piece, they touch along a rim only. The purpose of this precaution is to reduce the rate of flow of heat from the tube to the declination axis and thence to the rest of the massive parts of the telescope.

This flow of heat has a tendency to take place in the early evening when the telescope is first put into use and the result seems to be that the part of the tube next the polar axis becomes a little cooler or warmer than the outer parts. There is then a temperature gradient across the disc and this produces a slight cylindrical warping and consequent astigmatism in the reflecting surface. It is necessary that the face of the boss and the two faces, where the mirror cell is bolted on at the lower end and the tube at the upper, be at right angles. The operation of truing these up may be seen in Fig. 4.

Immediately above the mirror and attached to the centre piece is an iris diaphragm similar in construction to that used in ordinary cameras. It is operated by a graduated hand-wheel on the outside of the tube and may be used to stop down the telescope and serves also as a protection to the mirror when closed. It is shown at half aperture in Fig. 5.

The skeleton part of the tube, which is bolted to the upper flange of the centre piece, is made of duralumin I-beams, cross-braced by rods also of duralumin threaded right and left hand at the ends into T-anchors. By means of these brace rods, which can be seen in Fig. 3, the tube may be made to twist slightly or to shift in any direction without twisting. The rods are tightened to a point where each is under tension in any position of the tube. Careful

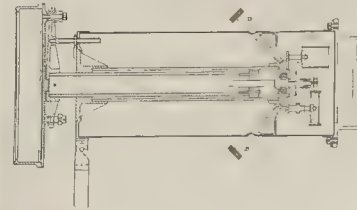


Fig. 6—Cassegrain Mirror Mounting.

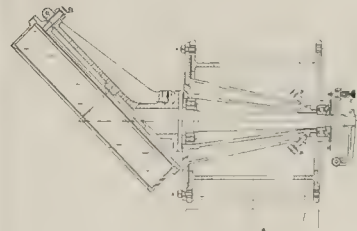


Fig. 7—Newtonian Mirror Mounting.

tests when completed have shown that, under the extreme position of flexure, the outer end does not sag more than one-sixteenth of an inch. The skeleton part of the tube supports at its upper end the secondary mirrors which may be of the Newtonian or Cassegrain form. The 74-inch telescope is designed to be convertible to either form. Figure 3 shows the Cassegrain mirror in place and in Fig. 1 the position the Newtonian mirror occupies is shown by dotted lines. Sectional elevations of these two attachments are shown in Figs 6 and 7. Very convenient arrangement is provided in both these mirrors for bringing them into alignment with the axis of the tube by means of the screws BB. By shifting the dovetail pin C (Fig. 7) in the Newtonian form the mirror can be rotated and the rays of light brought to the side of the tube in any one of four positions. This is almost essential because, in working at the upper end, it is convenient to bring the focus to that side of the tube which is the easiest to reach. By loosening the screws AA the Newtonian may be removed and the Cassegrain fastens to the same flange. When the Cassegrain mirror is in place the observer is situated at the base of the tube for observation. In a large telescope of the reflecting type the changes in temperature, causing expansion and contraction in the tube and mirror, alter the position of focus. Changes in the focus may be made by altering the position of the Cassegrain mirror and in the 74-inch telescope this is done by a small motor mounted in the tube which attaches the mirror in position. The observer has only to touch a button at the lower end of the tube and the focus is adjusted in or out as desired.

THE DECLINATION AXIS AND HOUSING

The declination axis, to which the tube of the telescope is attached, consists of a steel forging 13 feet long with a flange at the inner end 3½ feet in diameter. Its weight is about 3½ tons. A hole has been trepanned through its centre to permit the electric wiring to pass to the tube and to the outer end. It is supported at the inner end by a

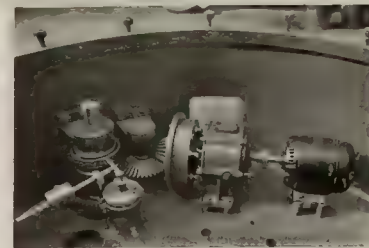


Fig. 8—Declination Quick Motion Gear.

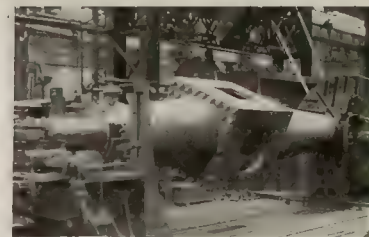


Fig. 9—Truing Polar Axis

radial ball-bearing 17 inches inside diameter, and at its outer end is carried by a radial bearing and double thrust bearing fitting into a tubular steel casting (H in Fig. 1) which is called the declination sleeve. All the bearings have been specially made with considerably less clearance than in commercial bearings. A spur gear (SG in Fig. 1) is keyed to the outer end of the declination axis which serves to turn the tube quickly in declination. The power for this motion is supplied by a motor inside the declination housing DII. A photograph of this motor and gearing is shown in Fig. 8. It can be clamped or unclamped by a magnetic clutch, the controls for which are mounted on the pier of the telescope. Bolted to the large spur gear is a graduated circle DC (Fig. 1) which by an index enables the observer to read the declination at which the tube is set. This declination circle is divided to single degrees, but in order to read the declination more accurately a fine spur gear FG bolted to the graduated circle drives two small drums DR, and these are graduated to a least reading of five minutes. Small electric lights illuminate the declination circle and drums so that the reading may be conveniently seen from the operator's position beside the pier. The declination housing is made intentionally heavy so that its weight counterbalances the weight of the tube. Additional weights are attached to the housing to adjust the balance. It can be seen from Fig. 1 that no extra weight has been attached to the declination axis proper, thus reducing flexure to a minimum and facilitating a free motion in declination. When the telescope has been brought to the correct declination setting, within a minute of arc, the quick motion gear is disengaged and the observer, on looking through the telescope, will see the star, but usually not in the centre of the field of view. It is necessary to have some fine adjustment whereby the tube may be moved by very small amounts. The mechanism which effects this is called the declination slow motion. It is shown in Fig. 1 and in Fig.

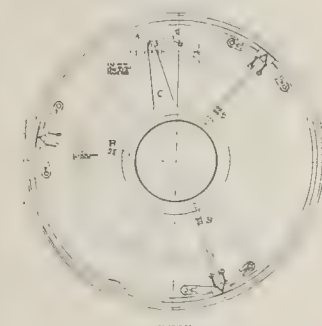


Fig. 10—Right Ascension Slow Motion Clamp

3 between the polar axis and the tube. A fabricated steel arm about seven feet long is mounted on a V-ring which can be clamped or unclamped to the polar axis. The outer end of the arm carries a nut mounted on a link motion which can be moved in either direction along a screw attached to a bracket fastened on the tube. A motor operates a two-speed gear to turn this screw for giving the slow motion in declination. The motor is operated by two separate switches. For the faster motion a dog-clutch, operated by a solenoid in parallel with the motor, connects the gear to the slow motion screw and moves the tube fifteen minutes of arc in one minute of time. For the slower motion an electro-magnet brings a differential gear into action and the tube moves thirty seconds of arc in one minute of time. The circuit which actuates the magnetic clutch in the declination housing for quick motion is interlocked with the clamping motor on the V-ring so that it is impossible to have the two clamps engaged at the same time.

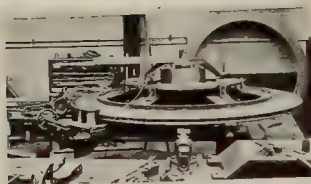


Fig. 11—Cutting Teeth in Driving Circle.
THE POLAR AXIS AND DRIVING GEAR

The polar axis, which carries the declination axis and housing, is best seen in Fig. 9 where it is shown on a lathe. It consists of three sections, a central hollow steel cubical casting to which is bolted two tapered hollow steel castings. Forged steel pivots are shrunk into the ends of the steel castings. Figure 9 shows the operation of turning the polar axis. It is 22 feet long and weighs 9½ tons. Having been trued up it was never taken apart but henceforth treated as a unit. It runs in self-aligning ball bearings with a ball thrust bearing at the lower end. The housing at the upper

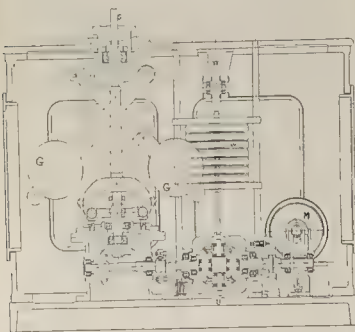


Fig. 12—Driving Clock.

end provides means of adjustment in elevation and azimuth to bring the axis *SS* in line with the celestial pole. Keyed to the lower end of the polar axis is a large spur gear *S* which engages with the quick motion motor *QM* in Fig. 1 for moving the telescope in right ascension. This motor is engaged with the gear by a magnetic clutch. When the clutch is engaged and the motor started, the telescope rotates in right ascension, and in order to bring the telescope quickly to rest when the motor is released, the clutch continues to engage for a few seconds and acts as a drag. This is a very convenient arrangement in preventing the overrunning of the telescope beyond the desired reading. The driving circle (*Dr. C* in Fig. 1) for moving the telescope, to follow the motion of the stars is mounted on the lower pivot immediately above the quick motion spur gear. It turns freely on ball-bearings but can be clamped to the axis through the spur gear. The clamping arrangement is shown in section at *RC* in Fig. 1 and a plan is shown in Fig. 10. A small motor (*A* in Fig. 10) rotates the arm *C* and the toggle arms *B*, which are mounted eccentrically, push out wedges into a V-groove on the inner edge of the driving circle. The driving circle itself is a steel casting, and a bronze rim, in which the teeth are cut, has been shrunk onto its outer edge. The pitch diameter is 8 feet cut into nine hundred and sixty teeth. The operation of cutting the teeth in this wheel is shown in Fig. 11, in which it will be observed that a finely divided silver circle is mounted directly over the worm wheel. The positions of the teeth are governed by the divisions on this circle so that the driving circle is a copy of this tested circle. This procedure eliminates the danger of periodic errors. The teeth were carefully hobbled and ground and the worm finally run in with rouge to ensure a perfectly smooth fit. Mounted above the driving circle at *FW* in Fig. 1 is a free wheel called the *sidereal circle*. It is mounted on the boss of the driving circle and can be easily rotated by any one of six hand-wheels attached to it and fitted with pinions gearing into a wheel attached to the driving circle. The metal rim of this wheel is 3 inches wide and graduated on both edges to one minute of time divisions. The lower set of divisions read against indexes fixed to the guard of the driving circle and therefore mark sidereal time, while the upper set of divisions are read by indexes which are attached to the polar axis and therefore read right ascension. On starting work at the beginning of the night the clock is started and the sidereal circle moved by means of the hand wheels till the lower graduations read the correct sidereal

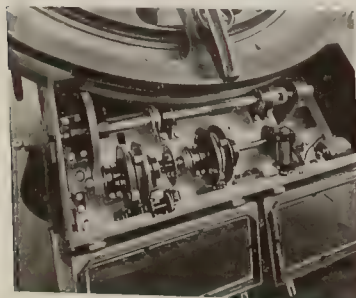


Fig. 13—Gear Plate and Driving Worm.

time. As the sidereal circle is attached to the driving circle the lower scale will continue to read the sidereal time. In order to set the telescope on any star it is necessary only to engage the quick motion and turn the telescope about to the polar axis until the upper scale reads the right ascension of the object. The driving circle is then clamped to the polar axis and the quick motion clamp freed. The quick motion clamp in right ascension and the clamp for the

driving circle are actuated by an interlocked circuit so that it is impossible to have both on at the same time.

THE DRIVING CLOCK

The position of the driving clock is shown at *I* in Fig. 1 and a cross-section through the vertical spindle is given in Fig. 12. The cross-arm governor *G* is driven by the weight *W*. The motor *M* turns at the right speed to keep the weight floating on its vertical guides. Ordinarily if the motor is turning at the correct rate the position of *W* will be stationary and the outer housing of the differential gear *U* will be at rest. But if the motor turns too fast or slow the weight rises or falls and the outer housing turns, without affecting, however, the torque applied to the governing spindle. In order to keep the weight *W* suspended and to govern the motor, the weight is attached to a lever arm, not shown in the drawing, which actuates a resistance in series with the motor slowing the latter down if the weight gets high and speeding it up if the weight falls low. As the governors speed up, the link motion spreads and brings into play two friction pins at *FF* and the speed is checked. The moment that these engage can be regulated by turning the heavy weights *GG* which are mounted eccentrically. A clock of this type runs very accurately but not with the precision necessary to keep the image of the star perfectly stationary. Therefore, instead of gearing the spindle *S* directly to the driving worm, it communicates to a gear plate shown in Fig. 12, and at *GP* in Fig. 1, the function of which is to correct for any small irregularities in the clock drive and at the same time provide means of slow motion of the telescope in right ascension if it is desired to shift the position of the star in the field of view. The spindle of the driving clock *S* connects with the worm and worm wheel of the gear plate shown at *D* in Fig. 13. If the clock is driving at exactly the right rate this shaft turns one revolution in twenty-four seconds. Mounted on this shaft at *CC* is a thin disc in the circumference of which are cut twenty-four V-shaped notches, and into these notches beats the armature of a solenoid which is connected to the seconds beat of a sidereal master clock. Provided the disc *CC* runs at the correct rate the armature accurately strikes the depressions but if the wheel runs too fast or slow the armature shifts the wheel slightly and it is connected with the differential *HM* to keep the worm drive in synchronism with the notched wheel *CC*. For a slow



Fig. 14—View of Site from Southeast showing Piers.

motion the differential *HM* may be rotated by the motor *M* and for a still finer motion the differential *GM* can be locked or released by a dog-clutch. The motion imparted by the motor *M* is never sufficient to completely overcome the normal rate of drive so that the driving screw is always engaged with the driving circle and hence all difficulties of backlash avoided.

The complete mounting as described above was received in Toronto on October 15th, 1933.

THE BUILDING AND DOME

The building to house the telescope, though out of the line of ordinary construction, did not require so long to complete. The contract for it was let in November 1931, and it was received on July 31st, 1933. In the meantime the piers to support the telescope and the foundations of the building had been built.



Fig. 15—Building with Dome Partially Erected.

Figure 14 shows the cement piers in place and the nature of the surrounding country. The site chosen is a hill about 800 feet above sea level and 100 feet above the surrounding country, 12 miles north of the city limits of Toronto. The piers consist of reinforced concrete and extend 25 feet underground to a foundation of hard clay which extends farther down for more than 150 feet. On the east side of the pier a pit 6 by 10 feet extends to the bottom of the foundations. An elevator is mounted in this pit for removing the mirror and mirror cell. The mirror can be readily taken off for the purpose of resilvering.

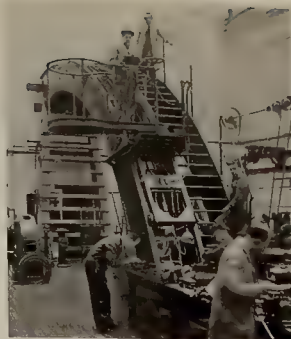


Fig. 16—Observing Platform.

Figure 15 shows the circular building and dome as it stood at the end of September. Since then the telescope has been put in place and the dome covered. As the construction of the building and dome presents a number of unusual features a brief description will be of interest. The entire building is of steel construction. The circular base and dome have double walls, and open louvers at the base of the building admit air which circulates to the top and exits through baffle plates at the top of the dome. The inside and outside of the dome is covered with "agasote," a kind of hard paper product, and the outside is further protected by a sheeting of copper. By this means the interior is kept cool during the day and in the evening when observations are started the whole building soon assumes the temperature of the outside air and the definition is not interfered with by heated air currents which would undoubtedly be the case if any quantity of heat were stored in the walls.

The dome, which weighs about eighty tons, rests on twenty-four rollers 27 inches in diameter mounted on self-aligning ball bearings and running on a circular track. Lateral rollers keep the dome in position as it is rotated. The rotation is effected by means of an endless steel cable passing around an annular channel ring fixed to the base of the dome and over two tangent pulleys to a winding gear. The dome is 61 feet in diameter and has an opening at one side, 15 feet wide, extending from the bottom to 7 feet beyond the zenith. Two parallel shutters run on rails at the top and bottom of the dome and are operated by wire ropes connected to a motor-driven gear which opens or closes the two shutters simultaneously. To protect the telescope from wind blowing in at the open shutter, two wind screens, motor operated and consisting of heavy sail cloth, can be used to cover the part of the opening not in use. One runs from the top of the opening downward and the other from the bottom upward and are guided in tracks fixed to the inside of the shutter opening. A novel feature of the dome, and one which should make it more convenient to operate than any previous models, is the method of access to the upper end of the tube. Two parallel segmental platforms are fixed to the inside of the dome, the lower one at the base of the opening and the upper one at a 16-foot higher level. Rails are mounted on these platforms and a circular air bridge connects the two platforms. The bridge is shown in Fig. 16. The entire bridge can be moved laterally on the rails of the platforms and a small observing carriage travels up the curved arch. The small observing carriage can be swung around on a pivot to enable the observer to place himself in a convenient position with respect to the eyepiece of the telescope.

No effort has been spared by the manufacturers to make the entire mechanism, mounting and dome in the best way possible. The workmanship is of the highest quality throughout and when finally completed the telescope will be the largest in the British Empire.

The magnitude of the telescope will make it a powerful instrument of research, and the major part of the programme of the observatory will be devoted to investigations into the motions and distribution of the stars. One evening per week however will be devoted to viewing the heavens when the public will be invited to visit the observatory and look at the planets, the moon and other objects in the sky.

NOTE: The illustrations, with the exception of Figs. 14 and 15, which are from photographs taken from the roof of the administration building of the observatory, are from photographs and drawings kindly supplied by the Sir Howard Grubb, Parsons and Company.

THE ENGINEERING JOURNAL

THE JOURNAL OF
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VOLUME XVII FEBRUARY 1934 No. 2

A New Canadian Observatory

Much of the recent progress in astronomical knowledge has related to the stars and nebulae, and is the result of patient observation with greatly improved equipment. This has furnished astronomers with a mass of information on such matters as the stars' temperatures, motions, distances and constitutions, all based on the study of the messages which the astronomer deciphers from the light sent to him from the skies. In fact, astronomy has become closely connected with atomic physics and the astronomer and the physicist are working hand in hand. For such investigations the astronomer needs instruments of a capacity and precision undreamt of in the days of Newton and Huygens. Of these, the most important is, of course, the telescope with its photographic and spectroscopic accessories.

The design and construction of a great telescope is now an undertaking which requires the highest skill of the mechanical and electrical engineer, and presents manufacturing and optical problems which very few engineering works are able to solve successfully. As Dr. Young points out in his paper, printed elsewhere in this issue of The Journal, it is necessary to design an instrument which can be moved slowly and with extreme accuracy, although weighing many tons, so as to follow exactly the motion of a celestial object. Further, if the image of the object is to have the necessary clearness of definition, the deformations of the telescope structure and the mirror it carries must be kept within very minute limits, whether such distortions are due to temperature changes or to the weight of the parts of the instrument itself.

The paper describes the way in which these objects have been attained in the construction of the new 74-inch telescope for the David Dunlap Observatory of the University of Toronto. The provision of this instrument and the

building which houses it was due to the efforts of Professor C. A. Chant, head of the Department of Astronomy of that University, who for the past twenty-five years has urged the establishment of a research observatory there. In numerous public lectures, in The Journal of the Royal Astronomical Society of Canada, and in the courses in astronomy within the university, the need of an observatory was constantly stressed.

During his life, the late David A. Dunlap took a keen interest in astronomy and in the meetings of the Astronomical Society. When, therefore, some time after his death, Dr. Chant approached Mrs. Dunlap with the proposal that she erect an observatory as a memorial to her husband, the project was very sympathetically received, and as a result plans were made for the observatory, the account of which we are now privileged to publish.

Some of the largest telescopes of the world are listed below:

1. The 100-inch reflector at the Mount Wilson Observatory, California.
2. The 72-inch reflector at the Dominion Astrophysical Observatory, Victoria, B.C.
3. The 69-inch reflector at the Perkins Observatory, Delaware, Ohio.
4. The 60-inch reflector at Harvard College Observatory, Boston.
5. The 60-inch reflector at the Mount Wilson Observatory, California.

The telescope for the new David Dunlap Observatory will have an aperture of 74 inches. It will therefore be the second largest in the world, and its construction gives Canada two reflecting telescopes of the first rank.

The new telescope will be devoted almost entirely to astronomical research. The advantage which it will possess over smaller instruments is due to its great light-gathering power. Taking the light-gathering power of the unaided eye as a unit, that of various sized telescopes may be expressed in terms of this unit. A 2-inch telescope would possess 64 times the light-gathering power of the eye, a 10-inch telescope 1,600 times and a 74-inch telescope 87,616 times. Possibly another and more striking way to illustrate the effectiveness of the large telescope is by a comparison of the number of stars which can be seen in telescopes of various sizes. There are about 4,000 stars visible to the naked eye in the whole celestial sphere on a clear moonless night. A 2-inch telescope will show about 300,000 and with the 74-inch telescope it will probably be possible to photograph over 500 million. Thus the ability of the large telescope to photograph faint stars enables it to reach far into the depths of space.

When a spectrograph is attached to the telescope the spectra of the stars can be photographed and it is possible to determine their constitution, temperature, distance and speed of motion in the line of sight. To photograph the spectrum of a star at the limit of naked eye visibility requires about two hours with a telescope of 15-inch aperture; with the 74-inch this exposure time can be reduced to ten or fifteen minutes. Thus a great deal more work can be done in a given time with the larger instrument. The actual difference in practice is greater than these numbers indicate, owing to the danger of the long exposures being ruined by clouds and other troubles.

But there is another advantage of the large telescope over the small, and that is the clarity with which it shows detail on any surface. The layman thinks of this as magnifying power; a better term is resolution. In order that detail may be seen on an object it is necessary that the images of two nearby points of the object be seen as two separate images in the telescope. The smaller the angular separation of two points which can be seen apart, the

greater the detail which can be made out. This ability to distinguish contiguous points is called resolving power and it varies directly with the aperture of the telescope.

In order that the full resolving power of a telescope be developed it is necessary that the beam of light from the source pass through a perfectly uniform medium and ordinarily even when the sky is clear the atmosphere of the earth is very far from this condition. In looking at the stars through a small telescope of 2-inch aperture the uniform beam of light needs to be two inches in diameter only, but in a 74-inch telescope there must be a uniform beam 74 inches in diameter. The former condition is met with on very many more nights per year than the latter. It therefore happens that the casual visitor to an observatory may be disappointed with the clarity with which he sees the moon or planet through the large telescope, often not better than through a much smaller instrument. The fault does not lie with the telescope but with the atmosphere. When the latter is steady, or, as an astronomer would say, when the "seeing" is good, the full aperture can be utilized to advantage and much more detail seen. The observer must seize these opportunities to do his best work.

The new telescope will be used on every available clear night from sunset to sunrise in photographing the stars or obtaining their spectra. An observatory however would not be fulfilling its complete purpose if used entirely for research. Part of the time must be devoted to those interested in seeing the wonders of the skies and it is planned to set aside one evening per week when the observatory will be open to the public and the telescope available to them. From such a visit some idea may be gained of the patience and skill which the observer needs in order to keep a star image on the slit of his spectroscope, or obtain a clear photographic record in spite of adverse atmospheric conditions and the necessity for prolonged exposure.

CANADIAN CHEMISTRY METALLURGY

March 1934

TORONTO CHEMICAL ASSOCIATION

Armour Plate Glass—Laminated Safety Glass—
Dunlap Telescope Lens
Large Lens for Dunlap Observatory

Prof. C. A. Chant gave a very interesting description of the difficulties encountered in securing the large 76-inch lens required for the new telescope at the Dunlap Observatory. The contract for the entire telescope was let to the firm of Sir Howard Parsons in 1930, and the first glass company approached was a subsidiary of that firm at Derby. They hesitated to tackle the job, and after much delay a commercial glass concern in England offered to produce the lens from crown glass in 15 months.

In the meantime Prof. Chant heard of the experimental work being done at the Corning Glass Works in the United States, in preparation for a 200-inch lens for California. A special grade of Pyrex glass was prepared for this purpose, with an exceptionally low coefficient of expansion. This huge lens was to be made with a honey-comb structure on the lower side, in order to cut down its weight and time of cooling.

About the time Prof. Chant communicated with the Corning Glass Works, they had produced an experimental 60-inch lens and had little hesitation in accepting the contract for the 76-inch one required here. The telescope having been built for the greater weight, the honey-comb principle was not employed, but the one side of the lens was cast concave to reduce the grinding.

Casting the Lens

In casting the lens, the glass from a great 75-ton furnace at a temperature of 1,500 °C. was transferred to the mould, which was kept at 800 °C. under a low pressure. After the pour, the temperature was raised to 1,250 °C. for 15 hours, to secure homogeneity. The lens was then transferred to the annealing kiln, soaked at 550 °C. for 9 days, allowed to cool 2.2 °C. per day, until a temperature of 380 °C. was reached, and then cooling then speeded up to 15 °C. per day. During the annealing, the lens was examined under the polar microscope and satisfactory results were obtained for grinding. The grinding process will require nearly a year, and the finished lens will arrive about the end of 1934.

The meeting was under the auspices of the Canadian Institute of Chemistry, with Mr. A. R. Bodham in the chair. The President of the Institute, L'Abbé Alexandre Vauchon, was present and brought greetings from Quebec to the Toronto group.

PLATE IV



FIG. 1. The great telescope in the observatory building.
 FIG. 2. The great telescope in the observatory building. When it is at the top of the dome, it is at the top of the observatory building. Photographed at Newcastle in December 1934.

THE JOURNAL

OF

THE ROYAL ASTRONOMICAL SOCIETY
OF CANADA

Vol. XXVIII, No. 3

MARCH, 1934

Whole No. 232

THE 74-INCH TELESCOPE OF THE DAVID DUNLAP
OBSERVATORY

By R. K. YOUNG

(President's Annual Address, January 16, 1934)
(With plate IV)

THE buildings and preliminary equipment for the David Dunlap Observatory are nearing completion. The Administration Building is finished. It will provide space for offices, for library, and for laboratories required in the measurement of the photographic plates and in subsidiary investigations. The copper-covered steel dome and circular building for the large telescope has been erected, although considerable work still remains to be done. The mounting for the telescope is in place. The large mirror is at present in England being ground (see Plate IV), and is expected to be ready within a year.

The two buildings stand out clearly on the sky line to the east as one travels north on Yonge St. between Thornhill and Richmond Hill. The Administration Building is built of stone along classical lines, and does great credit to the architects. The building which houses the great telescope does not lend itself to the same architectural possibilities, but nevertheless it harmonizes well with the Administration Building. The observatory stands on the highest ground in that vicinity, the ground floor of the great dome being about 800 feet above sea level. It is in the midst of an estate of 179 acres. This amount of land seems to many unnecessarily large, but there is a very important reason for having it. The large telescope will be used almost exclusively for photography and must not be interfered

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R. K. Young

with by any extraneous light. It is therefore essential to prevent too many buildings and houses being erected in close proximity to the observatory. The small village of Richmond Hill, about one mile to the north west, offers an objection to the site, but it is not serious because fully three-fourths of the work will be done with the telescope pointing south of the zenith and the lights from the village will not shine directly into the dome.

The land around the observatory is at present quite open, with a few trees and shrubs scattered here and there. It would be much better from an astronomical point of view if the land were more heavily wooded. It is intended to make the land into a park, to be known as the David Dunlap Park, and is very desirable to plant trees over the area as soon as possible. The reason for having the ground covered with a growth of trees is to shield it from the sun on the warm summer days. On an open plain the ground becomes quite hot during the day, and when, at sunset, this heat is given off again the warm air flowing upwards creates an unsteady atmosphere which interferes with the use of the telescope. The trees and vegetation absorb a great part of the sun's heat so that a steady state of the atmosphere is reached much more quickly. It is planned to have the site developed under the superintendence of the Department of Forestry of the University. Although the project of the observatory was started five years ago, this part of the scheme has been delayed by circumstances beyond our control. It should be carried out without delay.

Most of you however when you think of the observatory have in mind chiefly the large telescope which will be housed in the "Big Dome", and it is of it that I shall speak this evening. There will be smaller telescopes in the domes on the Administration Building capable of doing a great deal of good work. The small dome to the south will house the 19-inch reflector made some time ago in the University. The central dome will contain a 10-inch or a 12-inch visual refractor, and in the north dome will be mounted a photographic refractor. These subsidiary telescopes will be especially useful in an observatory so closely associated with instruction in the University.

The principle of the telescope was discovered about the year

The remainder of this address was the same as that at Montreal, for which see a previous page.

THE NEW YORK TIMES, SUNDAY, MARCH 4, 1934.

Casting a Giant Lens



A SPACE-PIERCING EYE.

Wonders Involved in Making the Mightiest Telescope Glass.

At Corning, N. Y., the mold is ready for the casting of a disk of glass 200 inches in diameter, which is to become the mirror or light-catcher of the greatest astronomical telescope ever constructed. Twice the diameter of the Carnegie Institution's famous 100-inch Mount Wilson reflector, this new instrument will be about ten times as effective. When it is done it will have entailed ten years of research and have cost \$5,000,000, advanced by the International Education Board.

It might be supposed that size of mirror is a matter merely of a mold big enough. Astronomers know better. It was difficult to find a glass-maker who would undertake the casting of the 100-inch Mount Wilson mirror. Even then the block obtained was not perfect. In the case of the 200-inch mirror there have been practice castings, one of them yielding a disk of glass 120 inches in diameter.

The method of casting having been thus established, the stage is set for an event which is to take place this month and which is bound to arrest the attention of astronomers all over the world. Glass will run out of a furnace, a dazzling white-hot river, into a truck. Its temperature will be 1,500 degrees C. By the time the truck has reached the mold the temperature will have dropped to 1,000 degrees. Twenty tons of liquid will flow like water into the mold.

Then for four months the block will be allowed to cool. "Annealing," the technicians call this slow reduction of temperature. Its object is to avoid the strains which would be set up if the outer parts cooled quickly while the interior was still hot and soft. The strains would crack the mirror, just as a lamp chimney sometimes cracks for no apparent reason.

Clearing With Electrons.

Once cooled, the glass must be ground and polished to give it a slightly concave surface. The light of a star falls on every part of the surface, and from the surface it must be reflected to a focus, which is a mathematical point. If the grinders and polishers are wrong by a few millionths of an inch the image at the focus will be blurred. As the polishers proceed they do less work every day. Gradually the glass is heated by friction. Heat causes expansion. To polish the surface in that condition means that there can be no true focusing after contraction. So work is stopped and the disk covered for the day with a foot-thick layer of felt. When the covering is removed, twenty-four hours later, optical tests are made to find out where grinding and polishing are needed in order to bring the final correct shape what the mathematicians call a paraboloid of revolution.

Five years of work already lie behind those who have carried the project so far. Most of the preliminary research has been conducted by the physicists and astronomers of Mount Wilson Observatory of the Carnegie Institution of Washington for the California Institute of Technology to which the final installation will belong. Two years must be spent in grinding, surfacing and polishing the mirror. Then comes the difficult task of installation. It may be 1935 before we shall hear of the first revelations made by the mightiest astronomical eye ever devised.



The Tentative Model for the 200-Inch Telescope. Its Huge Size Is Indicated by the Figure at the Base, Above-Casting an Astronomical Mirror.

Possible Revelations.

And what will these revelations be? Prediction is impossible. But if the triumphs that must be credited to the 100-inch mirror are any indication, we may expect a great visible addition to the 2,000,000 spiral nebulae, each a galaxy like our own; a piercing of space to greater depths and the discovery of systems undreamed of; a fine analysis of the spirals that whirl in such great agglomerations as the Great Nebula of Andromeda; a deeper understanding of the Milky Way and our own place in it; a verification or refutation of the theory that the universe is expanding; new light on Mars and its enigmatic "canals"; and a study of matter in the stars, at pressures of billions of tons to the square inch and at temperatures measured in unthinkable millions of degrees, which will help to clarify some of the mysteries of matter.

WORLD'S LARGEST TELESCOPE

CONSTRUCTION IN NEW YORK

FROM OUR OWN CORRESPONDENT
NEW YORK, March 26

Twenty tons of molten glass were poured yesterday into the mould of what will be by far the largest telescope mirror ever constructed. It was an anxious undertaking which took 10 hours to complete, but though there were some mishaps it is believed to have been successful. That, however, will not be known with certainty for another 10 months—so long will the mirror be allowed to cool in order to prevent cracking.

The mirror, which is to be 200 inches in diameter, is being made at the Corning Glass Works, Corning, New York, for the California Institute of Technology, and will be set up at the Mount Wilson Observatory, Pasadena.

When the pouring was half over yesterday some of the cores became detached and floated to the surface of the white-hot glass. The director of the work was confident, however, that these could be scooped out when pouring was finished without harming the mirror.

Among the 4,000 spectators of the work was a group of distinguished men of science, including Sir William Bragg, Dr. George Ellery Hale, Honorary Director of the Mount Wilson Observatory, Dr. W. S. Adams, Director of the Observatory, and Dr. C. A. Chant, of the Dunlap Observatory, Toronto.

The focal length of the mirror will be 55 ft. It will gather four times as much light as the largest telescope now existing, and will make possible the photography of objects four times farther out in space than those which can now be photographed. Photographs will be taken 10 times faster than they can be taken now. A hint of the possibilities opened up by the telescope is afforded by the fact that astronomers will probably be able to photograph with star clouds 12,000,000,000 light years distant. But a long time as human affairs go must pass before it will be set up and ready for use. Two years will be required after the mirror is taken from the mould and sent to California to grind it for the work it has to do.

The largest telescope in existence at present is the 100-inch reflector at the Mount Wilson Observatory. Funds for the new observatory are being provided by the International Education Board, which administers Rockefeller benefactions.

London Times, Tue. Mar. 27, 1934

MONDAY: The Daily Mail MARCH 26, 1934

MAKING THE WORLD'S BIGGEST MIRROR

FOR £1,200,000
TELESCOPE

MOLTEN GLASS OF
2,800 DEGREES

10 Months to Cool

From Our Own Correspondent
New York, Sunday.

ONE of the world's greatest experiments is in progress to-day at Corning, New York State, where 4,000 scientists from all parts of the world are gathered.

They are watching 20 tons of liquid glass being transformed into a mirror for the world's largest telescope.

The complete telescope is to cost £1,200,000, and the mirror will be 200 in. in diameter, compared with the Mount Wilson Observatory telescope of 100 in.

To-day's job is a major step in making the giant reflector. Glass is being poured from a furnace in which the temperature is 2,800 deg. Fahrenheit. The mould is already heated and can be cooled off only a few degrees a day, so that it will be ten months before the mirror is solid.

5 YEARS' PREPARATION

Dr. George McCauley, physicist in charge of the operation, is going without sleep for 48 hours until he is assured that the experiment is successful.

For five years astronomers and scientists in the glass industry have worked on the problem, yet many hearts are beating fast in the Corning glass works as the job proceeds.

The actual pouring occupies eight hours, the workers being provided with special clothing against the intense heat, but it has taken a month to heat the furnace to the required temperature.

The mirror being cast is not a solid cylinder of glass. Its back is honey-combed with holes arranged in geometrical patterns which make for lightness and a greater facility in use. It will have ten times the ability to gather light compared with its 100 in. brother.

Whether this new giant "eye" will solve the mystery of life on Mars is one of those fascinating questions which await an answer as the Corning glass men work on the lens.

TUESDAY: The Daily Mail MARCH 27, 1934.

200-INCH MIRROR'S
BLINDING FLASHES
THRILLING SPECTACLE OF
WHITE-HOT GLASS

CROWD SEIZE SOUVENIRS

MISHAP MAY MAR WORLD'S
GREATEST TELESCOPE

From Our Own Correspondent
New York, Monday.

THE world's greatest experiment for prying into the secrets of the universe has been completed.

The giant mirror for the £1,200,000 telescope, which will be the largest in the world, is now in its mould—liquid glass that will take ten months to cool.

While 6,000 people, including Sir William Bragg, the scientist, who is now lecturing at Cornell University, filed through the Corning (New York State) Glass Works, an official of the company explained that

in the opinion of physicists the whole casting operation is one vast experiment, and not until the great disc is removed from the mould and in one piece, without evidence of any strain, will the success of the casting be known.

Yesterday's operation was not entirely successful. The mould contained geometrically arranged tunnels or pylons across its floor, the object of which was to provide corresponding holes in the mirror to reduce its weight.

WHITE-HOT "SOUP"

Some of these tunnels, which were clamped by iron bolts which melted in the intense heat of the furnace, broke away and floated to the top of the liquid glass. They were fished out of the white-hot "soup" by workmen. When the mirror is ready for use

the places where the holes should have been will have to be drilled out with sand-blasting apparatus without cracking the mirror.

Visitors enjoyed the spectacle of the filling of the mould amid a terrific din and blinding flashes from the furnace fire.

Standing 50 ft. from the furnace, with its 2,800 degrees of heat, was the mould, a giant dome-shaped fire-brick structure heated by blasts of white-hot gases.

Grumpy, sweating workmen thrust iron ladles, each holding 400 lb. of glass into the blinding furnace and guided them to the great mould.

The front man carried a wooden shield round his neck, which he held by means of a multipiece in front of his face. This protected him from the dangerous ultra-violet rays in the furnaces.

Streams of glass which came from the ladle's "overflow" were seized as souvenirs by the spectators.

Scientists believe that such an experiment will never again be undertaken—first, because the giant telescope was conceived in the days when the United States was flush with money, and, secondly, because a 200 in. mirror is regarded as the limit of possibility in this field.

Dr. Walter Adams, director of Mount Wilson Observatory, who watched the casting of the mirror, said that it will enable astronomers to explore 27 times as much space round our solar system as is now possible.

THE TIMES
WEDNESDAY
MARCH 28 1934

THE LARGEST TELESCOPE

FROM OUR OWN CORRESPONDENT
NEW YORK, March 27

It is feared that the 200 in. telescope mirror, the largest in the world, now being made at Corning for the California Institute of Technology, will be found defective because of mishaps which occurred while the glass was being poured into the mould on Sunday. When some of the cores of the mould became detached, the iron bars which held them in place melted. It will be six weeks before the glass has cooled sufficiently to examine it for defects. Meanwhile another mould will be made in case it is necessary to do the work all over again.

SCIENTISTS MEET TO WATCH LENS MADE

CORNING, N. Y., March 26 (A. P.)—Distinguished scientists from all over the continent gathered today to see the creation of a cyclopaen lens, or the world's largest telescope.

Motion glass for a 200 in. lens will be poured tomorrow at the Corning Glass Works, an operation which has attracted the interest of the whole world of science.

Among those here for the event are Sir William Bragg, noted British scientist, and Dr. C. A. Chant, director of the Dunlap Observatory, Toronto.

\$6,000,000 Eye
Sees Across Back

Toronto Mirror Witness 20-Inch Mirror Cast

Corning, N. Y., March 26 A huge telescope eye, the latest I ever made, and expected to reveal the most mysterious details of the universe, is being slowly born in the molten glass of a newly created mold experiment. Twenty tons of molten glass, which were poured yesterday into the blue mold of the Corning Glass Works, will be fished out of the white-hot "soup" by workmen. When the mirror is ready for use

The Daily Mail

TUESDAY, MARCH 27, 1934.

The Giant Telescope

Scientists will regret the mishap yesterday in the casting of the reflector. For the gigantic new telescope which is under construction in the United States. It is certain to delay the completion of the instrument, and years may pass before it is at work.

When it is brought into action, with a power ten times greater than that of the 100 in. telescope at present in use at Mount Wilson, it should solve many problems. One of the most interesting of these is whether vegetation exists on the moon. Within the last few years some observers have thought they detected changes in the crater known as Linné. Most astronomers, however, still remain sceptical.

Another problem to be settled is the tremendous one of the existence of life on Mars. The new telescope should reveal whether the so-called canals really exist and whether they show signs, as the late Professor Lowell believed, of the desperate efforts of a dying race to cope with a desert world.

Montreal Daily Star
March 24, 1934



No 1038.

Edging the Main Mirror, General View. The disc is being made perfectly circular and 76 inches in diameter. May 1934



No 1039

Edging the Main Mirror-View showing the Grinding Tool. The disc continually rotates. May 1934

Reprinted from "Engineering" (England) March 9, 30, and April 30, 1934

74-IN. REFLECTING TELESCOPE

FOR THE

DAVID DUNLAP MEMORIAL OBSERVATORY

OF

TORONTO UNIVERSITY,
CANADA.

SIR HOWARD GRUBB, PARSONS AND COMPANY,

OPTICAL WORKS

NEWCASTLE-UPON-TYNE, ENGLAND.

*Proprietors, C. A. PARSONS AND COMPANY, LIMITED,
HEATON WORKS, NEWCASTLE-UPON-TYNE*

Reprinted from

"ENGINEERING,"

March 9, 30, and April 20, 1934

LONDON

OFFICES OF "ENGINEERING", 35 and 36, BEDFORD STREET, STRAND, W.C.2.

1934.

Publication No 10

74-in. REFLECTING TELESCOPE FOR THE UNIVERSITY OF TORONTO.

Constructed by
Messrs. SIR HOWARD GRUBB, PARSONS and COMPANY,
NEWCASTLE-UPON-TYNE.

It is pleasing to note that no mortals to prominent men who have passed away, instead of taking the form of some ornate and useless structure, are now often intended to benefit a local community, or even the whole human race, by helping to extend existing knowledge in some particular branch of science. That made by the widow and son of Mr. David A. Dunlap, late of Toronto, Ont.,

stalled, the site includes a block of administration buildings on which three domes for smaller telescopes are mounted. Few firms in the world are capable of constructing such an instrument as the 74 in. reflector, and it is particularly fortunate, therefore, that the order for it was secured by a British company, viz., Sir Howard Grubb, Parsons and Company, Optical Works, Walker Gate,



FIG. 1. OBSERVATORY BUILDING ERECTED BEFORE DISPATCH

of the latter type of instrument, a great advantage is provided for the use of the large telescope as the British Imperial Observatory at the said. This consisted of a reflector having a clear aperture of 74 in. and a focal length of 100 ft. and when completed will be presented to the University of Toronto by the Department of Astronomy of that University. It is located on a hill 177 ft. above sea level, and a few miles to the south of Toronto. In addition to the telescope, the building is also being in-

Newcastle upon Tyne, in face of both American and Continental competition. This firm, being associated with Messrs. C. A. Parsons and Company, Limited, is in a particularly fortunate position for the construction of large telescopes, since it has the use of Messrs. Parsons turbine shops for the accurate machining of heavy castings and other parts with which an optical works could not economically deal. The Optical Works are, however, fully equipped and staffed for the production of large mirrors and other optical parts, as well as for the light and heavy instrument work, of which so much is involved in the construction of a telescope of such large aperture. Work was commenced

on the Toronto telescope in November, 1930, and it was dispatched from the makers' works in September, 1933.

The instrument is housed in a cylindrical building, 61 ft. in diameter, sheathed inside and outside with steel sheeting and illustrated in Fig. 1 on page 1. The entrance is in the south side, through a steel porch with two pairs of doors, one pair of which can be seen in Fig. 1. The observing floor is 13 ft. above the ground level and a doorway leads from this floor to the top of the porch. From this stair-

constructed both the framework of the dome and the building to the designs of Messrs. Sir Howard Grubb, Parsons and Company. The dome is 61 ft. in outside diameter, and has a parallel opening 15 ft. wide, extending from the horizontal to a point 7 ft. beyond the zenith. Two parallel moving shutters running on rails at the top and bottom of the dome, as will be clear from Fig. 14, close the opening; these shutters being operated simultaneously by means of wire ropes connected to a motor-operated gear; emergency hand gear is also provided. Two motor-

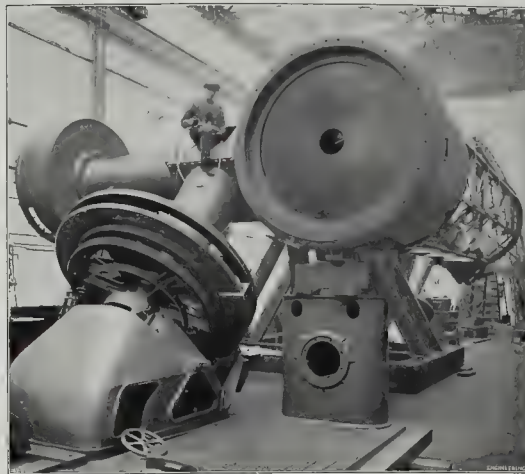


FIG. 2. TELESCOPE LOOKING FROM REAR OF CENTRE PIECE

ways give access to a gallery running right round the building at a height of 23 ft. above the ground level. A similar gallery inside the building gives access to a segmental platform in the revolving dome which covers the building. From the segmental platform the observer can reach an observing platform which is used when the telescope is employed as a Newtonian. A lift is provided inside the building for removing the cell and main mirror when the latter requires to be re-silvered.

Fig. 14, on page 8, shows the dome erected complete, except for the roofing material, in the yard of Messrs. The Cleveland Bridge and Engineering Company, Limited, Darlington, who

operated wind screens, of sail cloth, are mounted in the opening, one ranging from the bottom and the other descending from the top, so as to close the opening except for that part through which the telescope is actually pointed.

The dome is carried on 24 casted rollers, 27 in. in diameter, mounted in self-aligning ball bearings and running on a flat-bottomed rail. This rail is mounted on a strong annular girder which is supported on 24 steel columns, each 21 ft. high. Lateral rollers, of which 16 pairs are fitted, run on the inner and outer edge of the rail to keep the dome in position. Two segmental platforms, one at the base of the opening and the other at the back,

which, by means of the former, is automatically raised or lowered in the vertical direction, is supported on a platform on which rails a large in the form of a sun gear. This gear, which is 1 ft. 6 in. in diameter, is 5 ft. 6 in. in diameter, and is divided into 24 teeth, each 1 ft. 6 in. in diameter, and is the first of a series of gears which drive the telescope carriage on the telescope carriage. The instrument is arranged as a Newtonian. If used as a Newtonian, it is a 74 in. telescope, and if used as a Newtonian, it is a 74 in. telescope, and if used as a Newtonian, it is a 74 in. telescope.

its diaphragm provided to enable the aperture to be adjusted to suit observing conditions. We will describe this component in detail later, but may here mention that it is thought to be the largest diaphragm of the kind so far made, having a range of aperture from 74 in. to 12 in.

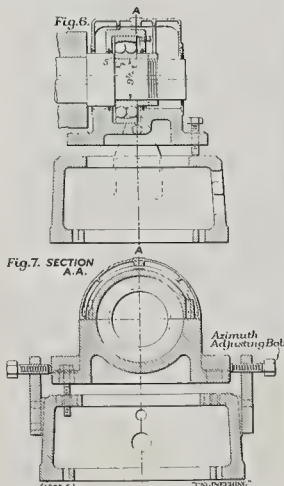
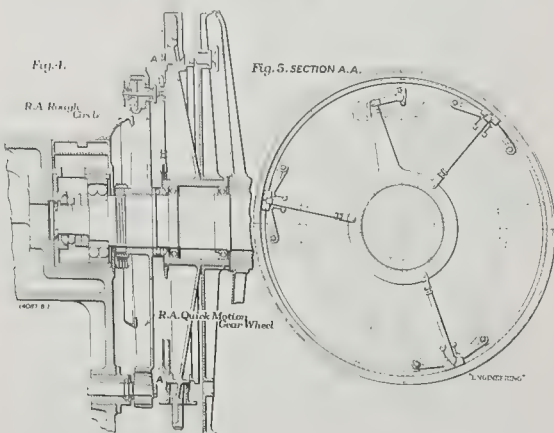
The telescope can be used either as a Cassegrainian or Newtonian. With the former, a convex mirror is used at the upper end of the tube, and this reflects the light back to a focus below the main mirror, which has a hole in the centre for the light to pass through. When used as a Newtonian, a

length and weighs 9 tons, is, however, best shown in the general arrangement drawing, Fig. 3, on Plate, from which it will be seen that it is built up of two tapered tubular steel castings bolted to a central cubical steel box and having forged steel pivots shrunk into the ends. It turns in self-aligning radial ball bearings at each end, a ball thrust bearing being fitted at the lower end, which is housed in a cast-iron base casting, as clearly shown in Fig. 3. The polar axis is, of course, set parallel with that of the earth, and the bearing at the

best shown in Fig. 5. It will be seen from this that when it is required to clamp the circle to the quick-motion gear wheel, and thus to the axis, a small motor mounted in the quick-motion gear wheel is set in motion, and this drives a small crank mechanism through worm reduction gearing. A connecting rod couples the crank to the radial arm extending from a floating central ring, to which are pivoted three rods, 120 deg. apart. The outer ends of these rods each bear on a pair of hinged shoes, which fit into a groove formed on a ring attached to the circle, as clearly shown in Fig. 4. The three rods, it may be pointed out, are not quite radial when the circle is unclamped, but the operation of the clamping motor, as explained above, turns them into the radial position and thus forces the shoes into the groove. When so clamped the axis can be driven by the clock, as will be explained later, but, when unclamped, the telescope can be turned about the polar axis at a comparatively high speed by means of a motor, and can also be adjusted by hand by the mechanism indicated by dotted lines in Fig. 3. We shall refer later to the quick motion in right ascension when dealing with that in declination.

The circle is a steel casting fitted with a bronze ring in which the teeth are cut, the ring being slightly shrunk on and fixed by a number of Delta metal screws. The pitch diameter is 8 ft., and 960 teeth of 8-mm. pitch were cut with the aid of a 4-ft. diameter master circle graduated on silver. On the extreme right of Fig. 4 is shown the submeral circle, which is 8 ft. in diameter, and is mounted on the loss of the driving circle. It can easily be rotated by one of six handwheels provided, five of which are distinguishable in Fig. 15, page 9, while one is shown in section in Fig. 4, page 3. These handwheels are fitted with a pinion which gear with a toothed ring on the driving circle, as will be clear on reference to Fig. 4. The submeral circle is of welded steel construction and is fitted with a Delta metal ring 3 in. wide, graduated on both edges to 1 minute of time. The lower set of divisions read against indices fixed to the guard of the driving circle and mark sidereal time, while the upper set of divisions read against indices attached to the polar axis and mark right ascension. An hour-angle circle, clearly shown in Figs. 2 and 11 on page 2 and page 6, respectively, is fixed to the quick-motion gear wheel.

Coming now to the declination axis we may first refer to Fig. 8, on page 5, which shows a longitudinal section through the declination axis, with part of the polar axis on the left. The declination axis itself passes at right angles through the cubical axis stiff passes at right angles through the cubical axis and is supported at its outer end in a tubular portion of the polar axis, bolted on to the cubical portion of the polar axis, weighing 3½ tons, and formed, as shown in Fig. 8, with a flange at the inner end, 3 ft. 5 in. in diameter, to which the telescope tube is attached. The tube turns in ball radial bearings with a double-thrust bearing at the outer end. The quick-motion gear wheel is keyed on to the outer end of the axis



by means of an endless steel rope passing almost round an annular channel curved brackets fixed to the base of the dome, over two tangent pulleys, and down to the turning gear, a tension dome and building were dispatched in June, 1933.

The general arrangement of the telescope can be followed by inspection of Fig. 2, page 2 of these. Fig. 15 is a view of the instrument in the makers' shops before dispatch, with the tube pointing south. Fig. 16 is a view from the south-east with the tube vertical, and with the mirror box moved, the tube lying horizontally. This view clearly shows the large

flat, set at an angle of 45 deg. with the axis, is fixed near the upper end of the tube, and this brings the light to a focus at a point outside the tube, where an eyepiece, or photographic camera, is mounted when required. A stellar spectrograph is being constructed for use with the telescope arranged as a Cassegrainian. It is of the single-prism type with a 21 in. collimator and two cameras of 12½ in. and 25 in. focus, respectively.

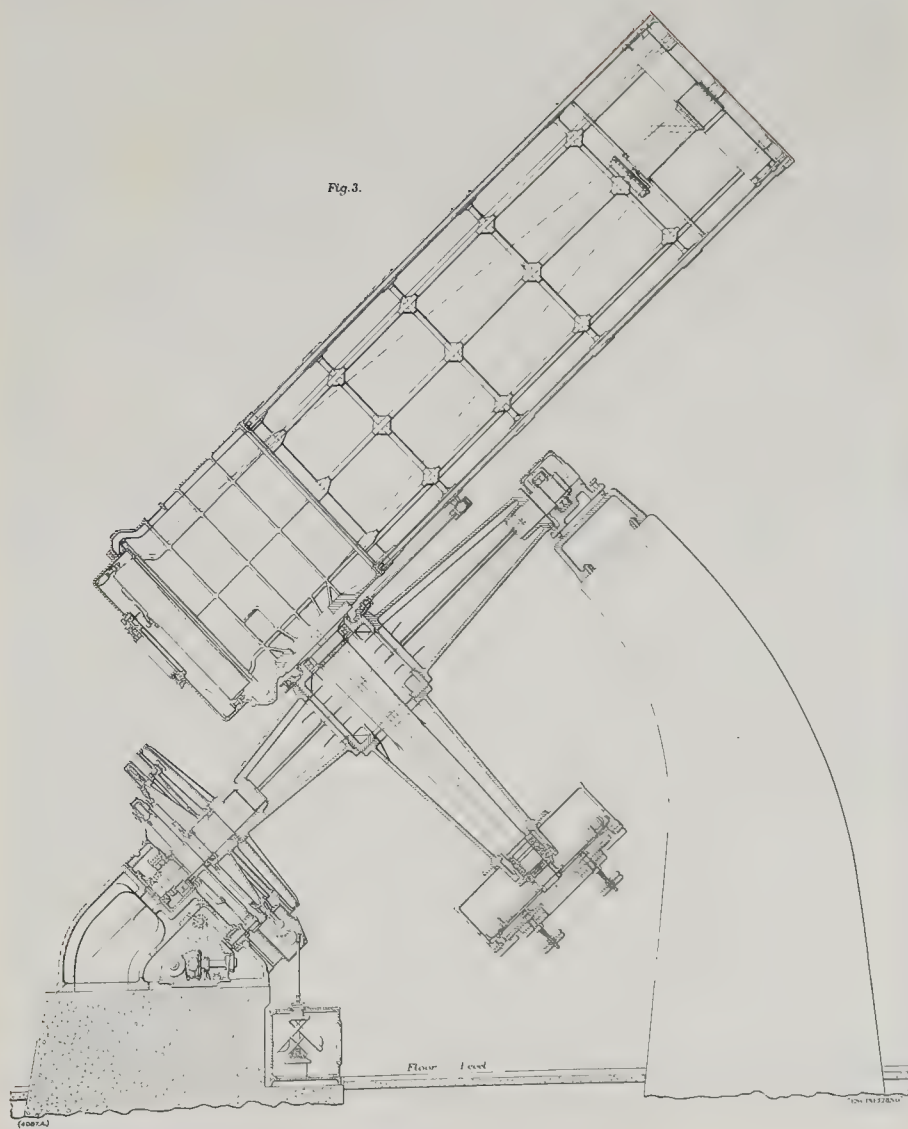
The polar axis complete with the necessary lifting into position in the erecting shop, is shown in Fig. 11. It will be recognised from the illustrations above that the mounting is of the English or composite type, in which the tube is placed on one side of the polar axis and the counterpoise on the other. The design of the axis, which is 22 ft. in

upper, or north, end is carried on a concrete pedestal, as shown in Fig. 3. Details of the design of this bearing are given in Figs. 6 and 7, annexed, which show the arrangements provided for adjusting the position of the axis in the vertical and horizontal planes to secure accurate alignment.

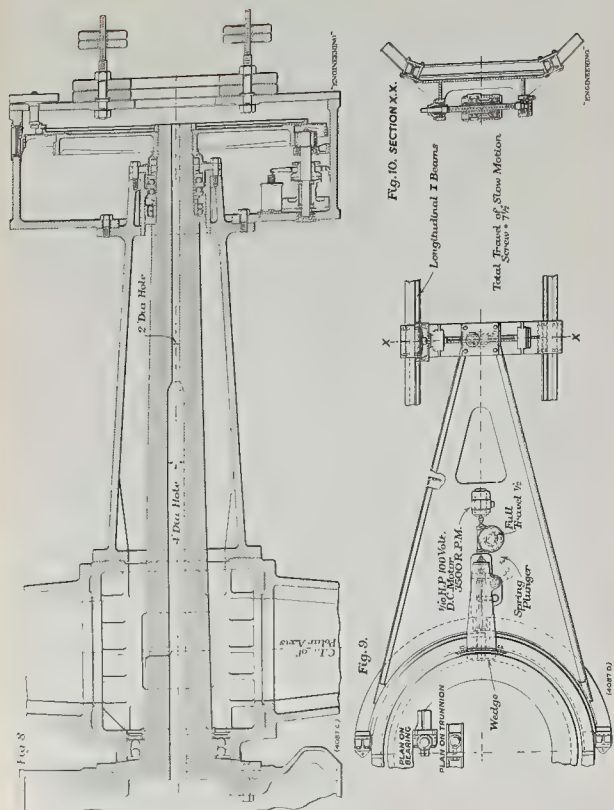
Longitudinal and transverse sections of the lower end of the polar axis, showing the right ascension clamping gear, are given in Figs. 4 and 5, respectively, on page 3. The actual driving circle, which is in the form of a worm wheel, is mounted free on the lower pivot of the axis on ball bearings, but it can be clamped rigidly to the quick-motion gear wheel, which is keyed to the axis, by means of a motor-operated clamping gear

PLATE

Fig. 3.



1-1000000



as also the declination circle, which is 6 ft. 3 in. in diameter at 1 in. graduated to 1 deg. of arc. The gear wheel and circle will be seen from Fig. 8, as shown in a drum-shaped casting fixed to the end of the tubular support carrying the axis, thus forming the counterpoise for the telescope tube. An interesting feature in connection with the declination circle is that it has been arranged to drive two drums, each 12 in. in diameter, the gear ratio being a step-up of 1:22. There is no backlash in the drive and the drums are graduated with 60 day lines, each of which represents 5 min. of arc. One of the drums can be seen in the top right-hand corner of Fig. 8, and a window, through

polar axis and the telescope tube, and is indicated in Fig. 3, although not shown in Fig. 8. Both the clamp and the slow-motion gear in declination are, however, clearly shown in Figs. 9 and 10, on page 5, and by the photograph reproduced in Fig. 12, on page 6. The clamp consists of a welded steel arm, about 7 ft. in length, mounted on a ring 4 ft. in diameter, attached to the side of the polar axis and having a Vee-groove in it. The arm can be readily clamped to the grooved ring by toggle gear operated by a small motor, as clearly shown in Figs. 9 and 12. The outer end of the arm is fitted with a nut mounted on a link motion, and engaging with a screw mounted in bearings attached to a

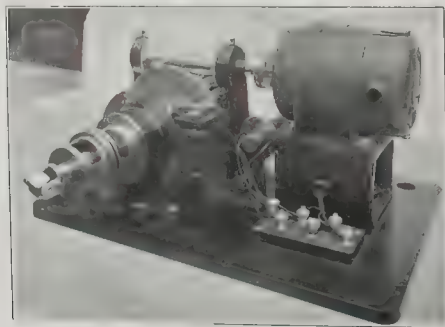


FIG. 13. DECLINATION TWO-SPEED SLOW MOTION GEAR.

which the declination circle can be read directly, will also be noticed.

Both the right ascension and the declination quick motions are operated by reversible motors, that for the former being of 11 h.p. and that for the latter of 7 h.p. These are both arranged to rotate the axes at the rate of one revolution in 8 min. The drives are taken through dog clutches operated by solenoids in parallel with the motor circuits, so that the motors and reduction gears are automatically disconnected from the telescope when not in use. A friction clutch is also incorporated, and the switchgear controlling the motor is arranged so that when the motors are switched off the dog clutches remain engaged for a few seconds, allowing the friction clutches to slip and thus bring the telescope gently to rest. The solenoid-operated dog clutch for the declination quick motion can be seen in Fig. 8, although the motor and friction clutch, which are, of course, housed in the drum-shaped casting, are not shown.

The declination clamp is located between the

bracket, which is firmly clamped to the side of the telescope tube. A motor, with a reduction gear giving two speeds, is connected to the screw by a universally jointed shaft, as shown in Fig. 12, and by this means the setting and guiding motion in declination are obtained. The motor and two-speed gear are shown separately in Fig. 13, on this page. The motor is controlled by two separate reversing switches. For the setting motion, a dog clutch, operated by a solenoid in parallel with the motor, connects the gear to the slow-motion screw and moves the telescope tube in declination at the rate of 15 min. of arc in one minute of time. For the guiding motion an electromagnetic in parallel with the motor brings a differential gear into action, giving a rate of motion of 30 seconds of arc in one minute of time. Similar rates of motion are provided for setting and guiding in right ascension.

Before describing the optical system of the instrument, we may mention that the tube, best shown in Fig. 3, on Plate, is made in three

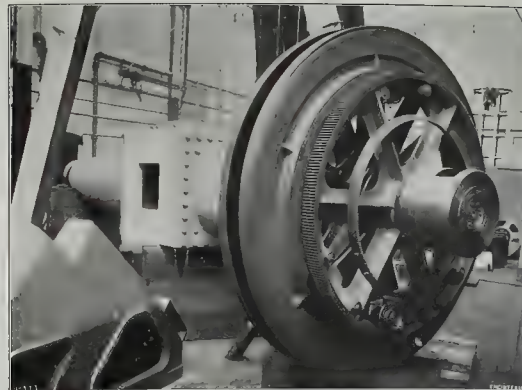


FIG. 11. POLAR AXIS READY FOR LIFTING.

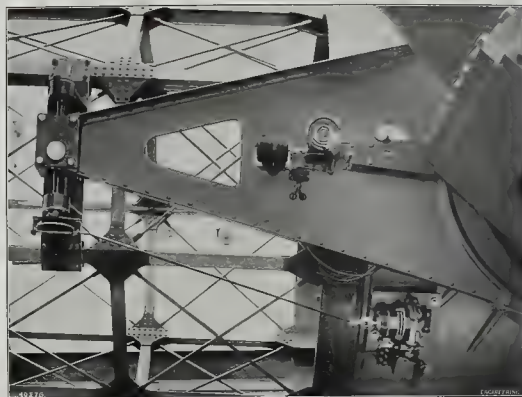


FIG. 12. DECLINATION CLAMP AND TWO-SPEED GEAR.

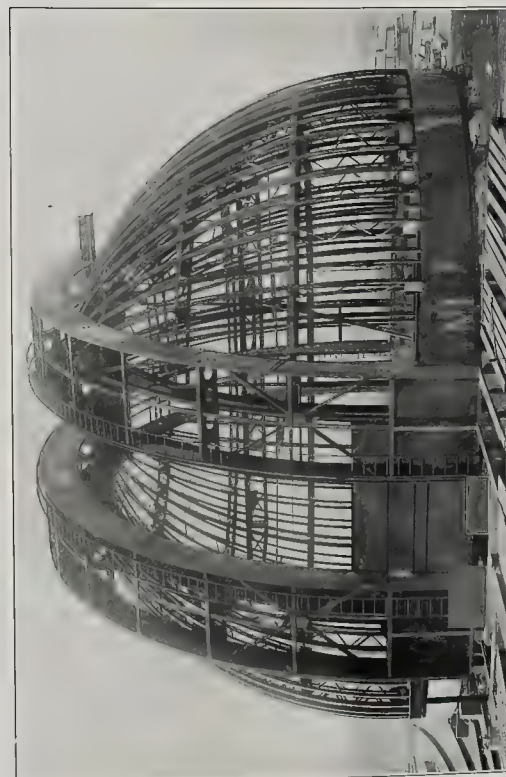


FIG. 14. DOME ERECTED IN MAKERS' YARD.

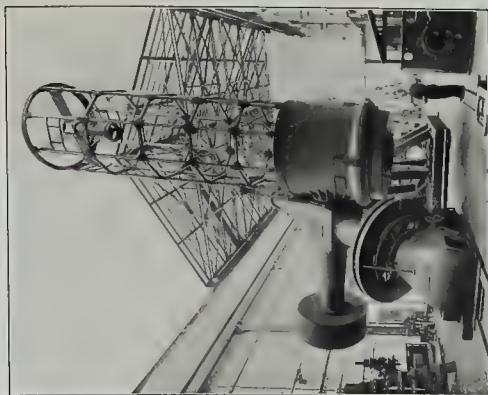


FIG. 14. TELESCOPE FROM SOUTH-EAST POSITION.



FIG. 15. TELESCOPE FROM NORTH-WEST POSITION.

ball bearings with spherical seatings. The latter are fitted on to screws, by means of which the mirror can be squared with the tube. For the lateral support of the mirror, when the tube is moved out of the vertical, 18 weighted levers, one of which is shown in Fig. 17, are provided. They are disposed round the inside of the cell and are mounted on universal joints. The short ends of the levers, as shown in Fig. 17, fit into holes in brackets riveted on to a flexible band, which is

but in order to prevent sagging of the leaves, owing to their considerable weight, and to keep the aperture central when the tube is near the horizontal position, it was found necessary to guide the moving ends of the leaves between bars having a ridges equal to the length of the leaves. Only four of the 20 leaves are thus guided, the other leaves being connected to the guided leaves by links and levers visible in Figs. 25 and 26. Of these illustrations, the former shows the diaphragm at its maximum

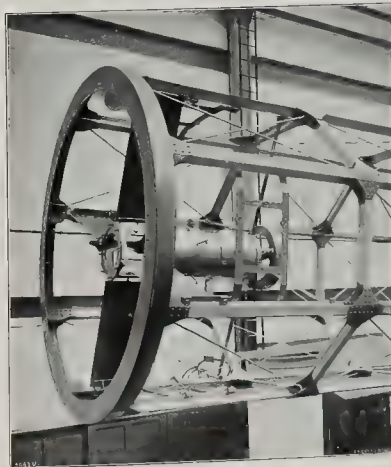


FIG. 22. CASSEGRAINIAN MIRROR MOUNTING.

loosely clamped round the mirror. This band, it may be noted, has blocks fixed at intervals on its inner face, and these blocks fit loosely into a groove formed in the edge of the mirror, thus serving to keep the band central. The object of the large worm wheel fitted to the back of the cell, as shown in Fig. 17, is to entry the spectrograph, which is partly indicated by chain-dotted lines.

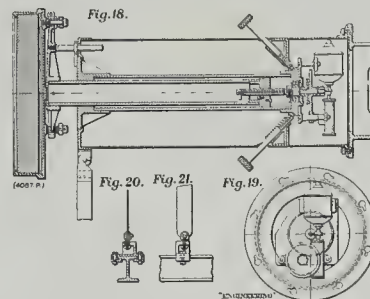
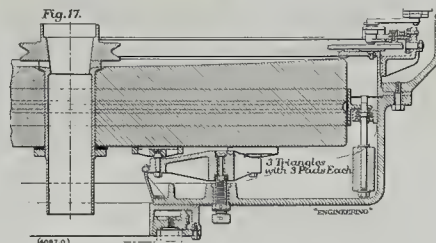
The iris diaphragm can be seen in section in the top right-hand corner of Fig. 17, and photographs of it before being fitted into the cell are reproduced in Figs. 23 and 24 on pages 13 and 14. The construction of the diaphragm is generally similar to that of the small diaphragm used in photographic cameras,

aperture of 74 m. and the latter at the minimum aperture of 12 m. The diaphragm is operated by a pinion and circular rack, the former being fitted with a handwheel which the man is holding in Fig. 25.

Coming now to the optical system, we may mention that the main parabola mirror is of a special Pyrex glass, the design for which was made in the United States. It is now being worked and figured by Messrs. Sir Howard Gribb, Parsons & Company; the focal length is 30 ft. The Cassegrainian and Newtonian mirrors are made of hard crown glass, and are 19 in. and 20 in. in diameter, respectively. The Cassegrainian mirror, which is convex and of

sections, of which the lower section is the cell in which the main mirror is mounted. The central section is a steel casting, 7 ft. in diameter, weighing 6 tons, and is formed with a large boss on one side to which the flange of the declination

diagonal tension rods of Duralumin. The rods are screwed with right-hand and left-hand threads, and are tightened up so that they are always in tension in any position of the tube. The construction of the tube is shown in Figs. 15 and 16 on page 9.



axis is bolted, as is most clearly shown in Fig. 8, on page 5. It will be seen, from Fig. 3, that just above the lower flange, to which the mirror cell is bolted, the casting is swelled out to 8 ft. 6 in. in diameter in order to accommodate the iris diaphragm; we shall describe the construction of this diaphragm later. The upper part of the tube is of skeleton construction and is octagonal in section. It is built up of I-sections of Duralumin connected by steel gusset plates, and braced with

the latter showing the Cassegrainian mirror in position, but the details can be more readily followed from Figs. 22 and 23, on pages 11 and 12.

The main mirror cell, a half section of which is shown in Fig. 17, above, is a ribbed steel casting fitted with four circular pads which support the back of the mirror, each carrying its proper portion of the load. The pads are mounted in groups of three, on spherical seatings carried by three triangles, which are themselves mounted on

hyperbolic form, is designed to give, with the main mirror, a focal length of 111 ft., so that the aperture is F.18. For mounting both these mirrors, a welded-steel box, of square section, with circular flanges, is suspended in the centre of the upper end of the tube on four strips of spring steel placed edgewise, so as to obstruct as little as possible of the light. These strips are clearly shown in Fig. 22, page 11, and Fig. 23, below. In this box, which is shown complete in Fig. 24, the

mirror mounting is illustrated in Fig. 23, below, and Fig. 24, opposite. It should be mentioned, in connection with this, that the reflected beam can be directed to any one of four points round the sides of the tube, where frames are fixed to which the photograph has been taken can be attached. In Fig. 24 the breech piece can be seen in position at the top of the illustration, and one of the three other frames to which it can also be attached is clearly shown in the centre of the illustration.

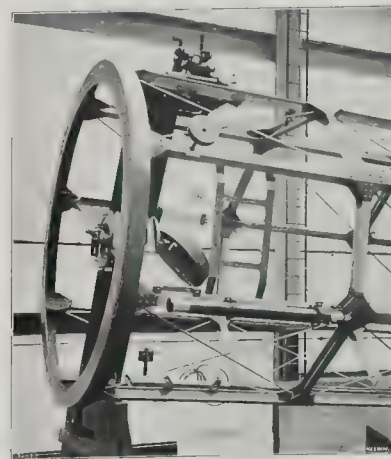


FIG. 23. NEWTONIAN GEAR IN POSITION.

mountings for either the Cassegrainian or the Newtonian mirror can be attached. Speed gear has been supplied for handling these mountings and interchanging them conveniently and safely. The Cassegrainian mirror mounting is shown in detail in Figs. 18 to 21, on page 10, and is illustrated in Fig. 22, page 11. So far as using gear, operated by an electric motor, is provided, as shown in Figs. 18 and 19, and the former of these shows the set screws for adjusting the Cassegrainian mirror into exact parallelism with the main mirror. Figs. 20 and 21 illustrate the straining device for the spring steel strips which support the lower end of the mounting. The Newtonian

The breech piece itself is shown separately in Fig. 27, page 15. It comprises focusing gear and a plate holder with two guiding microscope mounted on rods and operated by micrometer screws. Rotary motion is also fitted to correct for rotation of the field which sometimes occurs at altitudes, probably due to refraction. The plate holders take plates of quarter-plate size, and are interchangeable with a knife-edge focusing plate and with adapters for oculars.

Apart from the optical system, it will, we believe, be generally agreed that the most important part of a telescope is the driving mechanism. Its function is, of course, to turn the moving parts of

the instrument, in this case weighing about 35 tons, about the polar axis, to compensate for the rotation of the earth. When the telescope is being used for long exposure photography, extreme accuracy of drive is required, otherwise the star images would be drawn out into lines on the photographic plate.

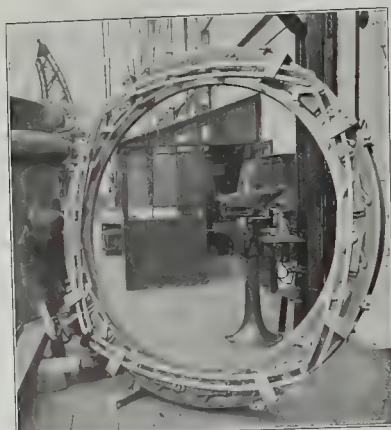
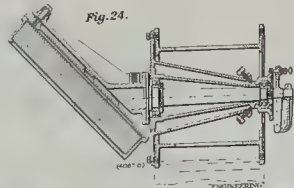


FIG. 25. IRIS DIAPHRAGM AT MAXIMUM APERTURE.

Dr. R. K. Young, President of the Royal Astronomical Society of Canada, has aptly illustrated this condition as follows: "A drift of the image of a star of 1/100 inch during the course of an exposure of 10 minutes would be sufficient to render the photograph useless for photographic purposes." The engineering difficulties confronting

the telescope maker in thus keeping the tube so accurately pointed are about equivalent to keeping a gun pointed on a target an inch in diameter at a range of twenty miles while the target is moving at the rate of 5 ft. per second.* The action, moreover, must be continuous and without any trace

of jerkiness. These arduous requirements are met by means of an electrically-driven clock and supplementary mechanism, which is corrected by a

* *The Refracting Telescope for the David Dunlap Observatory*, 1914, Paper before the Engineering Institute of Canada, Montreal, February 9, 1914, by R. K. Young, P. E.

entry of the armature into the notches has no effect, but if it should be fast or slow the disc is turned on way or the other relatively to the jack shaft. This relative motion operates a trigger connected to a switch which passes a lightning through the shaft and tips a two-way mercury switch at the right hand end of the jack shaft; the contacts of this switch are visible in Fig. 30. The effect of tilting this switch is to engage one of two electromagnets, the armatures of which

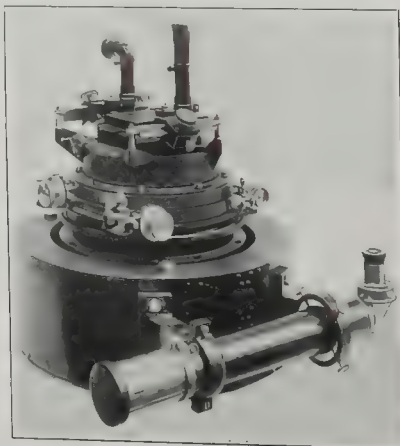


FIG. 27. NEWTONIAN BRIDGE PLATE.

are arranged to hold one of the discs carrying the pinion wheels of the epicyclic differential gear, and in this way the jack shaft is slowed down or speeded up, as required. The second epicyclic differential gear, viz., that on the left in Fig. 30, is controlled by hand and is used for correction errors due to refraction. It may be of interest to compare this gear plate with that for the reflector for Edinburgh, which was described on page 332 of our 128th volume (1929). If this comparison is made it will be seen that the two-way mercury switch takes the place of a considerable number of sliding and other contacts. The new gear, it may be mentioned, has only been fitted to the Toronto instrument and to a 36 in. reflector recently erected at Greenwich.

As already explained, the Toronto telescope can

be used both as a Cassegrainian and Newtonian, and since with the latter optical system the observer must be seated near the upper end of the tube, a specially-designed bridge, carrying an adjustable observing platform, has been provided for this purpose. The bridge, which is in the form of a semi-circle, runs on rails mounted on the edges of two segmental platforms fixed inside the dome, as mentioned earlier. A photograph of the lower portion of the bridge, after erection in the makers'

shops, is reproduced in Fig. 34, on page 19, and details of its construction are illustrated in Figs 31 to 33, on page 18. The horizontal distance between the running rails is 30 ft., and the difference of level between them is 16 ft. As already mentioned, the width of the bridge is 5 ft. 6 in., and this is about equally divided into a stairway and a track on which the observing platform runs. In the plan, Fig. 31, the method of driving the travelling wheels by a reversible motor through reduction gearing is shown dotted, and it will be noticed that power is transmitted to the lower travelling wheels through a long chain running over a pulley pulley.

The observing platform is mounted on a four-wheel trolley, the wheels of which run on the upper

very accurate seconds pendulum forming part of the equipment of all astronomical observatories.

A general view of the driving clock of the Toronto telescope is given in Fig. 28, on page 16, and a drawing of the mechanism is reproduced in Fig. 29. It comprises a heavy crossed-arm nearly isochronous governor driven indirectly by a 1-h.p. electric motor through a differential gear-box, shown in section in Fig. 29. A weight suspended from a chain passing over a pulley and attached to the outside of the

by a raising seen in the figure. The lower shaft, on which the correcting devices are fitted, as will be explained later, drives through spur gearing a second shaft, including a section on which an accurately-cut worm is formed, this worm engaging with an equally accurate worm wheel, called the driving circle, and mounted on the polar axis. The two shafts and the plate on which they are mounted are known as the gear plate.

Coming now to the correcting mechanism, we

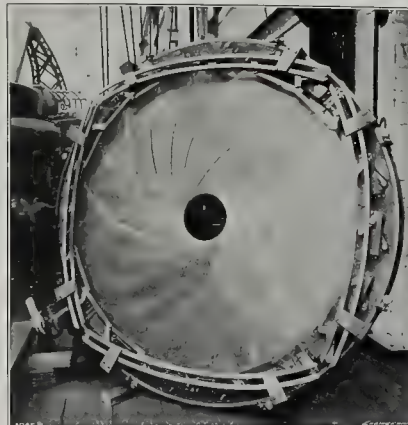


FIG. 26. IRIS DIAPHRAGM AT MINIMUM APERTURE.

differential gear-box, in the form of a Huyghen's loop, gives a constant driving force to the governor, and the gear-box carries an arm with a contact at its outer end, which passes over a number of studs arranged concentrically, and connected to resistances in the field circuit of the driving motor. When the motor is running at the correct speed, the weight is kept floating, but if the speed alters, the gear-box rotates, together with the contact arm, which then corrects the speed of the motor by altering the field resistance. The governor runs at 90 r.p.m., and is directly coupled by a vertical shaft, shown near the bottom of Fig. 3, on Plate, to a worm gear with a worm wheel mounted on the lower of the two shafts shown in Fig. 30, on page 17. The worm and worm wheel are shown at the right hand end of this shaft, the worm being protected

must first explain that the lower shaft, or jack shaft, is made in five parts, of which the first part reading from left to right in Fig. 30, carries the pinion driving the upper shaft, and this and the next two sections are connected together through epicyclic-differential gearing. The third, fourth and fifth sections are also connected through epicyclic-differential gearing, both sets of gearing being clearly shown in Fig. 30. On the centre section of the shaft is mounted, friction tight, a disc having 24 notches on the periphery, and opposite this disc is an electromagnet connected to the observatory seconds pendulum. This magnet, which is thus energised once per second, is provided with an armature of special shape, and this enters each of the notches in the disc, which is intended to make one revolution in 24 seconds. When the speed is correct, the

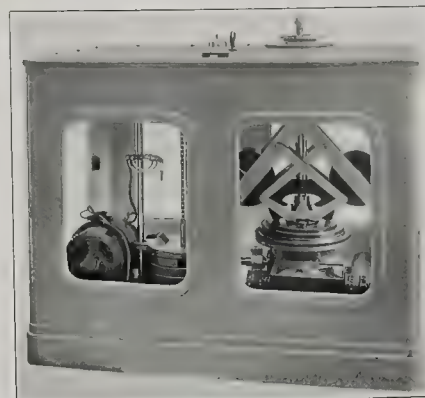
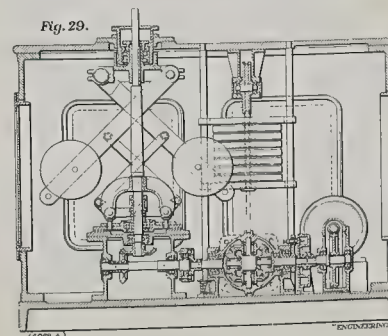


FIG. 28. DRIVING CLOCK.



17

flange of a pair of channels riveted to the side plates of the bridge. The trolley is held down by four wheels which run on the under surface of the upper flange, as is clearly shown in Fig. 33. The trolley is pulled up and down the bridge by wire ropes passing round an electrically-operated winch located on the upper end, as is best shown in the plan, Fig. 31. A smaller drum on this winch is used for a balance weight, which is shown dotted in its lowest position, and also half-way down, in the elevation, Fig. 32. The trolley on which the balance weight is mounted is shown in the plan, Fig. 31, and also in the transverse section, Fig. 33. It will be obvious that it is necessary to take steps to keep the observing platform level irrespective of

down the bridge, and the latter can be moved laterally in each direction. With these three movements the observer can obtain a convenient position near the Newtonian gear in practically any position of the telescope. There are two mirrors at the upper end of the telescope, one of 2½ in. and the other of 2 in. aperture. A third finder, of 4½ in. aperture, and with the eyepiece mounted on cross slides, is provided at the lower end.

To the brief reference to the mirror mentioned earlier, we may now add that the mirror disc, which has been supplied by the Corning Glass Works, New York, is made of a special Pyrex glass, having a coefficient of expansion on the Centigrade scale of about 2×10^{-6} , the corresponding figure for

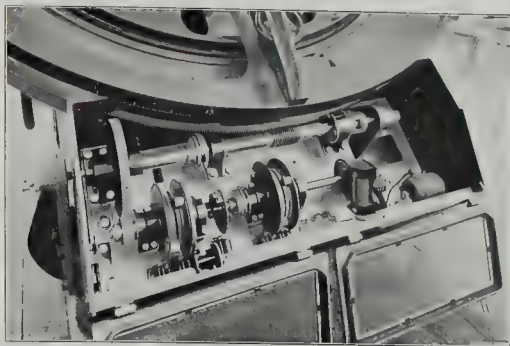


FIG. 30. GEAR PLATE WITH COVER REMOVED

its position on the bridge, and this is effected by mounting the platform on a short pedestal hinged to the upper part of the trolley, as is perhaps most clearly shown in Fig. 32. From the hinged pedestal extends an arm carrying at its outer end a small three-wheeled trolley which runs on a curved T-section attached to one of the web plates of the bridge. The curve of the T-section is such that the observing platform is turned about a horizontal axis as it moves up and down the bridge, just sufficiently to keep it horizontal. The platform can also be turned about a vertical axis, a ball race being provided for this purpose in the upper part of the pedestal, as shown in Fig. 33. To move the platform round the observer turns the handwheel shown in Fig. 33, and a pinion on the lower end of the shaft on which the handwheel is mounted engages with a fixed gear cut on the periphery of the ball race. The platform can thus turn about a vertical axis, and move up and

ordinary Pyrex glass being 3.2×10^{-6} . It was cast face downwards with a concave face of nearly the required curve, and also with the central hole mounded in. The disc reached the Walker Gate Works in November last, its weight being 2 tons 6 cwt. The disc, as received, was a full 14 in. in thickness, and as the mirror cell would only accommodate a thickness of 12 in., 2 in. had to be ground off the back. This has now been done and the back polished. At present the edge is being ground true, and the next operation will be to grind the central hole true. The concave surface has already been rough ground to approximately the right curve, but the final polishing and figuring will occupy some months. It is expected that the mirror will be completed by the autumn.

For grinding and polishing this mirror the machine illustrated in Fig. 35, page 20, was constructed by Messrs. Sir Howard Grubb, Parsons and Com-

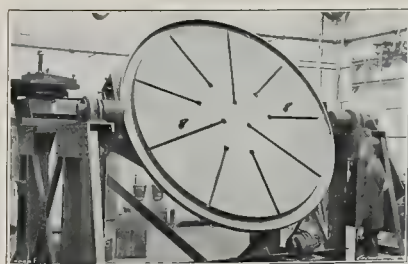


FIG. 35. GRINDING MACHINE WITH TABLE TILTED

pany. When clear of the ground a second crane lifted the northern end of the axis until the latter was tilted to about 44 deg., which is the latitude of the site. The southern end was lashed to the hook of the main crane, so that the inclination of the axis

could be maintained without the assistance of the second crane. The main crane then lifted the axis at the correct inclination, and passed it into the observatory building through the opening in the bearings. The whole of this operation, we understand, occupied less than an hour. The declination sleeve, weighing 3 tons, was raised into position and bolted on to the cubical portion of the polar axis. After which the declination axis itself was lifted and put into position. The erection of the mirror cell, tube and other details presented no particular difficulty, and the instrument now awaits the arrival of the spectrograph and the mirror, of which the latter, as already stated, will be completed towards the end of this year. As already mentioned, the observatory, including the 74-in. telescope, will be administered by the Department of Astronomy of the University of Toronto, and doubtless good use will be made of this magnificent instrument.

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pany. The table can take a disc up to 90 in. in diameter, but apart from its size its special feature is that the table can be tilted to the vertical position; it is shown tilted to about 15 deg. in Fig. 35. The machine is erected in a room provided with automatic temperature control, and located at one end of a long testing gallery. It is thus possible, by

interest. The circular building and dome, weighing nearly 200 tons, reached the site at the end of July, 1933, having been shipped directly to Toronto Harbour, and brought thence by road. The telescope, of which the total weight is about 50 tons, was shipped to Montreal, and was brought to Toronto by rail. From Toronto to the site the instrument

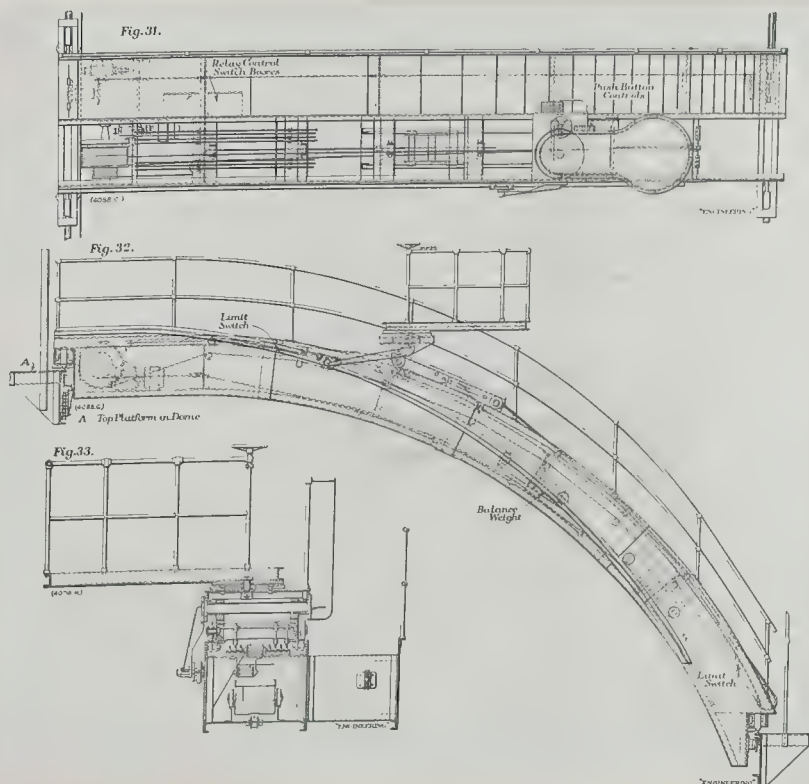


FIG. 34. LOWER PART OF BRIDGE FOR OBSERVING PLATFORM

tilting the table into the vertical position, to test the mirror without removing it from the machine. Many tests have to be made during the grinding and polishing operations, particularly when the final parabolic form is being produced, and the fact that these tests can be carried out without removing the mirror from the machine saves much time and trouble, and also avoids the risks attendant on handling such a heavy but fragile object.

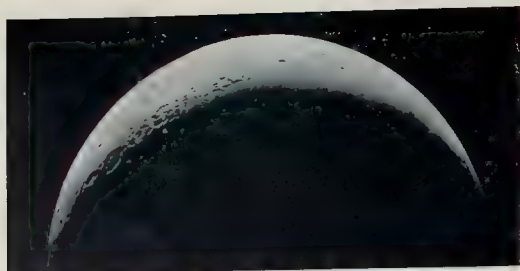
In conclusion, a few brief notes on the erection of the main parts of this telescope in Canada may be of

was transported by road. Both the building and the telescope were erected by the Dominion Bridge Company, of Toronto, the work being supervised by the foreman of the makers' erecting shop. The polar axis assembly weighed 16½ tons, and a 20-ton crane with a jib 92 ft. in length was erected to handle it. First the hoistings of the bearings were put in place on the north and south piers, shown in Fig. 3, on Plate 1. The driving and other circles were put in place on the south end of the axis, and the axis was then lifted by the 20-ton



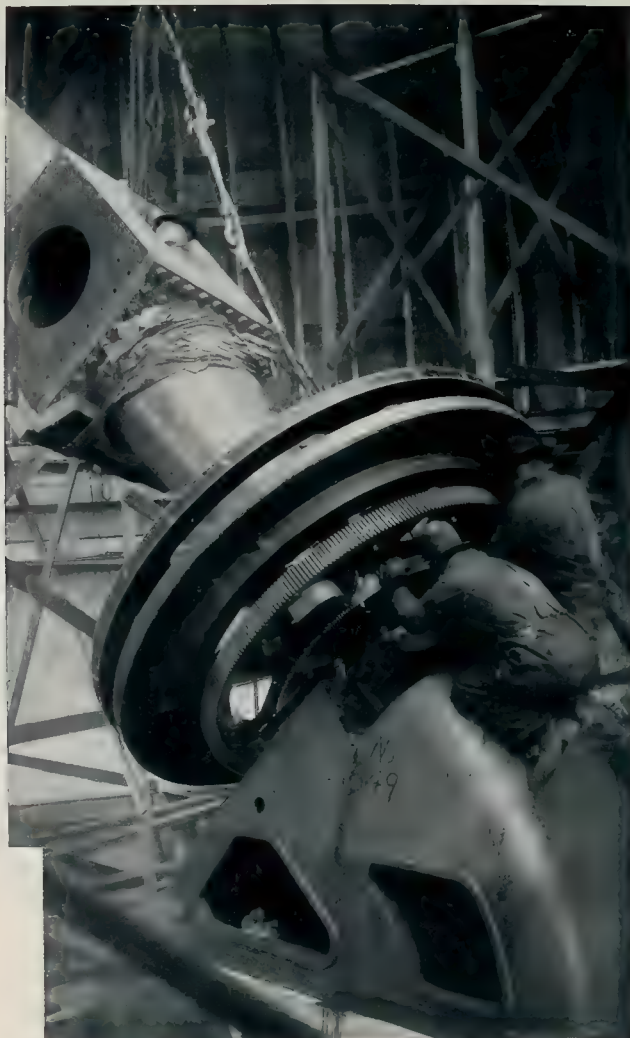
THE SPHERE

TWO GIANT TELESCOPES

April 21st. 1934

THE NEW 6-IN. REFLECTING TELESCOPE AT THE ROYAL OBSERVATORY GREENWICH which, together with the sunspot telescope in which it is housed has been presented by Mr. William Johnston to the Observatory in recognition of the work of Sir Frank Dyson as Astronomer Royal. In this picture the observer is seen using the guiding telescope with which he is able to correct small errors in the larger, clockwork actuated instrument.

These pictures of the new Greenwich telescope have been specially taken for THE SPHERE by William Davis and A. Connell, and are reproduced here by courtesy of Dr. H. Spencer Jones, the Astronomer Royal.



THE SECOND LARGEST IN THE WORLD The new David Dunlap Memorial telescope now being erected at Richmond Hill, near Toronto, Canada. (A full description of this instrument appears below)

It comes as a surprise to discover in Canada this new country, in an isolated spot atop a hill top in the midst of a pasture field, a magnificent telescope, the finest in the British Empire—finer far than anything in England, exceeded, indeed, by only one other in the world that on Mount Wilson, California. Yet this is what one finds today on Richmond Hill, fifteen miles north of Toronto, as he motors north on Yonge Street, that great artery of traffic which John Graves Simcoe, Canada's first Governor, cut through the bush, and which leads direct from the Queen City to that modern Aladdin's Cave, the rich mining region of Northern Ontario. The new David Dunlap Memorial telescope, which has been under construction for the last three years, is nearing completion. The administration building, of cut stone, is practically finished. The steel dome, 61 ft. in diameter, complete to the last word in equipment, which houses the telescope, is finished. The tube itself is installed. And the grinding of the great mirror—a long job—it must be correct to

one-millionth of an inch—which has been in progress for many months, is well advanced. The opening event is scheduled for this autumn.

An interesting but odd-looking affair, this telescope. With its mighty muzzle pointing to heaven—the tube is 30 ft. in length, nearly 7 ft. in diameter, and wide open at the top—at first sight it suggests a huge trench mortar, ready with a bomb in its throat to belch forth death and destruction on the beautiful city lying below. But this iron tube harbours no menace, far otherwise. In the very bottom of it lies the great mirror, face up, reflecting a circle of sky, jewel-like in its almost impossible beauty. It is about this costly treasure of a mirror that all the delicate machinery and expensive plant of this, the world's newest observatory, centres.

The mirror of the new telescope is 78 in. in width and nearly a foot in thickness. It is of Pyrex glass, a type of glass with a high silica content. Observations with a telescope must always be made through

(Continued on p. 124)

TWO GIANT TELESCOPES

(Continued from p. 105)

the widely opened shutter of the dome. Canada, for part of the year at least, is a cold country; and while temperatures in the dome of an observatory are kept as nearly as possible on a par with those of the outer air, there is usually a difference. As the temperature to respond. The small expansion or contraction of this type of glass means that a mirror made of it undergoes less change of figure and therefore gives more sharply defined images. It is a vitally important matter in all classes of astronomic work, but more especially in the study of the extremely faint stars in the spiral nebulae. The mirror is ground concave to a paraboloidal curve—a curve which differs somewhat from that of a sphere in that it focuses all the light rays at one point.

The polar axis of the David Dunlap Memorial telescope is 22 ft. in length and weighs 9 tons. It is supported at each end on masonry piers. Except that it revolves on its bearings like a great spindle, in slow steady motion, it does not vary its position. The declination axis, on the other hand, moves freely. It is the work-a-day axis of the telescope, on which the heavy tube turns when it is being depressed or elevated into position. This axis, of forged steel, is 13 ft. in length. It pierces the polar axis through its centre section cross-wise, and end projecting at right angles on each side. To one of these ends the tube of the telescope is securely bolted, and on the other end is loaded the counterpoise weight, a heavy, drum-shaped casting. Where the telescope is thus mounted—on one side of the polar axis, with a counterpoise on the other—the mounting is spoken of as the modified "English" or "Composite" type of mounting.

The tube of the telescope is not altogether unlike a cylinder—something resembling a great pipe on end or a section of smoke-stack, only the lower portion of it is enclosed. In the very bottom is fitted a ribbed steel casing which forms the mirror cell. Above this, in a swelled-out section, is lodged the iris diaphragm, a large camera shutter, with a spread of from 12 to 74 in. The middle section is a steel casting, very massive, supplied on one side with a large boss for attachment to a flange on the end of the declination axis. All the upper part of the tube is of skeleton construction. Within this, at the top, is the mounting for the supplementary mirrors. Underneath all at the bottom, is the fitting for the stellar spectroscope (supplied by Messrs Adam Hilger, Ltd.).

The seeing part of the human eye is very small. "How minute are our instruments, in comparison with the Celestial Universe!" once exclaimed El Karakat, an Arabian astronomer, who built a great observatory in Cairo in the twelfth century. He could see about 3,000 stars. Newton's little telescope increased

this a hundredfold, making visible about 500,000 stars (taking the sky as a whole). Canada's new telescope will have about 100,000 eye power, and thousands of those island universes which are so occupying the attention of astronomers to-day. The moon, when it comes up over distant only about fifty miles—if the man on the moon lives in a house the same size as ours and we can hit on just the exact spot we shall be able to see it. Mars, the nearest of the planets, is now, about 35,000,000 miles distant. Last spring it was little more than half that distance away. Five years from now it will be much less than that, only about 20,000,000 miles distant, which, astronomically speaking, is very close indeed.

By that time the students of the skies will probably have come to some pretty definite conclusions about this interesting neighbour—as to whether it is inhabited by men, about those great canals (?), the snow caps, the arid plains, the fertile sea bottoms, and whether life there, if there is any, corresponds to that on our earth. We'll be mighty interested in an object on Mars, to be seen here, would have to be pretty large, of course—at least thirty miles across. Regarding stars, the nearest of these, aside from our sun, is so very far away—a million times as far away as Mars—that we cannot hope for much in the way of intimate detail beyond what we already have, which is a fair idea of its size, weight, and composition. But nobody knows what it is made of, except that it will be a marvel. It will penetrate 140,000,000 light-years into space. How far is that? Light travels 186,000 miles per second. Nobody can comprehend it.

There is, of course, a story behind the building of this great telescope, as there is behind nearly every one of the great telescopes of America. The David

Dunlap Memorial telescope is the fruit of a suggestion by Professor C. A. Chant, M. A., Ph. D., chief of the Department of Astronomy of Toronto University to Jesse D. Dunlap and David Moffat Dunlap, the widow and the son, respectively, of the late David Dunlap, a prominent mining lawyer of Toronto, who during his lifetime was an enthusiastic student of the stars. This munificent gift places Canada in the position of having two telescopes of the largest class.

ONWARD APRIL 21, 1934

Canada's Place in the Stars

By G. H. MOSHER

THE University of Toronto has a new observatory, located on a height of land near Richmond Hill, a few miles north of the city. It is an impressive structure, with its grey stone walls and its four domes, and its setting of shrubbery and low trees, when you glimpse it from the nearby motor road of North Yonge Street. It makes you sit up and say sharply, "What place is that?"

It is the finest astronomical observatory in Canada, and, in Toronto U's opinion, long overdue. For years, Toronto's telescopic equipment, consisting of one twelve-inch instrument and several smaller, portable ones, has been obsolete. The great observatory

The David Dunlap Memorial observatory positively bristles with telescopes. Upon the main, or Administrative building there are three domes, a large, central one flanked by two smaller ones. Mounted in the central dome is a battery of three astronomical cameras, used for photographing heavenly happenings in the heavens, such as eclipses and showers of meteors. Incidentally, speaking of eclipses, by the time the next total eclipse of the sun occurs in the Toronto vicinity the Dunlap Observatory will be over two hundred years old, an ancient landmark, and you and I will be ancestors.

The smaller domes contain, respectively, a ten-inch "reflector" telescope for the use of the general public, and a nineteen-inch "reflector" telescope, made by Professor R. K. Young, assistant to Professor Chant. Although the observatory is at least fifteen miles from the Toronto waterfront if one of these telescopes were turned in that direction it would be possible to see a man painting the funnel of a boat at dock, and you could even see what color paint he was using. These two telescopes are, however, only "small potatoes" when compared to the show-piece of the observatory contained in the second building.

The second building is a steel dome, built alone rugged in its form, it must support the giant telescope, sixty tons in weight, almost as heavy as a railway locomotive. Two curved steel doors open and shut upon the aperture through which the telescope peers. They operate upon tracks built around the base of the dome. The telescope poking its snout out into the night through this aperture, looks not unlike a long-range gun.

Fifteen electric motors and innumerable

pieces of clockwork mechanism combine in the effort to draw the monster spy glass toward the spot in the heavens the astronomers desire to investigate. The telescope is set in a base of mercury to insure greater ease in swivelling.

The mirror lens of this telescope is seventy-four inches across, just two inches larger than the one at Victoria. It was ground to shape a concave hollow, by the firm of Sir Howard Grubb, Parsons & Company, engineers and giant telescope specialists of Newcastle-on-Tyne, England. The grinding job took three years. You have to place your order well in advance when you want one of these things. Telescope and dome cost \$500,000 and they are considered to be worth every cent of it because of the expert and highly skilled labour involved and the fine quality of material used.

Although the science of astronomy began in Europe with the construction of the first telescope by the Italian, Galileo, in 1609, and Europe subsequently saw the passage of many great astronomers, America now possesses the three largest telescopes in the world at Victoria, B.C., at Toronto, and at Mount Wilson, California. There the greatest telescope known to man peeps nightly through its dome on the steep, wind-swept side of one of the High Sierras. Its lens is eight feet four inches in size. Occasional newspaper reports advise one that Dr. Lesort is the man who constructed the present Mount Wilson telescope, and at work upon another instrument which will be exactly twice its size. When he finishes it, as the saying goes, "We shall see what we shall see."



Workmen Adjusting the Giant Telescope

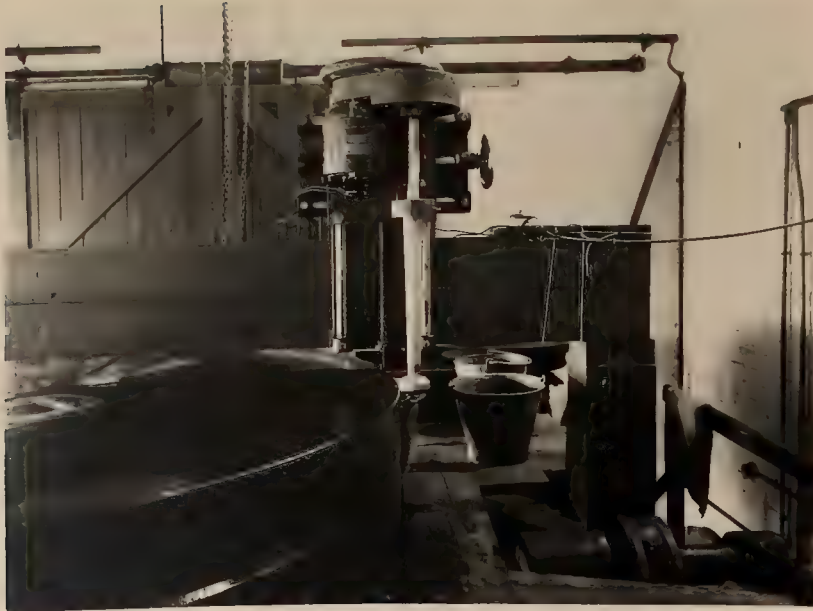
at Victoria, B.C., with its monster seventy-two-inch super-telescope, had caused Toronto University's astronomy students to turn jealous eyes westward.

Mrs. David A. Dunlap, widow of a wealthy man whose hobby had been astronomy, donated most of what it took to construct the new "Star Battery" at Aurora, and it took plenty of money!

Professor C. A. Chant, distinguished member of the Toronto University faculty and dean of Canadian astronomers, had been patiently putting up with inadequate equipment for a long time. Now he is happy because he has moved into his new home. You have to call it that for men like him spend more time in their "office" than they do in their own homes.



The David A. Dunlap Observatory, Near Richmond Hill, Ontario



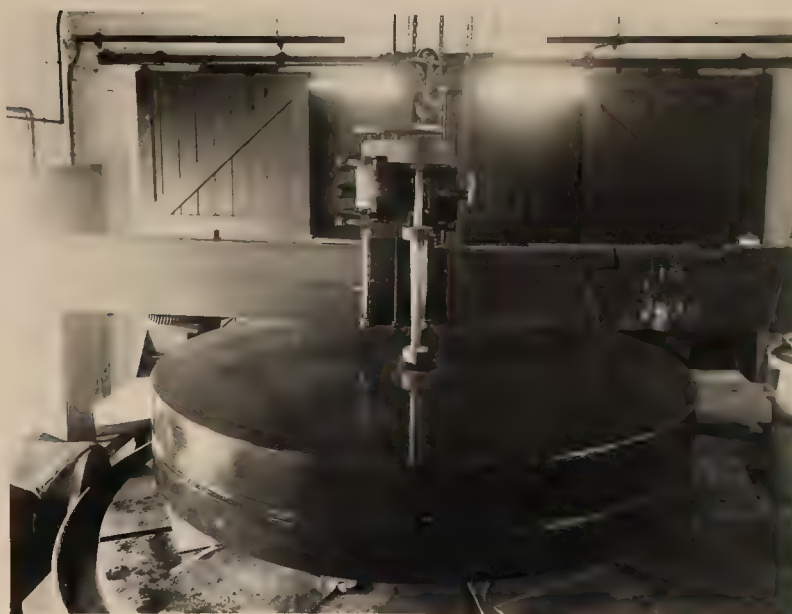
No 1043.

Revealing the edge of the Disc 76 inches in Diameter.

No 1044.

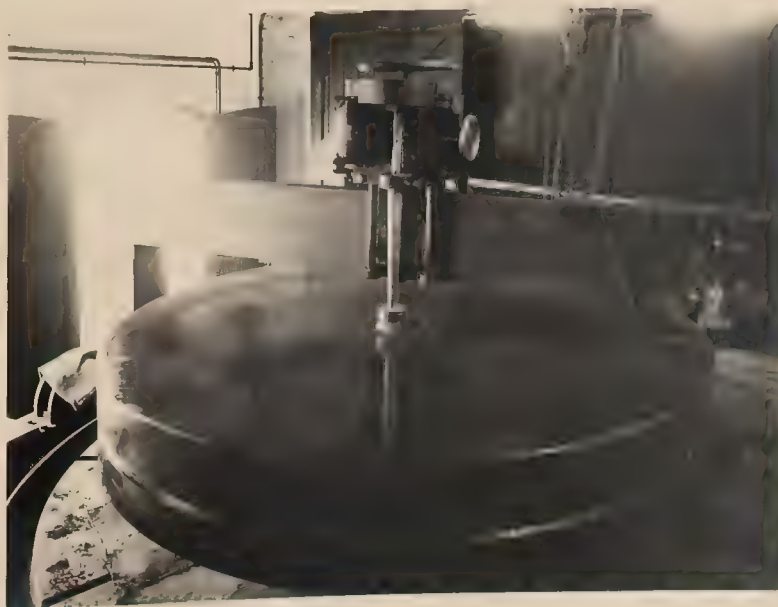
Grinding the Groove in the Circumference.

Two Views of the Great Mirror - May 1934



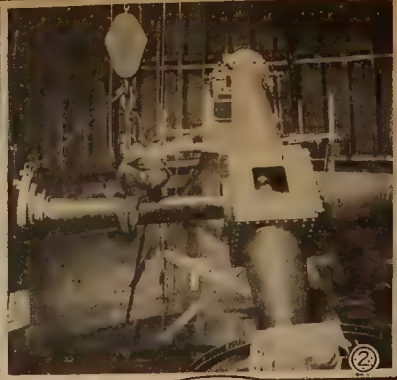
No 1045

Grinding out the Central Hole.



No 1046

Grinding out the Central Hole.



MOON BUT 50 MILES AWAY THROUGH GIANT TELESCOPE

Richmond Hill Instrument,
Second Most Powerful
In the World

FINISHED SHORTLY

The new David Dunlap Memorial
telescope at Richmond Hill is near-
ing completion.

The telescope, the second largest
in the world, and the first in the
British Empire, exceeded only by
one other in the world, that on
Mount Wilson, California, is pro-
viding wide-spread interest in the
astronomical world.

For three years the telescope has
been under construction. The ad-
ministration building, of cut stone,
is practically finished. The steel
dome, 21 feet in diameter, complete
to the 11-foot diameter which
houses the telescope, is finished. The
tube itself is a marvel.

The grinding of the mirror, per-
haps one of the loneliest jobs of all,
is well advanced. It has been in
progress for many months. The
mirror is not to be a mere one-
millionth of an inch.

The telescope presents a weird ap-
pearance. It might easily be taken
to be a mere pipe. The tube is 20 feet in
length and nearly seven feet in di-
ameter. It is a wide open at the top.
At first glance it suggests a huge
trench mortar, ready for action.

In the bottom of the tube lies the
great mirror, face up, reflecting a
circle of sky. It is about this costly
treasure of a mirror that all the
delicate machinery and expensive
plant of the new observatory centres.

The mirror is 74 inches in width
and nearly a foot in thickness.

The polar axis of the telescope is
22 feet in length and weighs a ton.
The declination axis moves freely.
It is the work-a-day axis of the tele-
scope on which the heavy tube turns.

The tube of the telescope is not
altogether unlike a cylinder. The
axis is made of a steel casting.
The axis is supplied on one side
with a large box for attachment to
a flange on the end of the de-
clination axis.

Canada's new telescope will have
about 10,000 eyepieces, and will
bring into vision more than 100,000
stars and thousands of those
faint universes which are so oc-
cupying the attention of the astron-
omers today.

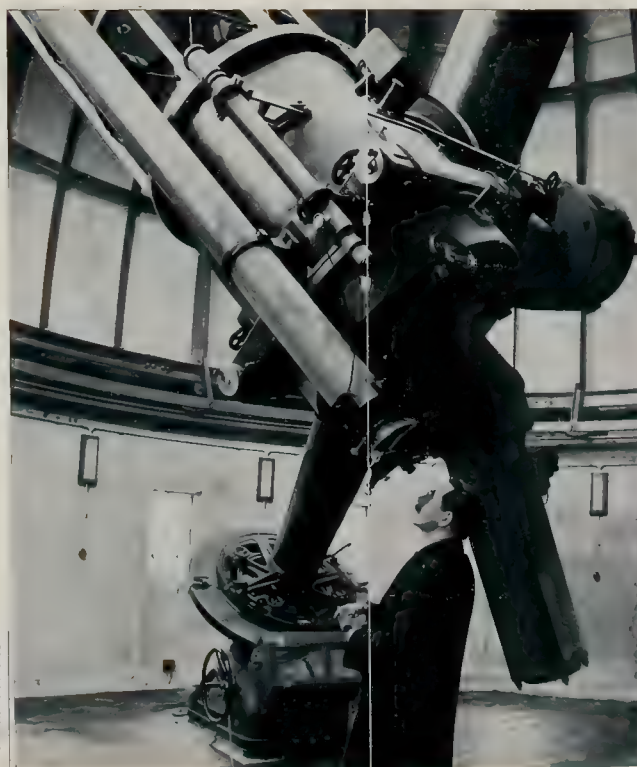
The moon, when it comes up over
Richmond Hill next September, will
look as though distant only about 50
miles. Mars, the nearest of the
planets, although about 35,000,000
miles away, will seem much closer.
The stars from now on it will be about
20,000 miles away from us.

The new telescope will penetrate
14,000,000 light years into space.
How far is that? Light travels 186,000
miles per second. Nobody can
comprehend it.

Prof. A. C. Chan is a busy man
these days with the new telescope
nearing completion.

NEW DUNLAP MEMORIAL OBSERVATORY AT RICHMOND HILL NEARS COMPLETION

The new Dunlap Memorial Observatory at Richmond Hill, which has been under construction for the past three years, is now nearing comple-
tion. The administration building, of cut stone, is practically finished.
The steel dome, 21 feet in diameter, which houses the telescope, is finished.
The tube itself is installed. The grinding of the great mirror is well
advanced. Photos show: (1) The new 36-inch reflecting telescope at the
Royal Observatory, Greenwich; (2) a view of the interior of the dome of
the Dunlap Memorial, showing the massive polar axis of the telescope.
The axis is 22 feet in length and weighs nine tons. (3) The steel dome of
the observatory, showing the declination axis being hoisted within. (4)
The tremendous axis itself.



The Sphere.
April 21, 1934

A MODERN MECHANICAL MARVEL
Another view of the new Greenwich
telescope which has been built by
the firm of Sir Howard Grubb,
Parsons and Co. of Newcastle on
Tyne. When "set" the instrument
remains pointed at the spot of
heavens under observation by
this keeps the polar axis turning
at the rate required to compensate
exactly for the earth's rotation. The
telescope is designed for use in what
is known as Cassegrain form, which
has many outstanding advantages
when used in the analysis of spectra.

Photographs taken May 8, 1934



*Digging the Trench for the H.E.
P.C. Cable. (Looking E.)*



Another View H.E.P.C. Trench. May 8, 1934



Digging Trench for Water Pipes.



*Dominion Bridge Truck removing
planks loaned from Oct. 1933*



View of Buildings from East.



Repairing the Lane-Looking West.



*Dr. J. S. Plashett, Director, Dominion Astrophysical Observatory
Victoria B.C. May 1934*



*C. A. Chant. Photograph by J. S. P.
May 10, 1934*



*Truck-load of 4-inch Iron Pipe for Water Main.
May 10, 1934*



Dumping Cinders on the Road. May 15, 1934



*Sawing up Rails, Posts, etc. May 15, 1934
Haldane Finney
(Garage behind)*



*Digging the Trench for the Water Main. Looking E. near Bayview Ave.
May 15, 1934*



Digging Trench for Water Main. Rear of Administration Building - (Looking E.) May 16, 1934



*Grading around Administration Building.
carrying in Books May 31, 1934*



Grading. May 31, 1934 Jordan-Roberts Truck



Another View of the Grading (Looking S.) May 31, 1934



David Dunlop Observatory,
Richmond Hill,
May 31, 1934

Dear Mrs. Dunlop, —

This is the first letter written by me in my office in the new observatory, and I think it should be addressed to you.

There is considerable work about the place this morning. The electricians are in the basement connecting up the big water supply pipes. Down at the well, the electrical connections are being made with the Hydro line in regard to the. The trench in front of the building for electrical power to come on under ground, is dug somewhat beyond the big dome. I am told that it is being made so deep that a 7' range



Mrs. D. & Dunlop,
73 Highland Ave.,
Toronto.



First Letter Written from the Directors Office.

THE GLOBE, TORONTO, SATURDAY, JUNE 2, 1934



THE HOME FORUM

"Many a night I saw the flames dance through the mellow shade, glitter like a swarm of fireflies tangled in a silver braid."
—Tennyson.

Dear Homemaker: I have enjoyed the recent letters on astronomy, for I, too, am a lover of stars. When a child I read and reread *Stories of Starland* by Adelaide Proctor, these stories being interspersed with lovely little poems, such as "Wynken, Blinken and Nod," "The Night Has a Thousand Eyes," etc. I have been interested in the subject ever since, and of late have been trying to find a few new constellations each summer when in the North country, where the night has not only a thousand eyes, but it would seem, a billion. I take with me Professor Chant's text-book, "Our Wonderful Universe," which is easy and fascinating reading, and from the Northern Cross, very conspicuous in the Milky Way, I can now find Lyra, Hercules, so like a flower pot; Corona, the little crown; Boötes, and then the glorious Arcturus, the brilliant orange star, in the southwest in September.

Did any one listen over the radio last June to the opening of the Chicago Exposition, when the ground, and buildings, were flooded with colored lights, undecipherably beautiful, the power being generated from a beam of light by the great Yerkes, Olcott story in Wisconsin, from the star Arcturus? It was a wonderfully interesting hour, and I seemed to actually experience a physical shock when the contact took place.

From one corner of the great square, Pegasus lies up in the southern sky, I can now find, too, the Nebula Andromeda, the only one of these great masses of stardust to be seen without the aid of a telescope.

So, my one is interested, I am, and in the great new observatory under construction at Richmond Hill. May I quote a little from an article in the *Sphere* of April 21 on telescopes in general and this one in particular?

The seeing part of the human eye is very small. "How minute are our instruments" in comparison with the celestial universe! once exclaimed El Karaki, an Arabian astronomer, who built a great observatory in Cairo in the twelfth century. He could see about 2,000 stars. Newton's little telescope increased this a hundredfold, making visible about 500,000 stars; Hubble, the sky as a whole. Canada's new telescopes will have a 30,000 eyepiece, and will bring into vision more than 100,000,000 stars, and thousands of those little island universes which are so occupying the attention of astronomers today. The moon, when it comes up over Richmond Hill next September, will look as though distant only about fifty miles.

Nobody knows what is in store for the great new David Dunlap memorial telescope, except that it will be a marvel!

I hope to come again to soon but I wanted to recommend Professor Chant's interesting and inexpensive book with keys to the summer and autumn stars, to all interested, and may I sign myself as

Arised.

Thank you, Arised, for this fine letter.



During construction water had been brought from Richmond Hill.

The last Tank Load of Water, brought from Richmond Hill.
June 12, 1934



Pump House in Course of Construction. June 12, 1934



Hydro Construction Party which removed the poles.
June 12, 1934



View obtained from the top of a Hydro Pole, just before the pole was removed. June 12, 1934

Hydro Construction Party, June 12, 1934
During the erection of the Buildings, power was supplied on poles. Then the wires were put underground, and on June 12, the poles were removed.

University of Toronto
DEPARTMENT OF ASTRONOMY

DAVID DUNLAP OBSERVATORY
RICHMOND HILL P.O., ONT.

July 25, 1934

Dear Mrs. Dunlap, -

I came down on Monday, as I proposed, and yesterday called at Rouse and Mann's, picked up Mr. Robson (who appears to be their chief authority on art) and we two went over to Highland Ave., to examine Mr. Massey's Book. With it before us we discussed your book for an hour, and in that time made our ideas definite.

Robson has the oak covers and has the style of binding, the hinges and lock, etc., definitely in his mind. He will send me some sketches and I will take them over to you and we can decide on what is best. I will also discuss with you the best method of arrangement, what all to include, and any other details which you may think of.

This afternoon I called at McDonald and Willson's, and they have some fixtures which will probably be suitable for your room and the library. It is astonishing how interested the people are in the Observatory. At McDonald's I met Crane (one of the head ones) and Black, and they insist -

Letter discussing this Album,
electric fixtures and other matters.

University of Toronto
DAVID DUNLAP OBSERVATORY
RICHMOND HILL P.O., ONT.

Indisposed - Excusing

Mrs. L. A. Dunlap

Zion-Alba Farm,

Godmorden P.O.,

Ont.

University of Toronto,
Yonge St.,
Richmond Hill, Ont.

SAMPLE COPY

From a Newspaper Wrapper, received July 25, 1934
(The University moved up Yonge St.!)



July 27, 1934
The Summer School in Astronomy visits the Observatory.

July 24, 1934



The Pump House. Aug. 23, 1934

Evening Standard
LONDON

Empire's Biggest Telescope Weighs 50 Tons

TORONTO, Wednes. day.
WHEN the new Dunlap Observatory is opened at Richmond Hill, north of Toronto, shortly, it will contain what is claimed to be the biggest telescope in the British Empire.

The observatory is being given to the University of Toronto by Mrs. D. A. Dunlap, a memorial to her late husband, a Canadian mining man.

The dome and building are practically completed, and part of the telescope is now being put in position. An English telescope firm has had charge of the main reflector, which will soon be installed.

The weight of the telescope and its 74-in. reflector is about 50 tons, and the weight of the steel dome about 80 tons.

DOMES ON ROLLERS

The dome is carried on 24 casted rollers of 2-in. diameter, mounted in ball bearings, and running on a rail. Sixteen pairs of lateral roller bearings on the inner and outer edge of the rail keep the dome in position.

Two segmental platforms, the lower one at the base of the opening and the upper one at the back, 16 ft. higher, are fixed inside the dome.

The centre section of the telescope tube itself is a steel casting seven feet in diameter and weighs 5½ tons.

Just above the lower flange, the casting swells out to 8 ft. 7 in. to accommodate an iris diaphragm for use in stellar photography.

The shutter of the camera opens to 6 ft. 2 in. and is operated by a hand wheel. It is expected to reveal new secrets of the stars.

The Dunlap Observatory is being constructed under the supervision of Professor C. A. Chant, head of the astronomy department of the University of Toronto. — Reuter.



Mr. Cyril Young, Manager of Sir Howard Grubb, Parsons & Co., putting the dark slide into the Newtonian Breechpiece. (See pages 114, 115)



Science reaches into space to discover new stars whose light started toward us aeons ago

ON a clear, dark night, when the rest of the world seems hushed in sleep, how many stars can you see? They seem innumerable, it is true, but actually all that are visible to the unaided eye are some 5,000 or 6,000 and of these not more than 2,500 are visible at any one time. These were all the stars the ancients dreamed of and these are all the stars of today can know, and all that even modern man knew about until the advent of the telescope.

Galileo's crude telescope of the 17th century, enlarged man's vision to add hundreds of thousands of other stars. Since that epochal day, every increase in the power of the telescope has brought out new stars that otherwise would have been invisible. Yet the great galaxy of stars which we are able to view through the ordinary telescope constitute but the nearer fringe of hosts of heavenly bodies so numerous that the mind is staggered by the vastness represented.

At first, in his arrogance or ignorance, man thought of his own planet as the centre of the universe. The world was the most important thing in space and dominated all. The sun, the moon and the stars revolved around the world. Once a day the lordly sun made his trip, in unquestioned obedience to the will of the centre of the universe. From the time man first became cognizant of the apparent motion of the heavenly bodies this idea persisted, persisted right up until a few hundred years ago, when the theory of Copernicus that the earth revolved around the sun was accepted, and the invention of the telescope finally revolutionized the study of astronomy.

Dr. C. A. Chant, Professor of Astrophysics at the University of Toronto and Director of the David Dunlap

Observatory which is to open early next year expressed the idea quite clearly when he stated that in early times "the universe was believed to be but of small extent. The heavens were just overhead and the gods often came down to mingle with mortal man."

"Today," he added, "we know that the sun itself is a star giving out light and heat of itself and the earth is one of a number of much smaller bodies called planets revolving about it. The size of their orbits when expressed in miles seem very large, but when we find out the distances to the millions of other stars which appear as bright dots on the sky

we realize that our solar system occupies but a bright pinhead of space, while the rest of the universe is inconceivably remote. The distances from star to star, are not infinite although they are extremely great.

"THE nearest star (outside the sun) is the Centaur, which is 270,000 times as far as the distance of the earth from the sun or 25 million million miles. We express this distance in another way. The speed of light is 186,000 miles per second or 11,000,000 miles per minute. While light requires only eight and a third minutes to come from the sun to the earth, it takes over four years to come from Centaur, which is therefore four light years away from us."

Four light years away! Can the human traveller comprehend it? The fastest trip ever made around the world by aeroplane aided by every modern means of travel, took some sixty days to accomplish. Yet light travels the same distance seven times over with every tick of the clock. And light travelling four years just manages to make the trip between Centaur and the earth.

If the nearest star is four light years away, how about the others? The distances from the earth of some 2,000 have been measured and they are found to be up to 500 light years away.

Gaze on any one of them. They are quite visible. (Turn to page 40)



Grinding the edge of the huge reflector lens

by Allan Strathglass

Top left. The Dunlap Observatory, that will house the second largest telescope in the world.
Left. The great telescope set up in England before shipment to Canada.
Right. The iris diaphragm.

HOW MANY WORLDS CAN YOU SEE?

(Continued from page 7)

new data might. Here is a star that looks down on us tonight. We see it brightly. But the light we see is not the light emitted by the star itself, but light which is now reaching us on its voyage through space when Drake was battling the pirate Armada.

But most stars are even farther away. There is another, the light from which started to earth after Alexander of Macedon was executed. It is so far that it will take more than 2,000 years to reach us. But Alexander died only 2,300 years ago, a comparatively recent date when compared to the distances which separate them from the stars.

Two great majority of stars are much farther off than this," according to Professor Chant and other astronomers. The globular star clusters for instance, range all the way from 10,000 to 220,000 light years away. The nebulae, which are so far away that they are hardly to be seen with the best of telescopes, are much more distant. In fact, the nearest is 400,000 light years away. And one recently discovered spiral nebula is assigned a light distance of 1,150,000 years.

Which means that those stars which are 50 light years away, some 1,500 million miles, are the nearest stars. A billion miles for instance is such a distance as to be almost beyond our mental grasp. But that is only a brief footnote compared with the distance to those stars which science has located a million times as far away again.

THE Neanderthal man, however, is believed to be the first human type on earth was supplanted by the Cro-Magnon some 50,000 years ago millions of years after the light which is just now visible to us started on its journey to earth. The world itself is believed by many scientists to be but a little over 1,500,000 years old yet the light from that newly discovered nebulae mentioned, started on its journey to earth 115,000,000 years before the earth was born.

INTERESTING conjecture is possible when considering these great distances. It is believed that the stars are moving away or appar-

ently moving away from us at an almost incredible speed. So that when one gazes at these distant stars or star clusters through a modern astronomical telescope, he is seeing the star or star cluster, not where it is but where it was anywhere up to 135,000,000 years ago. Millions of years ago that particular star or star cluster may have died out, or some catastrophe may have happened to it and yet it may be millions of years before that could possibly be known to man.

Or put it this way. If something happened to Centaur, the nearest star, tonight, say it blew up, that fact would not be known to even the astronomers of the world for four years.

Astronomers are probably gazing tonight at some phenomena which passed into history long before the earth broke away from the sun.

GALILEO started all this, with his invention of the telescope. And it brought him plenty of trouble because he insisted that the earth went around the sun. The Holy Office of The Inquisition denounced Galileo, affirming:

"The proposition that the sun is the centre of the world and immovable from its place is absurd philosophically, false and formally heretical because it is especially contrary to the Holy Script."

Galileo's telescope consisted of a simple convex lens about two and a quarter inches in diameter. At that, it increased the vision of the naked eye to such an extent that it was sufficient to add hundreds of thousands of stars to the two or three thousand previously seen.

FROM Galileo's simple arrangement of 300 years ago to the great reflecting telescope of today is a far cry. All telescopes until recent years were modifications of the refracting type as first introduced by Galileo. The modern refracting telescope today is built on the same principle. But the reflecting telescopes depend on mirrors for their efficiency. The light is collected on a great concave mirror, is in turn reflected to another mirror and the resulting image focussed on an eye

piece. The new telescope enlarges our vision, not by a paltry few hundred thousand stars, but by millions. Such a one, adding 1,000,000,000 stars and thousands of island universes to those previously seen, is rapidly nearing completion at Richmond Hill, near Toronto.

This is the David Dunlap Observatory, provided as a memorial to the late David A. Dunlap of Toronto.

This observatory will be the finest of its kind in the British Empire and will house the second largest telescope in the world, second only to the giant telescope on Mount Wilson in California. The observatory and administration buildings are completed, the tube of the telescope is installed and completion only awaits the installation of the great 74-inch reflecting mirror, at present undergoing polishing and silvering in England.

The building is designed to be warm in summer and cold in winter—in fact, the idea is to maintain the same temperature inside as prevails outside.

It is about the telescope itself that interest in both the lay and astronomical worlds concentrates. The great tube, thirty-one feet long and seven feet in diameter, is made of steel. With the shutters open and the tube pointed at the sky, it reminds one of a Gargantuan field gun trained on some far-away planet or star. As a matter of fact, to carry the simile to its logical or illogical conclusion whichever you wish—the range of the telescope is some 135,000,000 light years. In other words, it brings into view heavenly phenomena so remote that it takes 135,000,000 years for the light from them to reach us.

WHILE the tube is spectacular in appearance, the main part of the telescope, that upon which rests its successful operation, is the great mirror which is to be fitted into the base. All the delicate machinery and expensive plant of the observatory centre about the mirror. This mirror is made of a great circular block of glass, six feet two inches in diameter and a foot thick. When completed it will weigh approximately 5,000 pounds—two tons and a half. In the centre is a circular hole a foot in diameter, and the completed mirror resembles nothing so much as a massive, polished grindstone.

The casting of the three-ton disc of glass for the mirror was a piece of work which offered great technical difficulty. In the first place, it had never been done on such a scale before. It could not be done in Canada because no facilities for such work exist. The building of a plant capable of doing the work would certainly not justify the expense because the needs of the whole world will hardly permit of more than one or two such plants. As mentioned, such a large glass casting had never been made, though larger castings had been made in two parts.

When scientists first attempted the work, they found that the only known available material which satisfied all their requirements was fused quartz glass. But quartz required enormous temperatures for casting. Finally experiments turned on pyrex and it was of this material that the 74-inch disc of the mirror was finally cast.

But the casting of the disc—remarkable feat as it was—was just

one step in the production of the huge convex mirror for which it was intended. It had to be allowed to cool very slowly, else it might crack or crack, which would mean that it would be useless. It was six months before the makers ventured to strip away the matrix. When they did so, they were delighted to find that their calculations had proved correct. The great disc was found suitable.

But this was only the start. The disc had to be ground and polished correctly to one-millionth of an inch. And, for that work it was shipped to England, and there, for almost a year, it has been in the process of grinding prior to silvering. If you think this is slow work, please remember that grinding is only permitted for fifteen minutes a day, in case the increased temperature caused by the friction should distort the mirror.

One other feature of the telescope deserves mention. That is the iris diaphragm which is used for protection of the mirror and for the changing of the aperture for the admission of light. This diaphragm is built on the same principle as one in the shutter of the ordinary folding camera, only it is on a comparatively enormous scale, having a maximum opening of six feet two inches.

The Dunlap Observatory telescope is the second largest in the world, being inferior in size and—if one may use the expression—penetrating power, to the Mount Wilson telescope only.

But it is still only a beginning. Already there is talk of a 200-inch reflector. If the 74-inch Toronto reflector can reach out into space for stars 135,000,000 light years away and find them, how many more may we yet see when we have tapped the last resource of human ingenuity.

Canadian Telescope Records First Infra-Red Photographs Of Perseid Meteors at Toronto

Great Dunlap Mirror of 100,000 Eye-Power Searches Secrets of Stratosphere as Told by Annual Wanderers

Special to The Christian Science Monitor

TORONTO—At the Dunlap Observatory of the University of Toronto, Richmond Hill, the recent shower of Perseid meteors was for the first time photographed on infra-red plates.

On each of two nights Dr. P. M. Millman of the Department of Astronomy reported that astronomers counted 500 meteors, although they were disappointed on the third night.

The Perseid meteors, which are seen each year during August, travel at a height of about 70 miles above the earth, at an estimated speed of about 40 miles a second, it was explained. The meteors are usually fragments of some comet broken up by the solar system, traveling in orbits around the sun, and they may be seen when they intersect the path of the earth about the sun.

Pioneer Photographs

A photograph of these meteors in the infra-red has never before been obtained, Dr. Millman pointed out. The interest in them lies in what they may reveal of conditions in the upper stratosphere, 40 miles above the surface of the earth, as well as in the scientific interest with regard to the meteors themselves.

The telescope at the Dunlap Observatory, which has been under construction for the past three years, is the second largest in the world, and the finest in the British Empire. It is second in size only to that on Mount Wilson, California.

It is housed in a steel dome, 61 feet in diameter, completely equipped. At the bottom of a tube, 30 feet in length, and nearly seven feet in diameter, which is wide open at the top, lies the great mirror, 74 inches in width and nearly a foot in thickness. The polar axis of the telescope is 22 feet in length and weighs nine tons. The declination axis moves freely.

With approximately 100,000 eye-power, the new telescope will bring into vision more than 1,000,000,000 astronomical bodies and the moon

will look as though it were only 50 miles away. Mars will seem much closer than its 58,000,000 miles from the earth. The new telescope is said to be able to penetrate 140,000,000 light years into space.

Christian Science Monitor.
(Boston) Aug. 24, 1934

THE TELESCOPE

An Illustrated Magazine of Astronomy

Vol. 1

AUGUST, 1934

No. 5

Canada's Third Large Observatory

The New David Dunlap Observatory of the University of Toronto.

By P. M. MITCHELL
Dept. of Astronomy, University of Toronto

POINCARÉ remarked that, had man been unable to view the stars, he would still be living in an age of barbarism. Though some may consider this an extreme view, certain it is that man's first conception of natural law and order arose from his observations of the heavenly bodies, and it is equally certain that such observations have played a major part in the guidance of scientific thought from the dawn of history to the present day.

Important as theoretical work is, the greatest need in the development of a science like astronomy is for the accumulation of a vast amount of observational data. During the last twenty-five years the North American continent has easily led the rest of the world in this respect and most of the work has been done at a few great observatories located in the United States. Canada, in the past, has had only two important research observatories, the Dominion Observatory at Ottawa and the Dominion Astrophysical Observatory at Victoria. These are both operated by the Dominion Government. Before the end of 1934, however, regular work will be commenced at Canada's third large observatory, the David Dunlap Observatory, near Toronto.

Toronto has long been an amateur astronomical center, the

Toronto Astronomical and Physical Society having been founded in 1800. Later, this developed into the Royal Astronomical Society of Canada which has active centers in all large cities in Canada and members throughout the world. The astronomical department of the University of Toronto came into being about thirty years ago, mainly through the efforts of Professor C. A. Chant, who has been the head of the department since its formation. Toronto had no good observatory, however, and the need for one has been felt for some time. It is through the generosity and farsightedness of Mrs. Jessie D. Dunlap that the new observatory has been made possible. Mrs. Dunlap wished to erect a lasting and fitting memorial to her husband, Mr. David A. Dunlap, who had been a keen amateur astronomer and much interested in the establishment of an observatory in Toronto. Learning from Professor Chant the need for a modern research institution at Toronto, Mrs. Dunlap offered to the University of Toronto, in 1927, the funds for the purchase of land and the erection of an observatory to be known as the David Dunlap Observatory.

In planning a new observatory the site is all-important and tests were made of the atmospheric conditions at various localities near Toronto. The final choice was the

AUGUST

THE TELESCOPE

1931

top of a slight rise about fifteen miles directly north of the University. The elevation is 800 feet above sea level, the highest point for some distance, and while the atmosphere is very slightly more transparent further north, the difference is so slight that it is not enough to balance the increased inaccessibility which would result by locating the observatory at a more distant point. The observatory property consists of 170 acres which, at present, is mostly open meadow. An arboretum has been planned by the department of forestry and, when this has been completed, the trees will eliminate, to a large degree, the poor seeing that results from the heat which rises after sunset from ground which has been exposed all day to direct sunlight.

Excavation for the administration building was commenced in 1932 and the corner stone was laid

by Mr. D. Moffat Dunlap, son of Mrs. Dunlap, on September 10, 1932. This building is of white stone and is surmounted by three copper covered steel domes, the largest being twenty-five feet in diameter. The entrance hall is lofty and is lined with Italian travertine, the material of which the Colosseum in Rome is constructed. The dedication tablet, in marble, faces the visitor as he enters, and an exceptionally artistic stairway to the second floor winds around three sides of the entrance hall. The administration building contains offices for the staff, a lecture room seating 150 persons, library and stack room, photographic dark room and laboratories, clock room, and a completely equipped machine shop.

A 10-inch reflector is being mounted in the south dome of the administration building. This instrument was constructed in its



The Administration Building of the David Dunlap Observatory

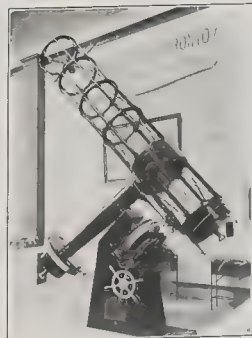
AUGUST

THE TELESCOPE

1934

entirety by Professor R. K. Young of the department of astronomy at Toronto. The mirror is of Pyrex and the telescope is driven by the Gerrish electrical drive. No instruments have yet been secured for the remaining two domes of the administration building but it is hoped that a photovisual telescope of about 12-inch aperture will be mounted in the center dome and a battery of fast photographic telescopes in the north dome.

The chief instrument of the observatory is a 74-inch reflector, constructed by Sir Howard Grubb, Parsons and Company of England. This telescope is housed in a 61-foot dome located some fifty yards north of the administration building and it will have, when completed, a light grasping power sec-



The 74-inch Reflector of the Dunlap Observatory

ond only to the 100 inch telescope at Mt. Wilson. The instrument is now mounted and lacks only the optical parts. The main disk is of the special type of Pyrex glass developed for reflecting telescopes by the Corning Glass Works of Corning, N. Y. Its overall diameter is 76 inches, its thickness 12 inches and it weighs in the neighborhood of 5000 lbs. It is now being ground and polished by the Parsons Co. at Newcastle-on-Tyne and will probably be ready for use by the end of 1934 or early in 1935.

The 74-inch has an English equatorial mount, very similar to that of the 72 inch at Victoria. The tube of the telescope is of duralumin, which saves considerable weight, and an iris diaphragm is located directly in front of the large mirror, a new departure for reflect-



The Entrance Hall of the Administration Building

[60]

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The 61-foot dome of the Dunlap Observatory under construction

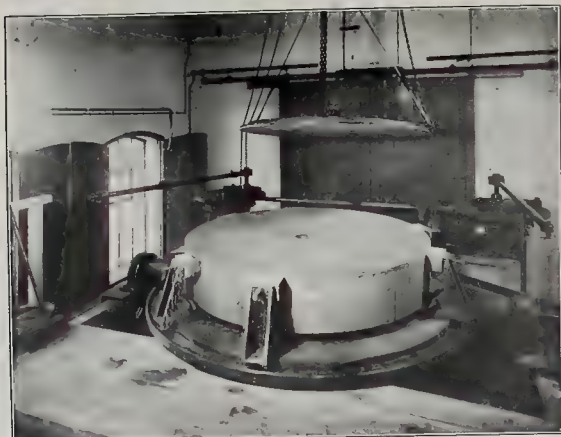


The 10-inch reflector of the Dunlap Observatory



The interior of the 61-foot dome of the Dunlap Observatory

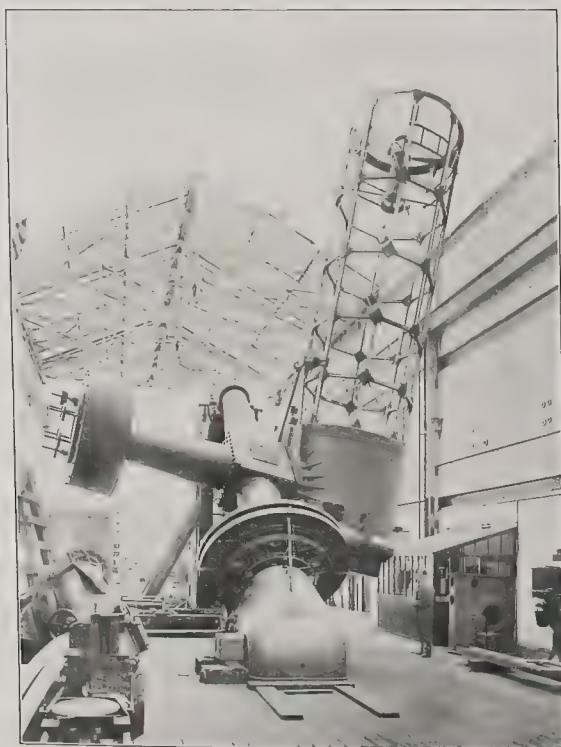
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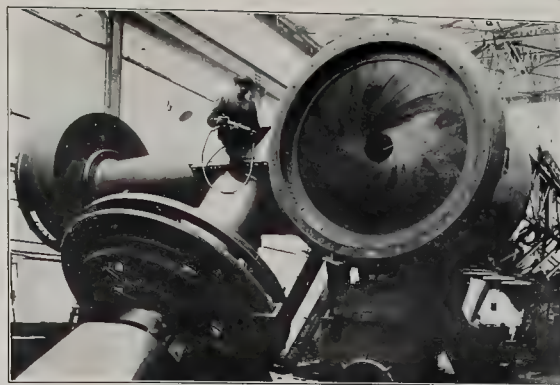
The 74-inch disc, on the grinding table. The rough grinding of the front surface is completed



The pier of the 74-inch reflector mounting before the dome had been erected
[62]



The 74-inch reflector in the shop at Newcastle-on-Tyne



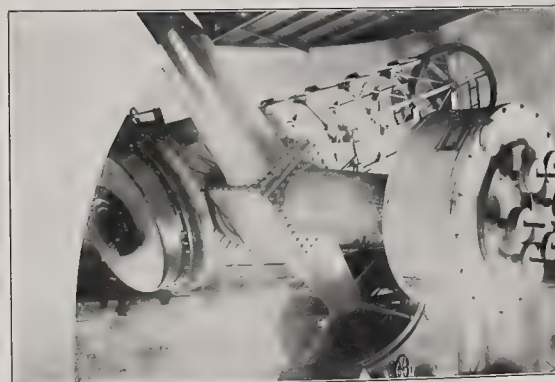
The tube of the 74-inch reflector in place, showing the iris diaphragm at minimum aperture

ing telescopes. The instrument is mounted on ball bearings and is supported by two reinforced concrete piers which extend to a depth of 25 feet below the surface of the ground. These are hollow and contain several small rooms which will be excellent for housing seismographs or accurate clocks. A small dark room is located in the upper part of the north pier.

The building housing the large telescope is constructed of metal, the walls being double to permit the constant circulation of air within them. The dome is covered with copper laid over a double layer of heavy agasote, a hard paper product. The observing platform for use at the upper end of the telescope is mounted in a different manner to that employed with any other large telescope. It moves up and

down a curved track the whole of which may be moved from one side of the dome to the other which thus facilitates placing the observer in any desired position, no matter how the instrument is placed. This platform is electrically controlled, as is the whole telescope as well as the dome and the shutters. Even the focusing of the cassegrainian mirror is performed by a small electric motor. The makers have spared no pains to make the telescope convenient in operation and efficient in performance.

The telescope is driven to follow the stars by a weight clock so geared that, in effect, the clock is being electrically wound all the time it is running and thus the weight remains in approximately the same position at all times. This



The 74-inch telescope as mounted at the David Dunlap Observatory

clock is synchronized with the sidereal master clock by what is known as the Grubb drive, an electrical synchronization on a somewhat different principle to that of the Gerrish drive.

A special room for resilvering the great mirror is located on the ground floor of the 74-inch dome. The mirror may be lowered into this room from the observing floor by means of a counter-weighted elevator which runs in a shaft extending 25 feet below the ground level. It is hoped eventually to have the surface of the 74 inch mirror aluminum coated, but no equipment for handling a disk of this size has yet been developed. The two secondary mirrors of the large telescope and both mirrors of the 10 inch reflector are now be-

ing coated with aluminum through the courtesy of the California Institute of Technology.

The work carried out with the 74-inch telescope will probably be largely spectroscopic, with particular stress laid upon faint objects, where a great field of investigation lies awaiting a telescope of this type. With many new large telescopes either in operation or under construction, the next fifty years should see the solution of a large number of astronomical problems which can only be attacked with great telescopes. Happily the gift of Mrs. Dunlap has made it possible for the University of Toronto to be one of the institutions in the advance guard of our attack on the astronomical problems of the future.

NEW ZEALAND ASTRONOMICAL SOCIETY (INC.)

MONTHLY NOTES.

1934 SEPTEMBER

DOMINION OBSERVATORY, WELLINGTON, N.Z.

No. 93

ANOTHER GREAT TELESCOPE FOR CANADA

The magnificent 72 inch reflector at Victoria, B.C., the second largest in the world and the largest in the British Empire, will soon be eclipsed by one of 74 inches aperture which is being presented to the University of Toronto by the Widow and son of the late Mr. David Dunlap of that City, and in his memory. It is pleasing to note that such dignified memorials as this, are beginning to take the place more frequently of the ornate and useless structures so often erected in memory of prominent men who have passed away.

It is a source of gratification also to see that the huge instrument is a British production, it having been made at the works of Sir Howard Grubb, Parsons and Coy., Newcastle-upon-Tyne, and the observatory building and dome at Darlington by another English firm. Both were shipped to Canada and are now erected and awaiting the main mirror which is ~~now~~ almost completed. This is also being made by Sir Howard Grubb.

The mounting is of the English or composite type and the weight of the moving parts alone is 35 tons of which the great mirror weighs 2. The driving mechanism is of an improved type that has been fitted to only one other large instrument - the new 36 inch reflector at Greenwich. The difficulties involved in driving such a heavy instrument with the extreme accuracy required for long exposure photography have been aptly summed up in the following interesting manner by Dr. R.K. Young, Pres. of the Royal Astronomical Soc. of Canada. "A drift of the image of 1/1000 of an inch during the course of a photographic exposure would seriously affect the definition of the picture. The engineering difficulties confronting the telescope maker in thus keeping the tube so accurately pointed are about equivalent to keeping a gun trained on a target an inch in diameter at a range of 20 miles while the target is moving at the rate of 5 feet per second." The driving mechanism is electrically wound and its speed will be automatically controlled every second from one of the observatory clocks.

All the requisite operations of setting, clamping, focussing and slow motions are made electrically under push button control. Some members of this Society doubtless long for the day when push buttons will be standard equipment on home built reflectors. Then it will be as easy to control the movements of a 6 inch telescope as to set, clamp, and drive the 35 tons of a 74 inch. A feature of the instrument is a large iris diaphragm which is mounted in front of the great mirror and can be opened from a minimum diameter of 12 inches to the full aperture of the telescope. It is claimed ~~that~~ that this is the largest iris diaphragm that has ever been made.

The tube, of the usual latticework construction, is fitted with interchangeable Cassegrainian and Newtonian secondary mirrors, each about 20 inches in diameter. The focal length as a Newtonian is 30 feet, and as a Cassegrainian, 111 feet.

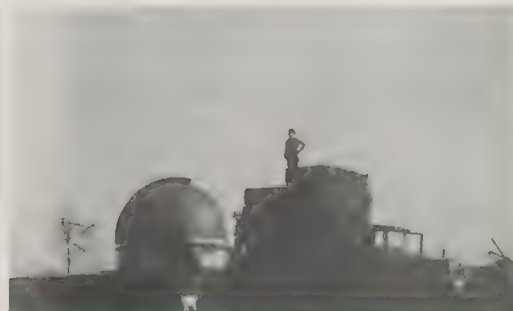
The mirror disc is 12 inches thick, of special Pyrex glass having a very tiny coefficient of expansion. It was produced at the Corning Glass Works which also cast the disc for the projected 200 inch instrument.

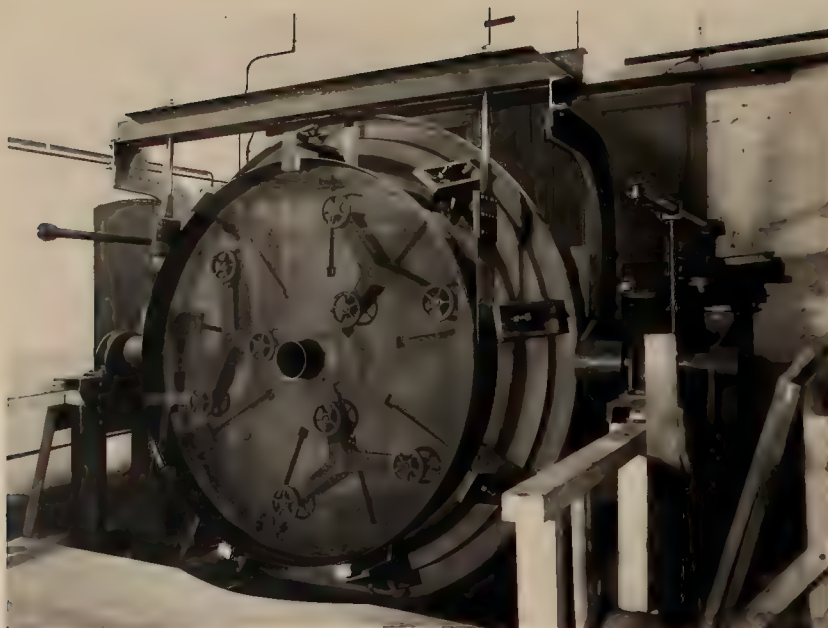
The Observatory itself is a round steel tower 21 feet high surmounted by a dome 61 feet in diameter with 2 parallel shutters opening 15 feet. It is located in a park of 177 acres on Richmond Hill a few miles to the North of Toronto.

The North American Continent will soon have more than half a dozen reflectors of 5 feet aperture or over, but as yet there are none at all comparable with any of these in the Southern Hemisphere where the possibilities of research with such instruments are almost incomparably grander. The Clouds of Magellan alone are worthy of study with the greatest telescopes that man can make.

ALLAN DRYCE.

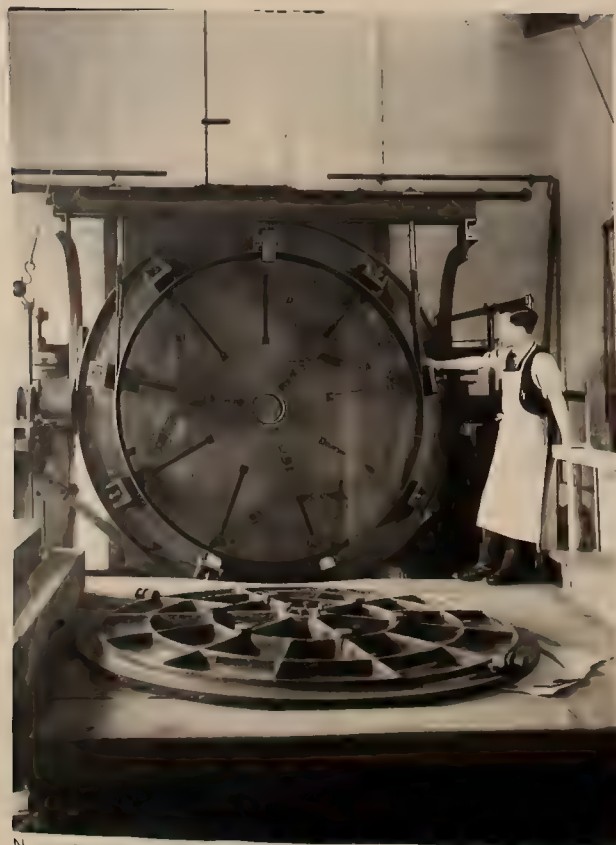
Hamilton.

Hamilton, N.Z.*Baling Hay**Sept. 19, 1934**(Hay crop short in 1934)**Loading Hay**Sept. 19, 1934*



No 1061

Main Mirror Tilted up for Testing (Side View Showing belt) September, 1934



No 1062

Main Mirror Tilted up for Testing (Front View) September, 1934



At the Front Door

October 26, 1934

*In the
Directors Office*



October 22, 1934



In the Entrance Hall.

October 26, 1934

In the Entrance Hall.





At the foot of the Stairs, in the Entrance Hall. October 26, 1934



View of Entrance Hall (Looking West)

October 22, 1934

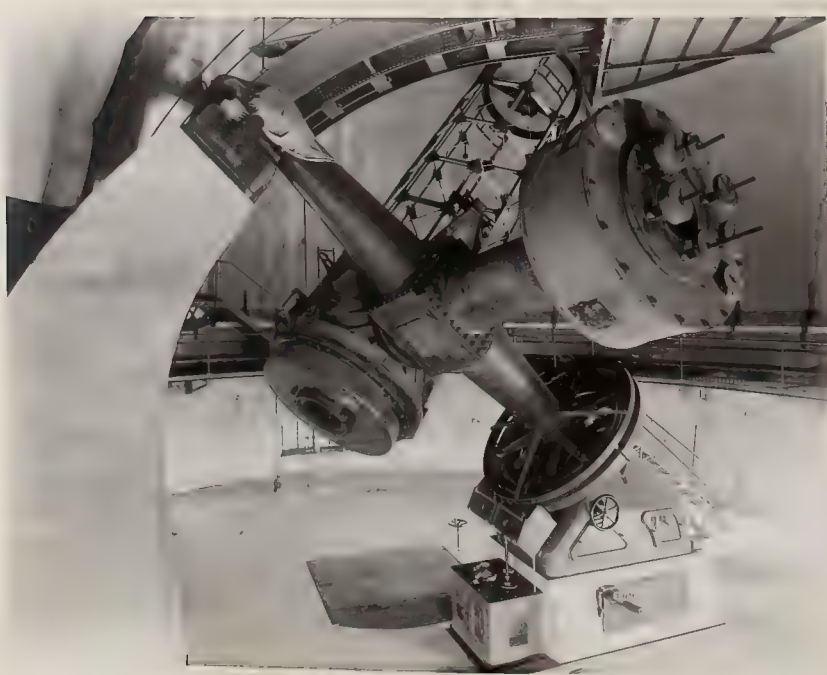


Entrance Hall from Stair Landing (Looking East)

October 22, 1934



*Mr. Dunlap
Photograph of Oil Painting by Joshua Smith in the
Library. Oct. 22, 1934*



The Great 74-inch Telescope. Oct. 22, 1934



David Dunlap Observatory

"The Heavens declare the Glory of God"

Christmas Greetings

from

Mrs. David A. Dunlap

*Hillside, Rosedale
Toronto*

Christmas Card, 1934

REVUE GÉNÉRALE HEBDOMADAIRE DES INDUSTRIES FRANÇAISES ET ÉTRANGÈRES

Administration et Rédaction : 5, rue Jules-Lefebvre, Paris (9^e).

LE NOUVEAU TÉLESCOPE DE 1^m 88 DE DIAMÈTRE
de l'Université de Toronto (Canada)

Un télescope de grandes dimensions vient d'être mis en service à l'Observatoire David Dunlap, généreusement créé à cet effet par la famille de ce nom, au voisinage de Toronto, à l'altitude de 250 mètres; comme pouvoir lumineux, il se classe immédiatement après le grand télescope du Mont Wilson, aux Etats-Unis, dont le diamètre d'ouverture est de 2^m 50.

Le télescope a été construit en Angleterre par MM. Howard Grubb, Parsons et Co, dont les établissements sont rattachés pour la partie mécanique aux ateliers de construction des turbines Parsons; nous nous proposons d'en décrire les particularités.

La figure 2 représente en élévation l'ensemble de l'instrument et de sa monture équatoriale; cette monture, ou pied paralactique, est du type anglais, le tube et son contrepiéd étant disposés de côté et d'autre de l'axe polaire ou horaire PP'. L'axe polaire est dirigé parallèlement à la ligne des pôles; à cet effet, le palier supérieur est porté par quatre vis et buté dans le sens transversal par quatre autres vis opposées deux par deux, au moyen desquelles on procède au réglage avant de serrer les vis de fixation. Le tube est boulonné à l'extrémité de l'axe de déclinaison.



de 70 hectares s'étend sur une hauteur, à Richmond Hill (fig. 4), à 15 kilomètres au nord de Toronto, dans un site particulièrement propice aux observations. En plus du bâtiment qui contient une couple du grand télescope visible à gauche de la figure 1, l'observatoire comporte un bâtiment d'administration auquel sont adjointes trois coupoles pour des télescopes beaucoup moins grands que l'appareil principal. Le bâtiment qui se trouve à l'extrémité droite est cylindrique, de 18^m50 de diamètre. Le plancher d'observation est, à 4 mètres environ au-dessus du sol, sur la longueur des galeries, intérieur et extérieur, tout le long d'un couloir de 7 mètres environ de hauteur.

naison DD', perpendiculaire à l'axe horaire; par une rotation convenable du système autour des axes PP', DD', on peut évidemment viser un point quelconque du ciel. Si alors on bloque l'axe de déclinaison par rapport à l'axe polaire, et si tout le système tourne autour de cet axe avec la vitesse angulaire du mouvement diurne, l'image de l'astre visé restera immobile dans le champ de l'instrument.

Le télescope proprement dit peut être agencé, suivant les circonstances, en télescope à vision latérale du type Newton (fig. 3), ou comme télescope à vision directe suivant la disposition inventée en 1672 par Cassegrain (fig. 4). Rappelons que le

A dix-huit contrepoids *b*, régulièrement répartis sur la circonférence, montés à l'extrémité de leviers articulés sur la cuvette; le bras court des leviers est en prise avec une patte, rivée sur une ceinture métallique souple qui entoure le miroir sans le serrer. Pour toute position du télescope, autre que la verticale, l'action combinée des contrepoids impartit au miroir une translation qui compense la flexion, due à la gravité.

Sur le bord de la cuvette est fixée la monture d'un diaphragme iris E, à vingt lamelles actionnées par volant à la main, qui permet de faire varier le diamètre d'ouverture entre 0^m 30 et 1^m 88, ouverture maximum; le diaphragme sert également à protéger le miroir lorsque le télescope n'est pas mis en service.

La virole centrale, également en acier moulé et nervurée, pèse 45 kg; elle est boulonnée à la cuvette. La portée du bossage est plane, mais un bourrelet en saillie sur l'embase de l'axe de déclinaison limite le contact à une ligne circulaire, de manière à limiter les échanges de chaleur pour prévenir dans le télescope, et par répercussion dans le miroir, une dissymétrie thermique, qui se traduirait pour ce dernier par un certain astigmatisme.

L'ossature octogonale, boulonnée à la virole, est en profilé

(2) Nous avons emprunté une partie des renseignements de cet article à une étude de M. Young publiée dans l'*Engineering Journal* (canadien) et à une série de trois articles parus dans l'*Engineering*, auquel nous sommes redevables, notamment, des figures 2 et 7.

Continued
at B.

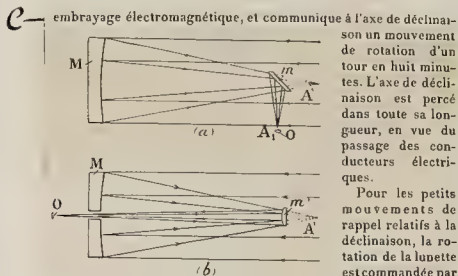


FIG. 3 et 4. — Principes du télescope à vision latérale du type Newton (a) et du télescope à vision directe du type Cassegrain (b).

sature octogonale; cette vis tourne dans un écrou, monté à l'extrémité du bras k , naturellement fou par rapport à l'axe de déclinaison, mais qu'une cale, actionnée électriquement au moment précis où le moteur de la boîte est débrayé, bloque sur l'anneau à gorge l solidaire de la monture de l'axe. Un moteur avec démultiplica-

teur à deux vitesses, fixé sur la virole centrale B, actionne par transmission à la Cardan et renvoi d'angle la vis i , et communique au telescope suivant la combinaison utilisée un mouvement angulaire de 15 minutes d'arc ou de 30 secondes d'arc à la minute.

L'axe polaire (fig. 5), dont la longueur est de 6^m 70 et qui pèse 9^k 5, est constitué par une boîte cubique, commune avec la monture

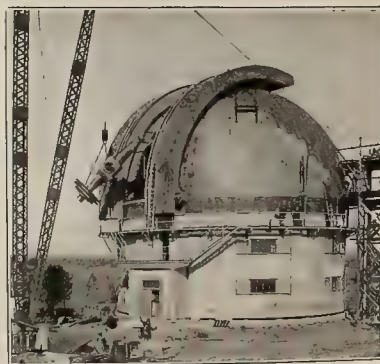


FIG. 5 — Vue prise pendant le montage de l'axe polaire.

de l'axe de déclinaison, par deux trompettes et par deux tourillons; ces pièces, en acier moulé ou forgé, sont montées sur des roulements à billes à portées sphériques, et prennent appui à la base sur une butée. Pour le dégrossissage de la visée, le mouvement de rotation est communiqué à l'axe polaire par la roue dentée n, clavetée sur le tourillon inférieur, en prise avec un pignon actionné à partir du moteur F par l'intermédiaire d'un embrayage électromagnétique; la manœuvre peut également être exécutée en utilisant les volants montés aux deux extrémités de l'arbre p. La visée étant supposée accomplie, il convient, comme il a été dit, d'interpéter à l'axe polaire la vitesse angulaire du mouvement diurne; le mouvement de rotation doit être rigoureusement uniforme, ce qui exclut la possibilité d'employer un mécanisme d'horlogerie de construction ordinaire, à mouvement saccadé.

saacé. Le mouvement de rotation d'accompagnement est communiqué à l'axe polaire par la roue tangente q , de 638 de diamètre, qui comporte 960 dents taillées dans une tôle de bronze rapportée; cette roue, actionnée à partir du mécanisme électrique G , est montée folle sur le tourillon, mais bloquée à la demande au moyen d'un petit mortet à l'échelle monté sur la toile de la roue n ou à la main en saisissant d'une roue de diamètre un peu plus petit montée sur le tourillon. La roue q est solidaire d'une roue de diamètre un peu plus petit montée au-dessus, laquelle entraîne, par six pignons régulièrement répartis sur la circonférence, le cercle sidéral s , que les volants v permettent de décaler par rapport à la roue q . La jante du cercle sidéral a 75 mm de largeur, elle porte deux graduations; l'échelle inférieure, en regard de repères fixés à la garde du cercle, indique le temps sidéral, et l'échelle supérieure, en regard de repères u fixés à l'axe polaire, indique l'ascension droite. Le mécanisme G étant mis en train au crépuscule, le

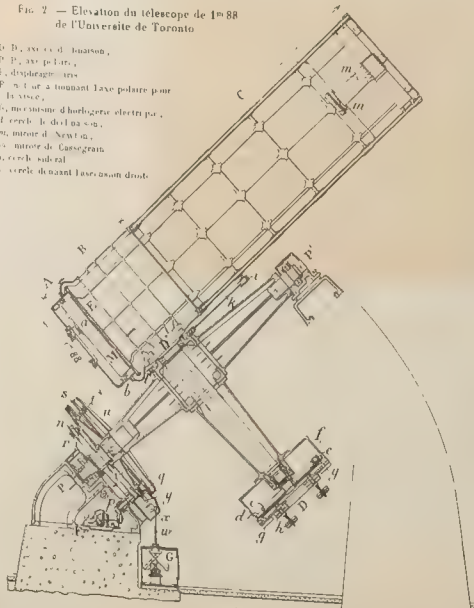
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LE GENIE CIVIL

TOME CV — N° 25

B — télescope « vison latérale » comporte un miroir plan diagonal m qui, au lieu de 45° sur l'axe, l'inclinaison est interposée avant la formation de l'image objective A' tournée par le miroir parabolique M en une image réelle A_1 qui est observée au moyen de l'oculaire O fixe dans la paroi latérale de la monture, dans le télescope « vison directe », le miroir objectif concave M est associé au petit miroir convexe auxiliaire m' qui transforme

FIG. 2 — Elevation du télescope de 1^m 88
de l'Université de Toronto

[illegible]

La nonlune du système optique se compose principalement de la cuvette A fig. 2, de la virole centrale B avec bossage houlonne sur l'axe de déclinason, et d'un ossature orthogonale C servant à fixer l'un ou l'autre des miroirs véciculaires, celui de Newton en *m*, celui de Cassegrain en *n*. La cuvette est une pièce nervurée en acier moulé, pesant 950 kg; le grand miroir, qui a environ 300 mm d'épaisseur et pèse près de 2 tonnes.

de duralumin. Les montures des miroirs véhiculaires s'ajustent dans un cadre en acier, centré par quatre larges rubans d'acier à ressort disposés de champ par rapport aux rayons lumineux. Pour la commodité des observations, quatre positions sur deux diamètres orthogonaux sont prévues pour le miroir de Newton; la monture tubulaire du miroir de Cassegrain contient un petit moteur électrique pour corriger dans le sens axial les déformations d'origine thermique.

Mouvements. — Les télescopes étant destinés aux recherches d'astronomie physique, les mouvements ont pour objet de pointer l'instrument et non de mesurer les positions avec une

librent le poids du télescope; à l'intérieur de la boîte est logé un moteur de 0,75 ch qui attaque la roue c par l'intermédiaire d'un

Continued at A.

Continued at C.

cercle sidéral est actionné en utilisant l'un des volants pour régler l'échelle inférieure sur le temps sidéral, après quoi on pointe le télescope avec le moteur F, l'échelle supérieure du cercle sidéral devant alors indiquer l'ascension droite de l'étoile cherchée; on bloque ensuite la roue g, tout en débrayant le moteur F. Un cercle gradué de l'axe polaire permet une lecture grossière de l'ascension droite, en regard d'un repère fixé au socle; un enclenchement s'oppose à la commande simultanée de l'axe polaire par les deux moteurs à la fois.

Le mécanisme d'horlogerie électrique est représenté en élévation sur la figure 7. Il comprend un moteur de 0,5 ch R, attelé par la roue tangente S et par le couple conique T à l'arbre vertical

qui aboutit au système de commande de l'axe polaire; sur l'arbre on est claveté un régulateur à bielles croisées, sensiblement isochrone. Les deux parties de l'arbre horizontal reliant la roue tangente au couple conique tournent en sens contraire par l'effet d'un différentiel dynamométrique, équilibré par le contre-poids U suspendu à une chaîne dont l'autre extrémité s'enroule à la périphérie du boîtier V. Les déplacements du contre-poids font varier une résistance, intercalée sur le circuit de champ du

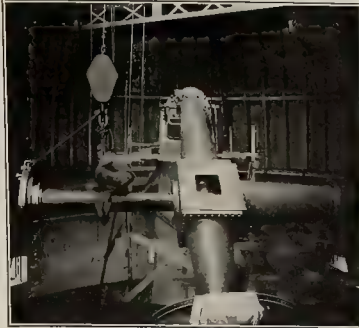


Fig. 6. — Vue prise pendant le montage de l'axe de déclinaison.

moteur; en outre, le régulateur met en action les frotteurs W en cas de dépassement de la vitesse.

La précision de ce mécanisme n'est pas absolument rigoureuse; sur la transmission qui relie l'arbre α et la roue g, par suite, été intercalé un mécanisme correcteur, ou filtre, comportant deux arbres horizontaux x et y. L'arbre x est interrompu par deux systèmes différentiels; sur la partie médiane est monté avec un certain serrage un disque à encoches, dans lesquelles tombe un

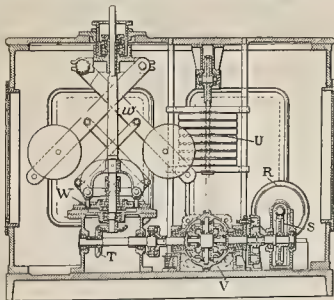


Fig. 7. — Mouvement d'horlogerie électrique communiquant à l'axe polaire la vitesse angulaire du mouvement diurne.

verrou actionné par voie électromagnétique à partir du pendule de l'horloge maîtresse de l'observatoire. Si le synchronisme est en défaut, le disque se trouve basculer un interrupteur à mercure à deux départs. Le courant bloque au moyen d'un électro-aimant l'un ou l'autre différentiel, ce qui a pour effet d'accélérer ou de ralentir le système d'arbre portant le disque, le deuxième système différentiel est actionné, soit par un petit moteur pour la mise au point de la visée, soit au moyen d'un volant pour corriger les erreurs dues à la réfraction.

'GUIDING STAR IS A LOT OF HOOEY'

So Declares Dr. C. A. Chant, University Professor

TORONTO, Jan. 1 (CPI)—And here's one to ponder in the new year. There is no such thing as a "guiding star" to which you can attribute your share of luck, good or bad, in the past or in the future. The idea that the heavenly bodies guide the destinies of man is just a lot of hooey, a survival of pagan beliefs and superstition, based on the mysterious which gullible individuals are too prone to accept as fact.

We have this opinion from no less authority than Dr. C. A. Chant, professor of astro-physics at the University of Toronto, regarded as one of the world's leaders in this branch of the sciences. The professor, who has devoted most of his life to the study of the mysteries of the universe, has no time for superstition, declaring that luck, chance, fate or whatever you choose to call it is not governed by remote control from the stars, but rather part of human experience, which he said was just as baffling as some of the mysteries which cropped up in the field of astro-physics.

Nelson News, B. C., Jan. 2, 1935

Also in
Charlottetown Guardian, P.E.I., Jan. 16, 1935

Le télescope est équipé de trois chercheurs, de 50, 70 et 115 mm d'ouverture, les deux premiers à l'extrémité de l'ossature octogonale, le troisième près du grand miroir; pour la vision latérale (miroir de Newton), l'observateur prend place sur la plate-forme d'un chariot électrique, mobile sur une charpente profilée suivant le quart d'une circonférence, roulant à l'intérieur de la coupole sur deux voies semi-circulaires dénivelées de 5 mètres environ. La coupole hémisphérique a 18°00 de diamètre; son ouverture, de 4°55 de largeur, s'étend de la naissance jusqu'à 2°15 au delà du zénith; elle pèse 80 tonnes, et

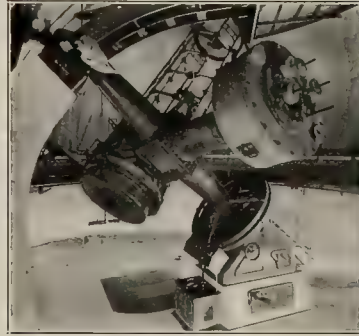


Fig. 8. — Vue du télescope monté dans sa coupole.

repose par 24 galets montés à billes sur le tambour, également métallique, pesant 120 tonnes. La coupole est actionnée par un cabestan électrique.

Le télescope pèse 50 tonnes, dont 35 tonnes pour la partie mobile autour de l'axe polaire; une fosse équipée d'un montage charge est prévue pour démonter la cuvette et le grand miroir, notamment en vue de réaligner ce dernier

J. E.

CANADA'S MILLION-DOLLAR OBSERVATORY OPENS IN MAY



The new Dunlop observatory, which has been erected at Richmond Hill, Ontario, is to be formally opened on May 31, 1935, when many prominent scientists will be present. Among the astronomers who will take part in the opening ceremonies will be Sir Frank Dyson, former Astronomer Royal. Our pictures above show parts of the imposing observatory which will be one of the most modern of its kind in the world. (Inset) Sir Frank Dyson.

Shaunavon Standard, Sask., Feb. 7, 1935

Also in

Kitchener Record, Ont., Dec. 31, 1934

Empress Express, Alta., Feb. 28, 1935

A. P. HARRISON
President of the University of Toronto
who is a member of the Royal Society
and other new officers of the Royal Society

LARGEST TELESCOPE IN BRITISH EMPIRE TO BE IN TORONTO

74-Inch Mirror Now Receiving Finishing Touches in England

DEDICATION IN MAY

Here's one to ponder in the new year. There is no such thing as a "guiding star" to which you can attribute your share of luck, good or bad, in the past or in the future. The idea that the heavenly bodies guide the destinies of man is just a lot of hooey, a survival of pagan beliefs and superstition, based on the mysterious which gullible individuals are too prone to accept as fact.

We have this opinion from no less authority than Dr. C. A. Chant, professor of astro-physics at the University of Toronto, regarded as one of the world's leaders in this branch of the sciences.

No Time for Superstition.

The professor, who has devoted most of his life to the study of the mysteries of the universe, has no time for superstition, declaring that luck, chance, fate, or whatever you choose to call it is not governed by remote control from the stars, but rather part of human experience, which he said was just as baffling as some of the mysteries which cropped up in the field of astro-physics.

Although he has encountered many setbacks in his work, Dr. Chant declared he had never reached the point of utter discouragement or doubted the wisdom of his specialized studies.

"I have no hesitation in saying that at least 90 per cent of astronomers have reached the conclusion that the universe is not the result of any blind law, but is regulated by some sort of intelligence, force or power, call it what you will. There may be some, however, who believe the universe is the result of chance. It all depends on the individual," Dr. Chant said.

"The planets (inhabited) That is a question for the future to determine," the professor declared. There was no reason, however, to believe the earth a tiny speck in the universe, was the only habitable sphere. But the limit of life is only on the other planets might be such as to defy the human imagination.

Huge Telescope.

He doubted if the 74-inch telescope mirror recently cast at Corning, N.Y., would turn out the shiffling. This huge "eye" however will be taken as large as any now in existence, and will enable United States astronomers to probe four times further into the heavens than ever before. It is expected to reveal objects 1,000,000 light years distant as they looked many millions of years before even the first form of life appeared on earth.

But Dr. Chant is primarily interested in his own miniature telescope, which will be officially dedicated on his 50th birthday, May 31, next. The small telescope has already been installed in the Dunlop Memorial Observatory on the Richmond Hill site, about 15 miles north of Toronto, and only awaits the 74-inch mirror which was cast by experts at Corning Glass Works. It is expected to reach Toronto within the next three or four months. It will give us the latest telescope in the British Empire, which will enable us to enter upon a new era in the history of research," Professor Chant declared.

Also in

Ottawa Citizen, Jan. 2, 1935

St. Thomas Times-Journal, Jan. 5, 1935

Cornwall Standard, Jan. 25, 1935

PROFESSOR NO SUPPORTER OF GUIDING STAR

University of Toronto Professor
Declares that Lifetime of Study
Has Proved that Fate Is Not
Guided by Remote Control
from Stars

Toronto, Jan. 2.—And here's one to put in the new year: There is no guiding star, no "guiding star" to which you can attribute your share of the good or bad in the past or in the future. The idea that the heavenly bodies guide the destinies of men is a lot of hooey, as Dr. C. A. Chant, professor of astronomy at the University of Toronto, declared today.

We have this opinion from no authority than Dr. C. A. Chant, professor of astronomy at the University of Toronto, recorded as one of the world's leaders in this branch of the sciences. The professor, who has devoted most of his life to the study of the mysteries of the stars, has no time for superstition or for that luck, chance, fate, or whatever you choose to call it, that is supposed to be made manifest from the stars, but in their part, it is a mere coincidence, which he said was just as likely as some of the events which occurred in the lives of astronomers.

Although he had encountered many setbacks in his work, Dr. Chant declared he had never reached the point of utter discouragement or doubted the wisdom of his special field and that.

"I have no inclination in favour of astrology, nor in favour of any other system which has reached the conclusion that the universe is not the result of a blind law, but a regularity and great intelligence, which we call it what it is. There may be some, however, who believe the universe is a mere chaos. It all depends on the individual," Dr. Chant said.

Sherbrooke Record, Que., Jan. 2, 1935
Also in
St. John Telegraph-Journal, N.B.
Jan. 3, 1935

MONTREAL DAILY STAR,
SATURDAY, JANUARY 19, 1935

OFFICERS NAMED BY ASTRONOMERS

New Toronto Observatory
Described by Toronto
Scientist

Members of the Montreal Centre of the Royal Astronomical Society last night elected new officers for 1935. George R. Lighthall was elected president of the society. Other officers were as follows: Honorary president, Mrs. C. P. Choquette; first vice-president, Dr. Julian C. Smith; second vice-president, O. A. Ferner; secretary, J. W. Spright; treasurer, Dr. A. S. Eve; Li-Co, W. E. Lyman; G. Harper; H. Hall; Dr. L. V. King; Russell Patterson; Henry F. Hall; F. D. Kinder; Dr. C. C. Birchard.

NEW OBSERVATORY
The meeting of the society was held in the McGill Physics Building. G. Harper Hall opened the proceedings and following the election of officers, handed the administration over to the new president. The feature of the new president, Dr. C. A. Chant, professor of astronomy at the University of Toronto, who described in detail the Dunlap Observatory, Toronto, which is now awaiting its "telescope eye."

Two telescopes, when completed, will be the second largest in the world. The 70-inch telescope is the reflecting, or mirror type. Dr. Chant explained that the mechanical difficulties made it impossible to construct a refracting type telescope of this size. The telescope is almost a foot thick. It is being ground down and polished in England.

The speaker was thanked by Dean A. S. Eve, director of the McGill Physics Building and member of the society, who paid glowing tribute to the new worldmanipulation which will be a telescope possible as well as to the vision and perseverance of Dr. Chant.

ASTRONOMICAL SOCIETY

Dr. C. A. Chant, of U. of T.
Will Speak at McGill

Dr. C. A. Chant, professor of astronomy at the University of Toronto, will speak on the David Dunlap Observatory at tonight's meeting of the Royal Astronomical Society of Canada, Montreal centre, in the Macdonald physics laboratory, McGill University, at 8:30 o'clock. Dr. Chant is director of the observatory which will have the largest telescope in Canada, second only at the present time to the great Mount William telescope.

Montreal Gazette, Jan. 18, 1935

BUILDING OF LARGE OBSERVATORY TOLD

Toronto University to Have
Largest Telescope in
British Empire

A photographic record of the construction of the David Dunlap Observatory at Toronto, which on completion will house the largest telescope in the British Empire, was shown to members of the Montreal Centre of the Royal Astronomical Society of Canada at a meeting held in the Macdonald Physics Laboratory, McGill University, last evening. Dr. C. A. Chant, professor of astronomy at the University of Toronto and director of the observatory, described the lantern slides which showed in detail the progress of the work which will be brought to culmination in about two months time when everything will be in readiness for the opening of the new observatory.

Prior to Dr. Chant's lecture, George R. Lighthall was elected unanimously as president of the Montreal Centre of the society for the year 1935. Other officers were also elected by acclamation as follows: Honorary president, Mrs. C. P. Choquette; first vice-president, Dr. Julian C. Smith; second vice-president, O. A. Ferner; secretary, J. W. Spright; treasurer, Dr. A. S. Eve; Li-Co, W. E. Lyman; G. Harper; H. Hall; Dr. L. V. King; Russell Patterson; Henry F. Hall; F. D. Kinder; and Dr. C. C. Birchard.

The David Dunlap Observatory, an adjunct of the University of Toronto, has been erected with funds provided by Mrs. Dunlap in memory of her husband, who died in 1914. Dr. Chant said it is situated near Richmond Hill, about 11 miles from the Toronto waterfront. The 70-inch mirror, which was poured at Corning, N. Y., is now being finished in England and should be ready for shipment to Canada early in March.

The observatory's administrative building, and the structure which will contain the large telescope, were completed some time ago. The great Mount Wilson telescope in California is the only one greater in diameter than that of the new University of Toronto observatory. Dr. Chant's lantern slides proved to be a remarkably complete record of the building of the observatory. Every detail of the construction of the building, from the time the foundations were started to the placing of the last bits of stone and steel, were shown as well as the pouring of the mirror glass for the telescope lens at Corning, New York, and the grinding of the mirror in England.

Montreal Gazette, Que., Jan. 19, 1935

Montreal Gazette, Que., Jan. 19, 1935

Montreal Gazette, Que., Jan. 19, 1935

Montreal Gazette, Que., Jan. 19, 1935

Montreal Gazette, Que., Jan. 19, 1935

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Montreal Gazette, Que., Jan. 19, 1935

Montreal Gazette, Que., Jan. 19, 1935

Montreal Gazette, Que., Jan. 19, 1935

SPEED WORK ON GREAT TELESCOPE

Will Have Largest Lens In
British Empire

AT RICHMOND HILL, ONT.

Observatory Opening Is Scheduled
For May 31

TORONTO, Jan. 22.—(CP)—Since discovery of the telescope little more than three centuries ago, the heavens have slowly yielded some of the profound secrets which through the ages baffled probing astronomers and in this march of investigation Canadian scientists have played an important role.

Among this group is Dr. C. A. Chant, professor of astrophysics at the University of Toronto, who has devoted most of his life to the study of the mysteries of the universe. He declines to discuss himself or his own contributions to science, in his opinion there is a far more absorbing subject—the giant telescope which is being erected under his supervision in the Dunlap Memorial Observatory at Richmond Hill, 15 miles north of Toronto.

The observatory is not only a monument erected through the generosity of Mrs. Jessie D. Dunlap to the memory of her husband, an amateur astronomer, but fulfillment of a cherished dream for Dr. Chant.

It was back in 1907 that he laid the foundation of the present important department of astrophysics at the University and since that day he has looked forward to the time when he would be able to lead his class to wider fields of research with the aid of a powerful telescope. But at the time he thought in terms of a more modest enterprise than the Dunlap 70-inch reflector which will be the largest telescope in the British Empire.

According to tentative plans the observatory will be officially dedicated May 31, Dr. Chant's 70th birthday. Already the huge dome housing the telescope has been erected and the grey sandstone administrative building completed. All that is awaited is the all-important mirror, or "eye" of the telescope, now undergoing its final grinding and polishing process at the Parsons Optical Works, Newcastle-on-Tyne, England. This firm is a subsidiary of the Sir Howard Grubb Company which undertook the contract to build the observatory.

Manufacture of the mirror was the greatest accomplishment of the whole undertaking in Dr. Chant's opinion. The experiment is casting the disc has been a long and tedious process and for some time the experts assigned to this technical work were compelled temporarily to acknowledge defeat. The commission was ultimately transferred to a firm in Corning, N. Y., where success attended the pouring of a mammoth 200-inch mirror for a California observatory.

The Dunlap mirror was poured into its mould on June 23, 1931 and on Oct. 1 the same year was declared a complete success. Then it was shipped to England to undergo further technical treatment. Even if the mirror is delivered in Toronto on schedule, nearly two years will have elapsed since work on the telescope "eye" was started.

"We shall certainly be happy when the mirror is installed, but we have never been impatient," Dr. Chant said. "It would suggest those afflicted with impatience might try astronomy as a cure." In that field results were far from spontaneous, he said, with a premium rather paid to concentration, determination and patience.

London Free Press, Ont.

Jan. 23, 1935

Also in

Leamington Post-News,

Jan. 24, 1935

Kirkland Lake Northern-News

Jan. 25

North Bay Nugget, Jan. 25

Quebec Chronicle-Telegraph

Jan. 25

Niagara Falls Review, Jan. 25

Port Arthur News-Chronicle

Jan. 25

Brandon Sun, Jan. 29

Moncton Times, N.B., Jan. 30

Prince Albert Herald, Feb. 1

Truro Victoria-Inverness

Bulletin, Feb. 2

Calgary Herald, Feb. 3

Kingston Whig-Standard

Feb. 4

The Pas Northern Mail

Man., Feb. 8

PLANETARY LIFE MAY BE FOUND BY NEW GLASSES

Astronomer Believes Super-Powered
Telescopes Will Reveal
Much New Information

Toronto, Jan. 26.—(By Edwin S. Johnson, Canadian Press Staff Writer).—The question of whether the planets are inhabited will ultimately be determined through the advance of astronomy, in the opinion of Dr. C. A. Chant, professor of astrophysics at the University of Toronto.

Research toward that objective has been carried on for many years and now with the advent of super-powered telescopes Dr. Chant believes much information hitherto unknown about the nature of the planets and stars will be disclosed to curious mankind. But many years might yet elapse before a definite answer is forthcoming.

With the official dedication of the Dunlap Memorial Observatory at Richmond Hill, on the outskirts of Toronto, Dr. Chant will take his place among the world's leading astronomical research workers. The dedication, set tentatively for next May 31, will bring into operation the largest telescope in the British Empire, a reflector with a 70-inch mirror.

With his staff of expert assistants Dr. Chant plans to devote his first programme of research to spectrum photography of the stars. This, he said, is one of the most fertile fields of study, but the heavens offered such a vast opportunity for research that it would be almost impossible to set a prescribed objective at any time.

Strange as it might seem, the Dunlap telescope could only function satisfactorily within a restricted field, Dr. Chant said. Its power was such that it must concentrate on the more distant celestial objects. He hesitated to speculate on the possibilities the new telescope opened up. Apart from the spectrum, much remained to be known of the nature of spiral nebulae, radiometry, direct photography of faint celestial bodies and the study of super-novae or "star suicides." Fifteen of these have been recorded or astronomical photographic plates during the last 50 years. Super-novae are stars which explode into flame 100,000,000 times as bright as the sun.

"No, we will not be out scouting for new stars or planets. What is left for the real watchdogs of the sky, the amateur astronomers," Dr. Chant said. "Of course we may encounter the occasional previously uncharted star, but that will be primarily through accident."

Asked what he thought of the theory recently advanced by a Spanish scientist that the sun was a cold planet and possibly inhabited, Dr. Chant remarked: "I read about that, but quickly threw it aside as so much rot. Why, do we not get our heat from the sun and is that not proof sufficient that it is not a cold planet? As for the sun being inhabited, that is an old theory. Some have advanced the belief that sun spots are huge craters leading to the heart of the sun where life might be sustained. We can't deny that possibility, of course, but there is life there, it is such that I might beggar the imagination."

Brandon Sun, Man., Jan. 36, 1935

Also in

Windsor Border Cities Star,

Jan. 26, 1935

Charlottetown Guardian, P.E.I.

Jan. 26

Moncton Times, N.B. Jan. 28

North Bay Nugget, Feb. 1



"Toronto, he admitted, had this advantage over Victoria — while both could see all the stars of the

Continued on next page

OFFICE OF THE PRESIDENT
H. J. CODY, M.A., D.D., LL.D.



April 13, 1935

Mrs. D. A. Dunlap,
93 Highlands Avenue,
Toronto.

Dear Mrs. Dunlap:

The Senate of the University unanimously desires you to accept the honorary degree of Doctor of Laws at the special convocation to be held on the evening of Friday, May 31, in Convocation Hall.

You will be back in time to learn all the detailed arrangements about this ceremony.

With kindest regards,

Yours faithfully,

H. J. Cody
President.

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THE ALL-BRITISH ROUTE

Printed & Published, May 1935.

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ICO-

TORONTO 10, MAY 3, 1935.

MRS DUNLAP
2 PRYTHOON STREET
PORTLAND PLACE
LONDON

MIRROR AND SPECTROGRAPH SAILING ON GREAT DUNE

CHANT

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STANDARD TIME

W. D. NEIL, GENERAL MANAGER OF COMMUNICATIONS, MONTREAL

RNA6 17 COML=LONDON 5*

NLT PROFESSOR CHANT=

TORONTO UNIVERSITY TORONTO=

THANKS FOR MESSAGE GLAD NEWS CONGRATULATIONS SAILING EUROPA ELEVENTH

=JESSIE DUNLAP=



IT'S HOME-COMING YEAR IN TORONTO

Richmond Hill Telescope Largest in British Empire

Newly-Completed Mirror
Study Heavens Through
Huge Lens

EXPOSE PLANETS

Toronto, April 26.—The largest telescope in the British Empire, at the Dunlop Observatory, Richmond Hill, Ont., is nearing completion. Some night in the near future, Professor C. A. Chant and his associates will realize their fondest dreams when they begin studying the heavens through the giant 74-inch lens, presented to the University of Toronto by Mrs. D. A. Dunlap and her son, D. Moffat Dunlap.

The monster mirror will be two inches larger than the one at Victoria B.C. This, however, according to Prof. Chant, will not give an appreciable advantage. They will, as a matter of fact, be virtually tame. The position of the new telescope will place it 500 feet above sea level. Besides being the largest in the British Empire it will be the second largest in the world, surpassed only by the 100-inch telescope at Mount Wilson in California.

The mirror was cast June 21, 1933, at Corning N.Y., and was then shipped to England where it has been undergoing a delicate grinding and polishing. The huge meniscus is 74 inches in diameter, 12 inches thick at the edge and 11 inches in the center. To ensure a perfect flawless mirror, the press of which it is made, was poured at a temperature of 1,500 degrees centigrade or 2,700 degrees Fahrenheit. It was then allowed to cool to 500 degrees F. before it was placed in an annealing oven for three months while the temperature was minutely lowered. "It was necessary to do this," said Prof. Chant, head of the department of astronomy at the University of Toronto and who will be in charge of the new observatory, "in order that it would show no strain or flaw. It came out perfect."

Magnifies 40,000 Times

The mirror will have a magnification 40,000 times that of the human eye. It will magnify to such an extent that it gives the size of Toronto, Winnipeg, or Calgary would be clearly visible on the moon, if the moon had such cities. The Great Lakes, if the great orb had any, would be as plainly seen by the observer. A telescope of the like power would not be seen. Dr. Chant, 2 H. Hogg, a young scientist who was an attendant astronomer of the Dominion Astrophysical Observatory at Victoria and who came to Toronto to work on the new telescope, pointed out the outline of a group of buildings, such as the Canadian National Library building at the University of Toronto might be seen.

out he would not commit himself further.

There on the high ground north of Toronto, the astronomers will work, like sailors during the Christmas rush, from sunset to dawn taking spectroscopic pictures of the heavens. Victoria has in the last 10 years taken and filed some 23,000 such photos. Nor will the scientists be the only ones to view the moon. In all probability the observatory will be opened to the public one night a week so that housewives and office workers may see the moon magnified 40,000 times.

Take Accurate Pictures

The observatory which has been under construction for the past four years is now awaiting installation of the mirror. The administration building of cut stone, provides space for offices, a library and for the laboratories necessary for accurate measurement of the photographic plates. The building which houses the telescope has a 61-foot dome weighing 50 tons and which rests on 24 columns, each 27 inches in diameter. This permits rotation of the dome as the planets in the heavens shift their positions.

Watching an object in the sky is not the easy job which may be imagined. Stars and all heavenly bodies are continually shifting, and it is necessary that the telescope be accurately pointed at the moving object. The difficulties encountered may be compared with the keeping of a gun pointed at a one-inch target twenty miles away while the target moves at the rate of five feet per second. Movement of the image by only a thousandth of an inch while it is being photographed will make a serious defect in the resulting picture.

The value of a telescope during an eclipse of the sun is seen in the photographs it will permit which will show the heaping prominence of the corona. Telescopes have shown the prominences extending a distance of 50,000 miles. With the arrival of the mirror, the scientists enter the final stage of the work. It is expected that the observatory will be formally opened on May 31, the 70th birthday of Prof. Chant.

North Bay Nugget
Ont. April 26, 1935

GIANT TELESCOPE SOON TO OPERATE

EMPIRE'S LARGEST LENS TO BE COMPLETE BY SPRING.

By EDWIN S. JOHNSON,
Canadian Press Staff Writer

Toronto, April 25 (P.P.S.)—Since discovery of Canada's first mine more than three centuries ago, the province has only yielded one of the precious metals which through its veins brilliant picture of a future is in the making. It is a future which Canadian scientists have played an important role.

Among this group is Dr. A. C. Chant, professor of astronomy at the University of Toronto, who has devoted much of his life to the study of the stars and planets. He does not see his time as an astronomer, but as a scientist. In his opinion, the giant telescope which is being erected under his guidance at the Dunlop Memorial Observatory, at Richmond Hill, is only the beginning of a new era in Canadian astronomy.

The observatory is not only a monument erected through the generosity of Mrs. Jessie D. Dunlap to the memory of her husband, an amateur astronomer, but also a gift of a life-long dream to Dr. Chant. It was a dream that he had since his childhood, and that he had been working towards it ever since. He had been working towards it ever since he was a child, and that he had been working towards it ever since he was a child. He had been working towards it ever since he was a child, and that he had been working towards it ever since he was a child.

Moose Jaw Eve. Times
Sask. April 2, 1935

Transporting the Great Mirror



Toronto Honors Observatory Participants

Special to The Christian Science Monitor
TORONTO—The opening of the David Dunlap Observatory at Richmond Hill recently was concluded at the University of Toronto when honorary degrees were conferred upon six prominent participants in the earlier ceremony.

Degrees of doctor of laws (honoris causa) were conferred on Sir Frank Watson Dyson, former Astronomer Royal of England; Mrs. Jessie Dunlap, donor of the observatory; and Dr. C. A. Chant, director of the observatory. Doctor of science degrees (honoris causa) were conferred on Harlow Shapley, director of Harvard College Observatory; Vesto Melvin Slipher, director of Lowell Observatory; and W. E. Harper, acting director of the Dominion Astrophysical Observatory, and a former pupil of Dr. Chant.

It was announced that Dr. Chant would become emeritus professor of astrophysics, emeritus director of the observatory, and honorary librarian and director of publications at the observatory. Dr. Chant graduated from the University of Toronto in 1890, returning the following year as a fellow in the department of physics, and since that time he has been connected continuously with the university.

The great telescope, the second largest in the world, was visualized as a means of studying double stars, analyzing the spectrum of the stars, photographing clusters and solving the problem of fragments by Dr. Shapley of Harvard Observatory. By the new observatory Canada joins the United States in making Western Hemisphere astronomy a thing for the future to think about, he said.

The need for greater observation and the study of distant planets was stressed by Dr. Slipher. Everyday uses of astronomy were outlined by Dr. Harper, who contended that a better understanding of it would eliminate superstitious fears.



Unloading the Mirror from the railway car to the truck.

*Arriving at the Observatory.
 Unloading from the truck to the Dome*

May 2, 1935



UNIVERSITY OF TORONTO
TORONTO 5, CANADA
REGISTRAR'S OFFICE

May 14th, 1935.

Mrs. D. A. Dunlap,
93 Highlands Avenue,
Toronto.

Dear Mrs. Dunlap:

The President of the University has directed me to inform you that the University will, if you so desire, provide you with the academic costume to be worn at the Special Convocation on Friday, May 31st, including the cap, gown, and hood, and will ask you to accept the hood as a souvenir of the occasion. In this connection, perhaps you would be kind enough to call at the Registrar's Office, Simcoe Hall, at your early convenience in order that you may be fitted with cap and gown.

I enclose herewith an outline of the ceremony of conferring degrees as observed by the University of Toronto.

I shall be glad to receive from you a list of your friends to whom you would wish to send invitations for the official opening of the David Dunlap Observatory and the Special Convocation.

Yours very truly,

A. J. Fennell

Registrar.

Abf:dw

UNIVERSITY OF TORONTO
THE CEREMONY OF CONFERRING HONORARY DEGREES

Each candidate for an honorary degree will wear the cap and gown appropriate to the degree which he is to receive, and will walk in procession to Convocation Hall where he will take the seat assigned to him upon the platform.

The presenter for each candidate, usually the President of the University, will rise and, after saluting the Chancellor, will deliver the presentation address. He will then formally present the candidate who will rise and advance to the Chancellor. The Chancellor will take the candidate's hand and confer the degree upon him, the deed placing the hood upon his shoulders.

The Chancellor will then request the candidate to sign the Golden Book. After having done so, the candidate will return to his seat.

HONORED BY U. OF T.



MRS. D. A. DUNLAP
On whom the University of Toronto will confer the degree of Doctor of Laws (honoris causa) at a special convocation on May 31 in connection with the opening of the David Dunlap Observatory.

WILL RECEIVE LL.D.



MRS. D. A. DUNLAP
on whom the University of Toronto will confer the degree of Doctor of Laws (honoris causa) at a special convocation in connection with the opening of the David Dunlap Observatory on May 31. Mrs. Dunlap is the wife of the late Dr. D. A. Dunlap, a leading philanthropist.

The Globe
May 11, 1935

Observatory Donor
To Receive Degree

University to Confer Honors on Day of Official Opening.

To mark the official opening of the David Dunlap Observatory at Richmond Hill on May 31 the University of Toronto will confer a degree, LL.D., on Mrs. D. A. Dunlap, widow of Sir Frank Watson Dunlap, K.B.E., F.R.S., LL.D., D.Sc., former Astronomer Royal of Great Britain. Mrs. Dunlap is the widow of the late Sir Frank Watson Dunlap, K.B.E., F.R.S., LL.D., D.Sc., former Astronomer Royal of Great Britain. Mrs. Dunlap is the widow of the late Sir Frank Watson Dunlap, K.B.E., F.R.S., LL.D., D.Sc., former Astronomer Royal of Great Britain. Mrs. Dunlap is the widow of the late Sir Frank Watson Dunlap, K.B.E., F.R.S., LL.D., D.Sc., former Astronomer Royal of Great Britain.

DUNLAP
OBSERVATORY
OPENING

Set for May 31

The David Dunlap Observatory will be opened to the public on May 31, 1935, at 10 o'clock. The observatory is situated on Richmond Hill, Ontario, and is the largest of its kind in Canada. It was built by the University of Toronto and is named in honor of the late Sir Frank Watson Dunlap, K.B.E., F.R.S., LL.D., D.Sc., former Astronomer Royal of Great Britain. Mrs. Dunlap is the widow of the late Sir Frank Watson Dunlap, K.B.E., F.R.S., LL.D., D.Sc., former Astronomer Royal of Great Britain.

The Mail and Empire
May 11, 1935

From Sir Frank Dyson

181

1935 May 13

27, WESTCOMBE PARK ROAD,

BLACKHEATH, S.E.3.

Dear Mr. Dunlap,

Prof Chant wrote to me that you have kindly invited me to stay with you when I come for the opening ceremony of the David Dunlap Observatory. He also gave me your address in London, but when my wife and I called to day we found that you had already returned to Canada. We have the pleasure of meeting you.

TELEPHONE
GREENWICH 5322

27, WESTCOMBE PARK ROAD,

BLACKHEATH, S.E.3.

go to a young naval officer and is expecting a baby early in June. She wishes to be remembered very kindly to you.

I am coming by the steamer on May 19 which should reach Montreal on May 24 and I shall come immediately to Toronto. I shall return by the Aquitania from New York which sails on June 7, so as to be in time for the International

"I think there is greater cooperation between astronomers among members of any other branch of science," he declared.



*At the Official Opening
May 31, 1935*

Hon. D. Marshall, Lieut.-Gov. Bruce, President Cody, Mrs. Dunlap, Dr. D. B. Macdonald.



UNIVERSITY OF TORONTO

ORDER OF PROCEEDINGS

AT THE PRESENTATION TO THE UNIVERSITY
AND THE OFFICIAL OPENING
OF THE

DAVID DUNLAP OBSERVATORY

BY

JESSIE DONALDA DUNLAP

MAY 31st, 1935

New Observatory At Toronto Has Huge Telescope

Mrs. Dunlap Donor of Astrophysical Plant to Honor Husband

Spoke in the Christian Science Monitor.
TORONTO, May 31—The David Dunlap Observatory, housing the second largest telescope in the world, was formally opened today when Mrs. Dunlap, the donor, opened its door with a key of gold. She then presented the building and its equipment to Dr. Bruce Macdonald, Chairman of the Board of Governors of the University of Toronto, the president of which, Dr. H. J. Cody, presided at today's ceremony.

Dr. Herbert A. Bruce, Lieutenant Governor of Ontario, conveyed the greetings of the province. Greetings were also brought by Cyril Young of Newcastle Eng., representing the Sun which had the task of building the big telescope, Sir Frank Dyson, former astronomer royal of England, Prof. Harlow Shapley of the Harvard Observatory, Prof. H. D. Guthrie, Observatory of the University of Michigan, and Dr. W. E. Harper, Acting Director of the Dominion Astrophysical Observatory, Victoria, B. C.

For many years it had been the purpose of Dr. C. A. Chant, professor of astrophysics at the university, to have an observatory and undertake definite research into the knowledge of astronomy as a department of the institution. Today, on his seventieth birthday, he realized that ambition when he received charge as its director of one of the greatest astrophysical institutions of the world, having at the present time the second largest telescope in the world.

Interviewed by The Christian Science Monitor, Dr. Chant described some of the mechanism he will use in the new observatory and explained the work upon which he will be occupied in the future.

The new 16-inch telescope of the



Mr. A. S. Mathers presenting the Golden Key to Mrs. Dunlap



The Chancellor, the President, the Governors and the Senate

of the

University of Toronto
request the honour of your presence
at the

Official Opening of the David Dunlap Observatory
at Richmond Hill

on Friday, May 31st, 1935, at 3 p.m.

R.S.M.P.
The Registrar,
University of Toronto

Invitation to the Opening

The Opening of the Observatory, May 31, 1935



Listening to the Addresses



Going to the Great Dome



President Cody and Rt. Hon. W. B. M. King



Dr. Cody Speaking



Dr. Bruce Speaking

Mrs. D. A. Dunlap



The President
of the University of Toronto
requests the pleasure of your company at

Dinner

on Friday May 31. at 7.30 p.m. sharp
Faculty Union Hall House
in honor of recipients of honorary
degrees.
R.S.N.P.
President's Office

UNIVERSITY OF TORONTO

SPECIAL CONVOCAION

FOR THE CONFERRING OF

HONORARY DEGREES

ON THE OCCASION OF THE
OFFICIAL OPENING OF THE
DAVID DUNLAP OBSERVATORY

FRIDAY, MAY 31st, 1935

H. J. Godey
F. W. Dyson
Robt. A. Kalinos.
W. Macdonald King
Harlow Shapley
Luman S. Arthur
H. Ferguson

A. H. Needler
Burgon Brickersteth
W. E. Harper
O. A. Gage
C. A. Gant
Amielee Acknowled
Vinduphan
J. P. Gibson
J. H. H. H.
H. Lennell
C. Young
Kynolok Young
C. Dr. Dr. Dr. Dr.
Mehy D. Curtha
J. P. Gibson
Jessie D. Dunlap

Those at the Dinner



THE UNIVERSITY OF TORONTO
DAVID DUNLAP OBSERVATORY
FRIDAY, MAY 31st, 1935

Dunlap Observatory, Largest in Empire Is Formally Opened

Outstanding Astronomers
Attend Ceremony as
Giant Telescope Is Pre-
sented as Memorial to
U. of T. — Unflagging
Effort Climaxed

TORONTO, through the University of Toronto, and by the generosity of Mrs. Jessie Donalds Dunlap, was virtually given the freedom of the world, and elevated to rank among the world's foremost cities of science with the dedication and formal opening of the David Dunlap Observatory, Richmond Hill, yesterday afternoon.

Before a crowd of more than 1,000, the dedication of the outstandingly modern and complete astronomical observatory, which cost \$500,000, was formally opened and dedicated to the University of Toronto. Received by Dr. D. B. Macdonald, chairman of the board of governors, the observatory was formally opened by the University of Toronto, and the observatory was formally opened by the University of Toronto, and the observatory was formally opened by the University of Toronto.

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Crowd Looks on As Big Telescope Is Bared to View

1,000 Guests Watch in
Awe as Observatory
Opened.

AT RICHMOND HILL
Presented to University
as Memorial to David
Dunlap.

A great hemispherical dome of 80 tons surfaced with copper, yesterday rotated about the top of a circular stone structure. Two parallel shafts slid open and the mammoth 74-inch telescope that is to unfold the secrets of the heavens here into view. A thousand invited guests looked on with awe as the telescope moved with ease to the sky.

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HUGE TELESCOPE DEDICATED MRS. DUNLAP GIVEN DECREE

Scientists See Enlargement of
Soul in Increase of Knowl-
edge of Infinite

PRaise For Chant

With the sun smiling its greeting, the David Dunlap Memorial Observatory was formally dedicated to the University of Toronto yesterday afternoon by Mrs. Jessie Donalds Dunlap when she turned a gold key in the door of the administration building before 1,000 guests.

Mrs. Dunlap, dedicating the observatory, said: "I recall and express the deep interest of my husband in astronomy. I hope it will help to advance the science as she turned the key to open the building."

"It will be a memorial to future generations," declared Dr. Bruce Macdonald, chairman of the board of governors, accepting the gift on behalf of the University of Toronto. "It points to the sky like David Dunlap, who was always looking up for the truth."

Dr. Macdonald's response carried a tribute to David Dunlap, whom he described as a man who was always looking up for the truth. He said that Dunlap was a man who was always looking up for the truth.

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within make real the Supreme Being. There are no limitations.

Get Degrees

At a special convocation in Convocation hall last night, the honorary degree of Doctor of Laws was conferred upon Sir Frank Dyson, former astronomer royal of England, Mrs. Jessie Donalds Dunlap, and J. D. C. Chant, degrees of Doctor of Science were conferred upon Dr. Harlow Shapley, director of the Harvard college observatory, Dr. W. M. S. Slipher, director of the Lowell observatory, and W. H. Harper, director of the Dominion Astrophysical observatory.

The Rt. Hon. W. L. Mackenzie King, recovered from his recent illness, announced that it had been his intention to attend the convocation. "Without being seen in Toronto or missed at Ottawa," to pay a special tribute to all those who contributed to the realization of the construction and completion of the observatory.

"This is a great day for the university and a great day for this country but also a day of astronomical achievement," declared Mr. King. "I did not come here for relaxation or recreation, but for inspiration. This noble work is a great inspiration."

He warmly extolled Mrs. Dunlap for her contribution to the Richmond Hill astronomical centre, he said, was a combined monument, resembling the small observatory that he had seen in the past. He said that Dunlap was a man who was always looking up for the truth.

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Degrees Follow Opening Of Dunlap Observatory

Noted Scientists and
Scholars Join in Event
at U. of T.—Six Are
Honored

A post-graduate student, who had been studying at the University of Toronto, was formally opened and dedicated to the University of Toronto. Received by Dr. D. B. Macdonald, chairman of the board of governors, the observatory was formally opened by the University of Toronto, and the observatory was formally opened by the University of Toronto.

The observatory, which cost \$500,000, was formally opened and dedicated to the University of Toronto. Received by Dr. D. B. Macdonald, chairman of the board of governors, the observatory was formally opened by the University of Toronto, and the observatory was formally opened by the University of Toronto.

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English Scientist Doubly Indebted

Double thanks were given to the University of Toronto last night by Sir Frank Dyson, former Astronomer-Royal of England.

Sir Frank was among other distinguished scientists who received an honorary degree from the University with the formal opening of the David Dunlap Observatory yesterday.

After thanking the University for conferring the degree of Doctor of Laws upon him, Sir Frank said: "I am also indebted to Toronto for a great medical discovery. My sister must have lived for a long time longer because of the discovery of Banting, Best and Macleod."

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The Telegram

Six Honorary Degrees Given by University

Professor Chant, Mrs.
Dunlap and Visiting
Astronomers Honored.

Six prominent participants in yesterday's opening of the new David Dunlap Observatory, last night were honored by the University of Toronto.

Degrees of Doctor of Laws (honorary) were conferred upon Sir Frank Dyson, former Astronomer Royal, Jessie Donalds Dunlap, former Astronomer Royal, and Clarence A. Chant, director of the observatory, and prize money degrees (honorary) were conferred upon Harlow Shapley, director of Harvard College Observatory, and W. H. Harper, director of the Dominion Astrophysical Observatory.

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Saturday Night, June 8 th.

The New York Times, Sat. June 1, 1935

NEW OBSERVATORY
OPENED AT TORONTOGift of Widow in Memory of
Husband, It Contains World's
Second Largest Telescope.

APERTURE OF 74 INCHES

Prominent Canadian, American
and British Astronomers At-
tend Ceremony of Dedication.

By JAMES STOKLEY.

SPECIAL TO THE NEW YORK TIMES.
TORONTO, May 31.—Equipped with the world's second largest telescope, a reflector seventy-four inches in aperture, the David Dunlap Observatory of the University of Toronto was opened at Richmond Hill near here this afternoon.

Several hundred guests, including Canadian, American and British astronomers and government officials, saw Mrs. Dunlap, who gave the observatory as a memorial to her husband, receive the keys from the architects, and, after opening the door to the administration building, turn the observatory over to Dr. D. B. Macdonald, chairman of the university's board of governors.

In introducing Dr. Clarence A. Chant, director of the observatory, the Rev. H. J. Cody, president of the university, said today had been chosen for the ceremonies because it was the seventieth birthday of Dr. Chant, who had sought such an observatory for years.

Dr. Chant revealed that in 1921, following a lecture on a comet then visible, he had mentioned the possibilities of such an observatory. Afterward Mr. Dunlap introduced himself and expressed interest in the project. Though he died in 1924, his widow later decided to give the project in his memory.

Ordered From England.

The great telescope was ordered from Grubb, Parsons & Co. at Newcastle-on-Tyne, England, in 1930. After it proved impracticable to have the large glass disk for the mirror made in England, it was ordered from the Corning Glass Works of Corning, N. Y., and was made of material similar to that used in the 200-inch disk now cooling which will ultimately form the eye of the great telescope of the California Institute of Technology.

The disk was sent to England, where it was ground and figured at the Parsons works. In addition to Dr. Chant, addresses were made by Dr. H. Bruce, Lieutenant Governor of Ontario; Cyril Young, general manager of Grubb, Parsons & Co.; Sir Frank Dyson, former Astronomer Royal of Great Britain; Dr. H. D. Curtis, director of the University of Michigan Observatory; Dr. Harold Shapley, director of the Harvard College Observatory; and Dr. W. E. Harper, acting director of the Dominion Astrophysical Observatory.

At a special convocation this evening, honorary degrees of LL.D. were conferred on Mrs. Dunlap, Sir Frank Dyson and Dr. Chant, and honorary Sc. D. degrees on Dr. Shapley, Dr. Harper and Dr. Vesto M. Slipher, director of the Lowell Observatory, Flagstaff, Ariz.

Away From City's Smoke.

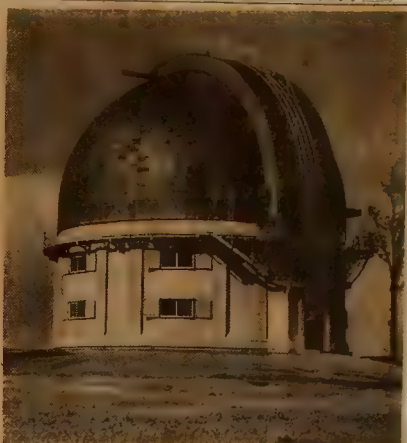
The observatory consists of two buildings erected on a hill two miles north of Toronto, where the prevailing winds, from the northwest, will blow the smoke from the city in the opposite direction.

The Administration Building, surrounded by three domes to contain smaller telescopes, contains the office, library and computing rooms. The 74-inch telescope is housed in a circular steel building, over 100 feet in diameter, which is the largest rotating dome weighing eighty tons. The moving parts of the telescope weigh thirty tons, and while they can be moved electrically, are so perfectly balanced that they can be operated by hand.

Dr. R. K. Young, Professor of Astronomy in the university, estimates that during the normal year about 125 nights will be satisfactory for observations, with 1,200 observing hours. Practically the entire time of the telescope will be used for photographing the heavens, either directly or through the spectroscopic, to determine the motions of the stars and the intensities of the various lines in stellar spectra.

Completion of this telescope gives Canada the second and third largest telescopes in the world. The 72-inch at the Dominion Astrophysical Observatory, Victoria, had previously held second place. Within two years, however, the 82-inch reflector being built in Cleveland for the Macdonald Observatory of the University of Texas will be completed, and possibly by 1940 the 200-inch in California will be searching the skies.

The world's largest telescope at present is the Carnegie Institution's 100-inch at Mount Wilson Observatory, California.



NEW OBSERVATORY OPENED IN CANADA.

The gift of Mrs. David Dunlap, which was turned over yesterday to the University of Toronto as a memorial to her husband. The building houses the second largest telescope in the world.

The Literary Digest
June 15, 1935

Toronto's Telescope

In 1921, Prof. Clarence Augustus Chant, in one of his popular astronomy lectures at Toronto, declared that Eastern Canada needed a large telescope. Out of the crowd, at the conclusion of the address, stepped David Alexander Dunlap, wealthy geologist and mining lawyer, to inquire what such a telescope might cost.

Three years later, Mr. Dunlap died. His widow decided, in 1926, to build an observatory, dedicating it to his memory. A fortnight ago, Professor Chant's four teen-year-old dream of a fine mirror for the stars at Toronto came true, on his seventieth birthday. With Professor Chant, as Director, routine observation of the stars began last week at David Dunlap Observatory of the University of Toronto.

The observatory's seventy-four-inch reflecting telescope is the largest in the British Empire, eclipsing the huge reflector of the Dominion Astrophysical Observatory at Victoria, British Columbia, by two inches. It was made at Newcastle-on-Tyne, England, by the British firm of Grubb, Parsons & Co., but is not all British; the glass disk for the mirror was cast at Corning, New York, of the same low-expansion type of material as that used in the disk for the forthcoming 200-inch telescope of the California Institute of Technology.

The David Dunlap Observatory also has a nineteen-inch reflecting telescope, made by Dr. R. K. Young, Associate Director and Professor of Astronomy at the University. The large reflector is exceeded in size by only one telescope in the world, the 100-inch at Mt. Wilson Observatory.

It will remain second in size for several months, pending completion of the eighty-two-inch telescope of the McDonald Observatory of the University of Texas, to be finished next year.

Professor Chant will retire at the end of June from his professorship at Toronto, to devote his time to the observatory. He joined the University staff in 1891, became Professor of Astrophysics in 1918. He has been Editor of *The Journal of the Royal Astronomical Society of Canada* since 1907.

At seventy, he bears a strong resemblance to John D. Rockefeller, Sr., as he was in his middle age; his nose is long, lips narrow, eyes deeply set. His hobbies are boating and fishing, but in a life devoted almost entirely to astronomy he never has taken much time to enjoy them.

The work of the new observatory has been planned for several years. It will consist of studies of the velocities, temperatures, and densities of stars.



Dr. C. A. Chant



OPENING THE DUNLAP OBSERVATORY. This group includes Cyril Young, general manager of the British firm which built the telescope, Prof. C. A. Chant, director of the Observatory (who was also celebrating his 70th birthday), Principal E. W. Wallace of Victoria College, the Rt. Hon. W. L. Mackenzie King, and (behind at right), A. S. Mathers, of the architect firm of Mathers & Haldenby, which designed the structure.

—Photograph by "Star"

TORONTO

"THE fault, dear Brutus, is not in our stars, but in ourselves, that we are underlings," so it is not surprising that we are becoming more astronomically-minded to know our stars better and thus better understand ourselves in this age of research.

Mrs. David Dunlap's magnificent memorial gift in memory of her husband—the greatest Observatory in the British Empire—has uplifted our eyes to something beyond the passing show. It also brought the most noted astronomer to Toronto—Sir Frank Dyson, the former Astronomer-Royal at Greenwich, was Mrs. Dunlap's guest. And it is remarkable enough to repeat here the comment by a very wonderful old lady aged ninety-six years of age, when she heard of the occasion. Fancy one's grandmother at the age of 96 quoting to me Addison's Ode, "word perfect." "The spacious firmament on high, with all the blue, ethereal sky, and spangled heavens, a shining frame, their great Original proclaim. Soon as the evening shades prevail, the moon takes up the wondrous tale, and nightly to the listening earth repeats the story of her birth, while all the stars that round her burn, and all the planets in their turn confirm the tidings as they roll, and spread the truth from pole to pole. Forever singing, as they shine, the hand that made us divine." I think that must be one of the finest of the first tributes Mrs. Dunlap's great gift inspired—a woman of 96 quoting a poem of the seventeenth century and it was that sort of a really memorable occasion—(unless when Dr. Bruce Macdonald as chairman of the Board of Governors of the University, arrested this deed of 2 ft. which Mrs. Dunlap had opened with a golden key and presented in a charming manner with a few well-chosen words. That magnificent set of buildings will be certain to attract society as well as scientists to astronomy—they defy a layman's description. His Honor, Dr. Bruce, quoted some very appropriate phrases in contemplation and President Cody's appreciation was but a prelude to Convention ceremonies when Mrs. Dunlap was the one woman among the distinguished astronomers to be honored with the honorary LL.D. and to attend Dr. Cody's Convocation dinner at Hart House.

PROFESSOR CHANT, who on this 70th birthday was realizing the hopes of a lifetime, was a centre of attention, but a prelude to Convention ceremonies when Mrs. Dunlap was the one woman among the distinguished astronomers to be honored with the honorary LL.D. and to attend Dr. Cody's Convocation dinner at Hart House.

Mr. Mackenzie King had come all the way from Ottawa, but Oxford and ex-convict Dr. Moffat Dunlap to be present at the opening of this wonderful memorial to his father. Two favorite relatives of Mrs. Dunlap there were Mrs. Ray Hodge and Mrs. R. L. Blackburn, of Ottawa, and Mr. Moffat Dunlap, who is a brother of Mrs. Moffat Dunlap, brought his lovely wife, who was formerly Miss Anne Oeler. Her track, car and violet flowered, was



MRS. DAVID DUNLAP, the donor of the magnificent David Dunlap Observatory which was presented to the Toronto University in memory of the late David Dunlap.



OBSERVATORIES ARE WINDY PLACES. His Honor, Dr. Bruce, President Cody, and Mrs. Dunlap at the Dunlap Observatory opening.

—Phot by "Star"

tribution building blazed with luncheon of this planet who did the outfit of the tented most luxuriously later Sir Frank Dyson, whose quarters in the Greenwich Observatory were just nearby the site of that great Wolfe-Montreal commemoration that was a unique Canadian occasion a couple of years ago, was involuntarily making star gazers out of tea-drinkers with his charming enthusiasm for his hobby. And star-scattered on the glass, guests were quite literally turned down-cups empty in their haste to cheer observe the Observatory.

Mr. Mackenzie King had come all the way from Ottawa, but Oxford and ex-convict Dr. Moffat Dunlap to be present at the opening of this wonderful memorial to his father. Two favorite relatives of Mrs. Dunlap there were Mrs. Ray Hodge and Mrs. R. L. Blackburn, of Ottawa, and Mr. Moffat Dunlap, who is a brother of Mrs. Moffat Dunlap, brought his lovely wife, who was formerly Miss Anne Oeler. Her track, car and violet flowered, was

Mrs. Herbert A. Bruce had got tonnes of floral tone on her blue ensemble. Mrs. H. J. Cody and Mrs. D. Bruce, Macdonald were lucky ones with wide, shady hats and Mrs. Ribbell, with Mr. Justice Ribbell, were a native Gainsboroughish one. Mr. J. H. McConnell wore a pretty party one too. Mr. Justice Meston wore one of the few grey toppers, pretty Mr. George Hesse had a tinfo hat. Mrs. E. A. Dunlap came with her son, as escorted by her son, Mrs. Vincent Massey, escorted by her husband, Mr. Arthur Kirkpatrick, was explaining that he really was not ill. Mr. Mathers and Major Haldenby, the architects, both wore a paternal "as they" survived their poverty. Professor R. B. Thomson, the botanist, had momentarily turned his thoughts to sky diets instead of moon-does and Dr. W. H. Lowry, the artist, was training his own vision this time on the distant spectrum. So one and all were high-minded and it was an unforgettable occasion.



SIR FRANK DYSON, the noted Astronomer, former Astronomer-Royal (right) snapped with Mrs. Vincent Massey, Mrs. Dunlap and the Hon. Vincent Massey at the opening of the David Dunlap Observatory.

1935 June 7

ON BOARD
CUNARD WHITE STAR
"AQUITANIA"

Dear Mr. Dunlap,

I owe you an immediate letter for your most kind hospitality. But I shall send a fuller one later on. I did not see Mr. David Mulligan but saw the Assistant Manager. They were most kind and let me have a room which they did not charge for. Schlenger - the Yale Professor - had sent a young Columbia Professor named Eckhardt to meet me. I took him

1082

NATURE

JUNE 29, 1935

The David Dunlap Observatory, Toronto

ALTHOUGH most of the astronomers of Canada are graduates of the University of Toronto, hitherto the University has possessed no observatory. This want has been recently supplied by the munificence of Mrs. D. A. Dunlap, who has presented the University with a 74-inch reflecting telescope, as well as a handsome administrative building. The inauguration ceremony took place on May 31, the chair being taken by the president of the University, Canon Cody, in the presence of the Lieutenant-Governor, Dr. Bruce, Sir Robert Falconer (a former president of the University), Mr. Mackenzie King (a former Prime Minister of the Dominion), professors of the University, and astronomers from Great Britain, Canada and the United States, and a thousand interested visitors.

After a dedicatory prayer by the Rev. E. W. Wallace, Chancellor of Victoria College, the chairman read letters of congratulation from the president of the International Astronomical Union, the president of the Royal Astronomical Society of Canada, Sir James Jeans and Sir Arthur Eddington. He then called on the architect, who handed a golden key to Mrs. Dunlap. She opened the door with the words, "In loving memory of my husband, David Alexander Dunlap, I now present this astronomical observatory to the University of Toronto, believing this memorial will express his deep interest in astronomy, and I hope through its equipment great advances will be made in the science," and handed the key to Dr. B. M. Macdonald, chairman of the Board of Governors.

After a warm expression of thanks by Dr. Macdonald, a sincere tribute to Mr. Dunlap was paid by him and by the Lieutenant-Governor. The president then called on Prof. Chant, the director of the Observatory, and noted that the inauguration had been fixed on Prof. Chant's seventeenth birthday. Prof. Chant referred to Mr. Dunlap's great interest in astronomy, and said that in 1926 he ventured to lay the project of an observatory in Mr. Dunlap's memory before her. The foundation stone was laid in 1932 by Mr. Moffatt Dunlap. To Mrs. Dunlap all the credit is due; were it not for her there would be no observatory here.

Mr. Cecil Young, manager of the firm of Sir Howard Grubb, Parsons and Co., then gave an account of

the large telescope and dome which were described and illustrated in an article in NATURE of October 14, 1933. Sir Frank Dyson gave the good wishes of the Royal Astronomical Society, and congratulated Prof. Chant on the great interest in astronomy in Canada, which was in large measure due to him. Short addresses were given by Prof. H. D. Curtis, director of the Observatory of the University of Michigan, Prof. V. Slipher, director of the Lowell Observatory at Flagstaff, Prof. H. Shapley, director of the Harvard College Observatory, and Dr. W. E. Harper, director in charge of the Dominion Observatory at Victoria.

The Observatory is situated in 160 acres of ground on Richmond Hill, some twelve miles north of Toronto, from which a beautiful view extends in all directions. It is sufficiently distant from Toronto to avoid smoke and glare, and yet near enough to the University. The number of good observing nights is estimated at about 120 in the year. The administration building is a handsome structure of stone, surmounted by three domes for smaller telescopes. It contains office and computing rooms, a library and well-equipped workshops. The large telescope is in a steel dome with the necessary insulating material to diminish changes of temperature, and was constructed by Messrs. Sir Howard Grubb, Parsons and Co. The general appearance of the telescope in relationship to the dome is very satisfactory. The clockwork and electrical movements of the telescope and dome fulfil all requirements. The mirror of 74 inches is of pyrex, made by the Corning Company of New York. The grinding and figuring of the mirror were carried out at Newcastle, under Mr. Cyril Young's direction, by Mr. Armstrong, the very competent artist of the firm. The spectroscopy was constructed by Messrs. Adam Hilger, Ltd., and is admirably adapted to determine velocities in the line of sight for which the instrument will be generally used. Telescope and spectroscopy were thoroughly tested by Mr. R. K. Young. In Mr. Young, Mr. Hogg and Mr. Millman, Prof. Chant has an able, experienced and enthusiastic staff. We may look with confidence for an excellent output of work from the David Dunlap Observatory.

F. W. DYSON.

The Globe
June 1, 1935Christian Science Monitor
June 6, 1935

Montreal Herald, June 7, 1935

PUB. SPIRIT IN TORONTO

SOME of the most beautiful and the nation's finest buildings are now being erected in Toronto, and the city is becoming a more beautiful place than ever before.

The most beautiful in which the Queen City is rich is the David Dunlap Observatory, which contains the largest telescope in the world. It is a gift to the University of Toronto by Mrs. D. A. Dunlap and was dedicated to the memory of her husband, David Alexander Dunlap, who was a pioneer in the field of astronomy. The mirror is 74 inches in diameter, and the telescope is 74 feet long. The power of this great telescope is estimated at 100,000 times as powerful as the unaided human eye. It can see objects millions of light years distant, and is the most powerful telescope in the world.

THERE are some people who think that because charity is a good thing, it is a waste of money. They are wrong. Charity is a good thing, but it is not a waste of money. It is a way of giving to the poor and the needy, and it is a way of making the world a better place.

We can do much to help the poor and the needy. We can give them money, but we can also give them our time and our energy. We can help them to find a way out of their poverty, and we can help them to become self-sufficient.

To our men of wealth and to our men of business, we appeal. We ask you to help us to make the world a better place. We ask you to give us your money, your time, and your energy.

Not that we ask you to do this for the sake of the poor and the needy. We ask you to do this for the sake of the world. We ask you to do this for the sake of the future. We ask you to do this for the sake of the human race.

But no man who has more money than he needs should pass through this life without contributing to the betterment of the human race. We ask you to do this for the sake of the world. We ask you to do this for the sake of the future. We ask you to do this for the sake of the human race.

Wealth is a great blessing, but it is not a curse. It is a tool, and it is a tool that can be used for good or for evil. We ask you to use it for good. We ask you to use it to help the poor and the needy. We ask you to use it to make the world a better place.

The New Outlook, June 12, 1935

Canada's Greatest Observatory

THEIR was general congratulation on the occasion of the opening of the David Dunlap Observatory at Richmond Hill near Toronto, Friday, May 31st. The huge seventy-four-inch telescope is the largest in the British Empire and second largest in the world. The observatory is a gift to the University of Toronto by Mrs. D. A. Dunlap and was dedicated to the memory of her husband, David Alexander Dunlap, who was a pioneer in the field of astronomy. The mirror is 74 inches in diameter, and the telescope is 74 feet long. The power of this great telescope is estimated at 100,000 times as powerful as the unaided human eye. It can see objects millions of light years distant, and is the most powerful telescope in the world.

CANADA'S NEW OBSERVATORY

The Canadian public—at least that large section of it not greatly interested in astronomy—may be surprised to learn that yesterday, at Richmond Hill, the largest observatory in the British Empire, and the second largest in the world, was formally opened. This splendid contribution to the Dominion's scientific equipment has been made possible by the generosity of Mrs. Jessie Donaldson Dunlap, at whose expense it has been erected and presented to the University of Toronto in memory of her husband, the late David Dunlap.

A pleasing incident connected with the ceremonies yesterday was that Dr. C. A. Chant of the university staff, who is to be Director of the observatory, was celebrating the seventeenth anniversary of his birth. With the excellent equipment now at his command for study of astronomical phenomena, Dr. Chant will be a happy man. He has devoted the best years of his life to this study, and has attained high rank among the world's astronomers.

In this connection it should be noted also that another Canadian who has won renown in the realm of astronomy—research work is John Stanley Plaskett, who supervised construction and erection of the 72-inch telescope on Little Spanish Mountain, on Vancouver Island, B.C. As Director of the Dominion Astrophysical Observatory, to which this great reflector belongs, Dr. Plaskett has been honored by universities and scientific bodies.

Astronomy is one of the oldest of man's studies. Early in the world's story observatories were founded in Eastern countries. The first observatory in Europe was erected at Nuremberg in 1472. Now all countries have up-to-date telescopes, by means of which man is adding to his knowledge of heavenly bodies. In America the Yerkes Observatory at Williams Bay, Wis., and the Lick Observatory in California, are regarded as among the best in the world; though a 300-inch telescope being constructed for the California Institute of Technology will give a field many times greater than any other. In the neighborhood of Vesuvius there is an observatory for the special study of volcanic action. In England privately owned observatories carry on important work, and these indicate the keen interest of the individual in astronomical research.

In view of the universal nature of astronomical study, it is worthy of note that Canada will, with the addition of this observatory at Richmond Hill, be well to the fore. The donor, Mrs. Dunlap, is to be commended for her generosity in making possible this addition to the Dominion's standing in the world of science, and the University of Toronto must be congratulated on its opportunity to take a leading part in work aimed to provide for mankind more intimate knowledge of "the mysterious universe."

New Observatory Opened

Many Guests Attend Ceremony At Richmond Hill, Ontario

A great hemispherical dome of 80 tons surfaced with copper, rotated about the top of a circular stone structure. Two parallel shutters slid open and the mammoth 74-inch telescope that is to unfold the secrets of the heavens hove into view. A thousand invited guests looked on with awe as the telescope moved about pointed to the skies. The David Dunlap Observatory was ready to study the mysteries of astronomy.

Situated on a hill-top 800 feet above the sea-level overlooking a vast expanse of countryside that suggested something of the limitless possibilities of the explorations of the telescope in the decades and centuries to come, the David Dunlap Observatory was opened at Richmond Hill, Ontario, with a golden key, by its donor, Jessie Donaldson Dunlap, as a memorial to her husband.

Leading students of the heavens, amateur scientists, outstanding educationists, representatives of the church and state, and men and women prominent in various walks of life, gathered to triumph in the opening of the world's second largest observatory in the British Empire. Dr. C. A. Chant, director of the observatory, who for three decades had looked forward to this moment and who was celebrating his 70th birthday, evidenced the pride that was bursting within him as he saw the culmination of his dreams.

Every portion of the mirror had been worked out in theoretical contour to two one-millionths of an inch, said Cyril Young, of England, who explained the evolution of the big mirror, 76 inches in diameter, 12 inches thick and weighing 5,000 pounds. He recalled how it had been poured from a special type of pyrex by the Corning Glass Works, put into the annealing oven and sent to England to be ground and polished.

Journal of the Royal Astronomical Society of Canada, 1913.



PLATE VIII.

BEFORE THE ADMINISTRATION BUILDING.
C. A. Chant F. W. Dyson R. K. Young

PLATE VI.



AFTER THE FORMAL OPENING.
Sir Frank Tyson Mrs. Danday Mr. John B. Holden

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PLATE V.



MRS. DANDAY OPENING THE DOOR OF THE OBSERVATORY.

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PLATE VII.

WITHIN THE GREAT DOME.
The spectrograph attached to the mirror cell of the 74-inch telescope is seen.

mounting arrived in Toronto on October 15, 1933, and within two weeks was mounted on its pier.

There was some delay with the immense disc of glass for the main mirror of the telescope, but on June 21, 1933, a happy and interested party from Toronto saw the 2½-ton disc poured from a superior type of "pyrex" glass at the Corning Glass Works, Corning, N.Y. After remaining in the annealing kiln for three months it was examined, declared good and then shipped to England to be ground and polished. Within its 5-ton shipping case, on the second day of this very month, it was moved into the great dome through an opening made in the wall, and five days later the mirror was safely placed in its cell and then attached to the telescope. The telescope is now completed.

Before resuming my list I wish to make some acknowledgments. Upon the name of Mrs. Jessie Donald, Dunlap we need not dwell. Were it not for her outstanding benevolence there would be no observatory here. To her all the credit is due.

Then I must mention the name of a person who, I know, would strongly dissent if I told him I intended to do so. I refer to Mr. John B. Holden, the former business associate and friend of Mr. Dunlap and the legal adviser of the firm of Messrs. Dunlap. Mr. Holden is a faithful guardian of her interests, but he appreciated at once the fine quality of the memorial and his heartily co-operated in making the observatory worthy of its object.

Of the work of the architects, Messrs. Mathers and Holden, I need only say that their creation is acclaimed by everyone as beautiful, appropriate and effective. The contractors also, Messrs. Sullivan and Reid, the Ramsey Contracting Co., Jordan-Roberts of Brantford, the Dominion Bridge Co., as well as the numerous staff of the Superintendent of Buildings of the University, all deserve commendation. One must specially mention the English firm, Sir Howard Grubb, Parsons and Co., which supplied the great telescope and the steel dome. It is a miracle of mechanism, and our relations with the firm have been very cordial. It is much to be regretted that Sir Charles Parsons, the distinguished head of the firm, did not live to see the telescope completed. He had a lively interest in it. I have been intimately associated with the great project in all its details, and everyone employed on it seemed to me to realize that it was a very special undertaking which demanded his best efforts.

I wish also to refer to the assistance rendered by the Royal Astronomical Society of Canada, the Council of which has loaned us the valuable technical portion of its library. We hope the Society will find it even more accessible here.

I may remark that the first three volumes entered in the accession book of our library bear the autograph of D. A. Dunlap.

Finally let me refer to the ability and enthusiasm of my colleague, Professor R. K. Young. No mathematical problem which we encountered was too hard for him to solve, no mechanical intricacies were too difficult for him to master and no labour was too long for him to undertake and carry to a conclusion.

Friends have remarked to me that I have spent a long time in my endeavour to secure this observatory. True it is that thirty years is a considerable time, but it is as nothing compared to the decades and centuries during which a stream of astronomical research will proceed from this noble memorial, bringing honour to the name it bears and to its donor, and also conferring distinction on our land.

The chairman then asked for a few words from the one who was directly responsible for the construction of the great telescope.

MR. CYRIL YOUNG, General Manager of Sir Howard Grubb, Parsons & Co., Newcastle-on-Tyne, England.

When I paid my first visit to the Observatory last Tuesday I could not help feeling proud that the firm of Sir Howard Grubb, Parsons & Co., with whom I am associated and whom I have been very kindly invited to represent today, should have had the honour of building the 74-inch Reflector, the principal instrument in this observatory, so kindly provided by Mrs. Dunlap's munificence.

In designing and constructing the Telescope and Dome, we have had to collaborate with Dr. Chant and Dr. Young, and I am glad to say that our relations throughout have been most happy. We have been very fortunate in having on the spot Mr. F. J. W. Bell, who so ably looks after Messrs. Parsons' interests in Canada. I wish to take this opportunity of acknowledging the assistance given by my chief, Mr. Bedford, who is here today. He has given us every facility to obtain the help and aid of the experts at the Heaton Works who could in any way contribute to the successful manufacture of the Telescope. This assistance has been most generously given and I am sure they would feel proud if they could see today the results of their contribution.

I must specially mention the name of Mr. Armstrong who was responsible for working the 74-inch mirror. When I explain that every portion of the surface of an astronomical mirror should be worked true to its theoretical contour to within 2-one-millionths of an inch you will understand that the task of making a successful mirror is a very arduous one.

Some weeks ago Dr. Young paid a special visit to Newcastle-on-Tyne in order to test Mr. Armstrong's work, and from the results of these tests we have every reason to believe the mirror to be a very good one.

I cannot conclude without a reference to the enterprise of The Corning Glass Works, which manufactured the "pyrex" disc for the mirror. This was the largest disc they had made up to that time, although since then they have produced several larger, including one of 200 inches diameter. The greatest credit is due to Dr. O. A. Gage and his colleagues at Corning.

My only regret is that Sir Charles Parsons did not live to be here today.

The next speaker was introduced by the chairman as a distinguished representative of British astronomy.

SIR FRANK DYSON, K.B.E., Astronomer Royal 1910-1933

In the first place I should like to congratulate the University of Toronto on the beautiful observatory and magnificent telescope given to them by the generosity of Mrs. Dunlap. Then I think Mrs. Dunlap should be congratulated on the worthy memorial she has made to her husband Mr. David Dunlap. As the Latin poet says, "It is a memorial more lasting than brass," and Mr. Dunlap, who was keenly interested in astronomy, would have desired no better tribute. Lastly, Professor Chant is to be congratulated on the consummation of his hopes on this day, his 70th birthday. Prof. Chant has been a great power in furthering the study of astronomy in Canada. Besides the number of students he has taught he took a large share in establishing the Toronto Astronomical Society which in after years became the Royal Astronomical Society of Canada and has spread right across the Dominion.

I have had the opportunity of seeing the telescope. Its general appearance is most workmanlike. I was very glad the contract was given to the firm of Sir Howard Grubb, Parsons and Co. Sir Charles Parsons was the famous engineer who by his invention of the steam turbine came next to James Watt in getting power out of steam. Sir Charles was the son of Lord Rosse, who made the biggest reflecting telescope of the nineteenth century, and was interested in the manufacture of optical glass. Later he associated himself with the famous firm of Sir Howard Grubb, which had made many large telescopes, and thus acquired the experience of Mr. Cyril Young, who became the manager of the telescope works at Newcastle, and of Mr. Armstrong who was the first artist in the figuring of mirrors and lenses. I am quite sure the 74-inch mirror is a very perfect instrument. Both the mirror and the spectroscopic which was constructed by the firm of Adam Hilger, have passed the severe scrutiny of Prof. R. K. Young.

Then I cannot imagine a more beautiful site for an observatory than this. It is far enough from Toronto to escape smoke and glare and sufficiently near for the teaching purposes of the University.

But a telescope is not much good without suitable astronomers to work it. I should like to congratulate Prof. Chant on his staff. I have only recently met Dr. Young but know of his work at Lick and Victoria. But I had the pleasure of meeting Prof. Hogg at Victoria. Prof. Plaskett with whom I was not staying took me to the 72-inch. Naturally Prof. Hogg was there. I was surprised to see Mrs. Hogg, because I knew she was a Harvard graduate and was working on Cepheid variables in the cluster Messier 2. But I was surprised to see a bride a few weeks old sleeping in the dome. That shows the enthusiasm of her parent. With enthusiasm and skill great things will be accomplished, and I look forward with every confidence to the future of the David Dunlap Observatory.

President Cody then requested Professor Curtis, of the comparatively near University of Michigan, to address the gathering.

DR. HEBER D. CURTIS, Director of the Observatories of the University of Michigan

It is a great pleasure and honour that it is mine today in bringing to the University of Toronto the congratulations and good wishes of the University of Michigan, a sister institution of learning lying on the other side of that imaginary line that divides two people who are so essentially the same in every respect.

In all our great universities there are groups of men who attempt to find out things about this universe and to translate and pass this knowledge on to the rest of the world. The geologist teaches us the lore of things beneath the crust so that we may mine our coal and iron; the botanist learns to know the life and needs of plants and makes this knowledge help us with our crops; the engineer builds our bridges and teaches us how to make steel, steam and electricity our slaves; the historian or economist shows us the errors of the past and how we may avoid them in the present. But I should not personally care to be the historian; we are apt to believe implicitly the lessons that the geologist, the botanist, or the engineer teaches us, but we humans seem to go right on making the same old mistakes in so far as the lessons of history are concerned.

The astronomer, by and large, renders no such utilitarian service to mankind and, on the whole, I am rather glad that this is so. He gives us rather something that enlarges our souls in a knowledge of the universe without; the inflexible and beautiful laws of the greater spaces; he brings to us aggregates that are so inconceivable in mass, or distance, or in the time factor that we can not go far wrong in applying to this external universe the tremendous adjective, —infinite. And it is a wonderful thing that our souls may thus come in contact with the infinite, for we do not live by bread alone.

It is unnecessary to repeat here the details of telescopic equipment and power that make the David Dunlap Observatory so notable an adjunct to the scientific work of the University. We of the field of science prophesy for the new observatory a long and useful life. In this fine plant there will be carried on in all the decades to come researches of value on the world of the stars without—their elements, constitution, and motions. Many of the studies thus made will doubtless need too highly technical and abstruse for the man in the street, yet all will have their individual parts in that greater picture of the universe that the astronomer is trying to paint for the world.

It is thus an unusual pleasure for me to join with this audience in congratulations to the University of Toronto at the dedication of this fine plant and new tool of research; our felicitations as well to Dr. Chant on this his natal day that marks also the successful fruition of his decade-long dreams made possible for the University an acquisition so notable as the David Dunlap Observatory.

The next speaker was the representative of the great Harvard University at Cambridge, Mass.

PROFESSOR HARLOW SHAPLEY, Director of the Harvard Observatory

Dr. Shapley congratulated Canada on the opening of such a magnificent observatory. As judged by the quantity and quality of equipment, Canada was now the outstanding country in the astronomical world. The only other place of comparable size in which there was similar interest was South Africa, and there the interest was not native, but rather due to outsiders who had set up observatories in that country owing to its situation in the southern hemisphere and to its very fine astronomical climate. He particularly stressed that in an observatory such as this, the balance should be maintained between popular interest in astronomy and freedom of the staff to pursue the researches that such generous equipment warranted. He welcomed the observatory as a further important element in the making of western hemisphere astronomical research a thing for the future generations to think about. Dr. Shapley complimented Mrs. Dunlap on the magnificence and suitability of her contribution to science. He also congratulated Dr. Chant on his birthday, and especially on the "pretty trick" he had received as a birthday present. He feared Dr. Chant's tastes might become so extravagant that in the future, suitable birthday gifts might be really difficult to supply.

Dr. Shapley also spoke of his personal interest in Toronto as an astronomical centre in that three of the astronomers here, Frank S. Hogg, Helen Sawyer Hogg, and Peter M. Millman, had been "doctored" by him at Harvard University. He thus felt he had a greater interest and responsibility here than at any other observatory but his own, and suggested that this mutual interest be a basis of constructive co-operation.

The last speaker of the afternoon was introduced as a graduate of the University of Toronto, who had achieved distinction in the astronomical world by his work at the observatories at Ottawa and Victoria.

W. E. HARPER, M.A., Acting Director of the Dominion Astrophysical Observatory, Victoria, B.C.

It is certainly a great privilege and honour to be invited to participate in the formal ceremonies connected with the dedication of this great observatory. There is, moreover, a peculiar pleasure that comes to myself, and I am sure to all who know Dr. Chant well, in the thought that the dream of his life has at last been fulfilled. The establishment of this observatory with its marvellous telescope for studying the secrets of the heavens is a fitting climax to his professional career in the University.

As Editor also of the *Journal of the R.A.S.C.* since its inception over thirty years ago he has devoted himself unselfishly to building up astronomical centres across Canada where leadership is given to the study of the stars, the oldest of the sciences. It is not at all surprising, then, that in addition to the formal greeting which has just been read from the General Executive of the Society there should come from these various centres across Canada congratulations and good wishes to Dr. Chant and to the University which receives this splendid gift.

I understand that Dr. Chant, who seems to remain ever young, will now be free to devote his whole time to astronomical research. He will have associated with him several who have been co-workers with myself at Victoria. It might, therefore, be thought that we would cherish some animosity towards him and the University for casting away some of our best workers in the persons of Dr. Young, Dr. Millman, Dr. Hogg and his talented wife. Such is not the case, for while we regret losing them we recognize that they have greater opportunities here. Moreover, in the early days, before we could secure our own, Dr. Chant was kind enough to loan us much needed equipment and he has ever been a good friend of the Victoria observatory. Surely, then, we can in some measure attempt to repay his kindness by allowing him and the University to have some of our best workers.

This telescope, as has just been mentioned, has a diameter two inches greater than ours at Victoria and now ranks as the second largest in operation. By reason of this taking precedence in size over our own it might be suspected that we would be envious. Again, such is not the case. Like a loyal British Columbian, I shall only say that with the wonderful climate we enjoy and the consequent high quality of our seeing, we feel we can afford to accept a handicap of a mere couple of inches. That is a mere trifle, you know.

There is a well-known saying that one may judge of the degree of the civilization of any nation by the provision which that nation makes for the study of astronomy. Now that you good people of the east have a really worthwhile observatory we westerners will believe there is hope for the east yet! Seriously though, speaking for the Dominion Astrophysical Observatory, we welcome this new observatory into the field of astronomical endeavour.

We expect great contributions from this observatory and I am sure we shall not be disappointed. It is a great thing to be able to make possible such achievements and greater still to be the one willing to do so.

Mrs. Dunlap may well be assured that for generations to come people will point with pride to this establishment as a worthy example of how to perpetuate the memory of a loved one. We may well use scriptural language and say that it will be as a city that is set upon a hill which cannot be hid or as a lamp which giveth light to all the world.

May her example serve to stimulate others to apply their wealth to similar memorials seeking to discover Truth, for "The truth shall make you free."

II.

SPECIAL CONVOCATION OF THE UNIVERSITY

THE second impressive ceremony of the day took place in the evening when the University of Toronto called a special convocation to confer honorary degrees on six persons who had made notable contributions to astronomy. The degree of Doctor of Laws was conferred on Mrs. Dunlap, Sir Frank Dyson, and Prof. Chant; and the degree of Doctor of Science on Dr. Harlow Shapley, Dr. V. M. Slipher, and Mr. W. E. Harper. The Chancellor of the University, Sir William Mulock, was unable to be present because of illness, and President Cody occupied his chair. Present on the platform, besides university officials, the recipients of degrees, and those who read the citations, there were His Honour the Lieutenant-Governor, Dr. Bruce; the Hon. L. J. Simpson, M.D., Minister of Education, and His Deputy Minister, Dr. Duncan MacArthur; the Hon. J. A. Faulkner, M.D., Minister of Health; His Worship the Mayor of Toronto, Mr. James Simpson; Sir Joseph Flavelle, and Mr. W. L. Mackenzie King.

The candidates for degrees who would ordinarily have been presented to the Chancellor by President Cody were presented by various prominent members of the faculty who read the president's citations. After each candidate had been given the degree, he inscribed his name in the Golden Book of the University, which contains the names of all the recipients of honorary degrees. The ceremonies were enhanced by the fact that the leader of the opposition in the House of Commons of Canada, Mr. Mackenzie King, was present and spoke brilliantly though briefly.

During the course of the evening, President Cody spoke with regret of the retirement of Professor Chant as head of the Department of Astronomy and Director of the Observatory, but stated that he would remain on the staff in an honorary capacity. Dr. Cody announced the appointment of Dr. Raymond K. Young as the new Director of the Observatory.

Preliminary to the conferring of degrees the President remarked that this was a red-letter day in the history of the University of Toronto. "It marks," said he, "a decisive advance in our scientific equipment. Today has been formally opened the David Dunlap Observatory, a magnificent gift to the University made by Mrs. Dunlap in memory of her honoured husband, that his name may be perpetuated in connection with the high realm of astronomical research, and that this University may share with other observatories throughout the world the task of investigating the field of stellar velocities and spectral photometry and in similar co-operative labours."

He then called on Professor E. F. Burton, head of the Department of Physics in the University, to present Sir Frank Dyson for the degree of LL.D. Professor Burton read the following citation:

Mr. Chancellor, I present to you one who has just retired from the historic post of Astronomer Royal and director of the Royal Observatory at Greenwich, an institution founded in 1675 by that Stuart monarch with a scientific bent, King Charles II.

Frank Watson Dyson was born in an English Baptist paragon and has never forgotten his fundamental religious associations, as is witnessed by his acceptance in 1921 of one of the vice-presidencies of the British and Foreign Bible Society. Educated at Bradford Grammar School and Trinity College, Cambridge, he graduated with the mathematical law honours on Second Wrangler and Smith's Prizesman. In 1894 he became chief assistant at the Royal Observatory, Greenwich, and in 1899 secretary of the Royal Astronomical Society. In 1905 he resigned these duties to become Astronomer Royal for Scotland. This post he held till 1910, when he was appointed Astronomer Royal at Greenwich. For twenty-three years he held this high office retiring in 1933 at the remorseless demand of his age limit.

Many honours have come to him in recognition of his astronomical researches. He was knighted in 1915 and made a K.B.E. in 1926. He was elected a member of the Royal Society in 1901. Edinburgh, Oxford and Cambridge have given him their highest academic distinctions.

From 1891 onward he has published innumerable scientific papers. Three outstanding features of his scientific achievements may be mentioned: (1) at Greenwich Sir Frank's chief work lay in the astronomy of position—the determination of the proper motions of the stars. To determine these motions there is need of a clearly defined programme, great persistence, and observations of the utmost precision. This kind of work may lack the spectacular features which appeal to the many, but it constitutes the very foundation on which astronomical progress rests.

(2) Sir Frank accompanied theclipse expeditions at 1900, 1901, 1905, 1905 as Greenwich chief assistant and his services in connection with the work that he would have no equal in astronomy.

(3) When Eratosthenes in 1915 published his paper on general relativity it was at once recognized that the new theory had an important bearing on astronomy. Einstein himself had suggested that the tests of its predictions were astronomical in nature and that one of them could be applied to observations made at a total eclipse. Two English expeditions were planned, one when the Great War was raging to observe a total eclipse in May, 1919, the other went from Cambridge to Arica under Sir Arthur Eddington, the other from Greenwich to Brazil under Sir Frank Dyson. Both were successful and obtained evidence in support of Einstein's theory. This brilliant expedition brought great fame to English astronomy and in making them Sir Frank Dyson took a leading part.

We honour our University in adding to our company this great man of singular modesty and singular achievement.

I, therefore, present to you Mr. Chancellor, Sir Frank Watson Dyson, K.B.E., F.R.S., LL.D., D.Sc., former Astronomer Royal, to receive at your hands the degree of Doctor of Laws, *honoris causa*.

Hon. N. W. Rowell, K.C., presented Mrs. Dunlap for the degree of LL.D. in the following words:

This University has been glad to invest with its highest academic honours certain notable women whose services to others have richly deserved public recognition. To-night we add another to this selected group—the gracious lady whose wife's affection and wise benevolence have made possible the erection of this great observatory, Mrs. Jessie Donald Dunlap.

From those to whom much is given, much by divine law is required. Mrs. Dunlap has ought faithfully and with discrimination to discharge her duties of stewardship in the use of her possessions. It is not for me this evening to mention the manifold causes she has generously helped. I speak only of some of the educational institutions which have been able to advance because of the munificence and which her husband and she herself have given to them. St. Andrew's College, the Toronto General Hospital, the Medical and Psychological departments of this University.

The climax of her helpfulness to the University of Toronto has come in the provision of the observatory and its site as a memorial to her late husband, David Alexander Dunlap. She shared her husband's interest in astronomy, that science which in our lambs and exalts the mind and soul of man, and responded with considered, yet prompt generosity to the suggestion that

she should do this great thing for the furtherance of astronomical research in Canada. Today she has handed to the University a superb addition to its scientific equipment, in the confident hope that through these new facilities the confines of astronomical knowledge may be ever enlarged, the mysteries of the universe may become more intelligible, and the starry heavens above us, as Immanuel Kant said, in conjunction with the moral law within, may come to bear witness to Him "who is before all things and in whom all things consist".

The University of Toronto wishes to mark its appreciation of Mrs. Dunlap's public philanthropies to this community and its profound gratefulness for her many benefactions to this institution by conferring upon her the highest academic distinction within its power.

I, therefore, present to you Mrs. Jessie Donald Dunlap that she may receive at your hands the degree of Doctor of Laws, *honoris causa*.

Professor Lachlan Gilchrist, President of the Royal Astronomical Society of Canada, then presented Professor C. A. Chant for the degree of LL.D., in the following statement prepared by President Cody:

It is not the happy lot of many to see their dreams come true and attain the goal of their life's effort. To Professor Chant this rare experience has come. Today, his seventeenth birthday, is the crowning day of his long career. It is our privilege to rejoice with him in his rejoicing.

He graduated from this University in 1890. It was my good fortune to be a fellow-student, graduating one year earlier. We were both officers in a fellow-student, graduating organization—the University College Literary and Scientific (now Athletic) Society. Another member of our committee was George Howard Ferguson. A year after his graduation, Dr. Chant was appointed Fellow and next year Lecturer in Physics. When the new department of astrophysics was established in 1907 he was made associate professor, and in 1918 he became full professor. This post he has held till the end of his career, when he retires under the age limit. I am glad to say that the Board of Governors has appointed him Professor Emeritus of astrophysics and Emeritus Director of the Observatory, as well as its Honorary Librarian and director of publications.

His whole life has been devoted to the teaching of astronomy and to emphasizing its value as a cultural subject in a liberal education and as a fine discipline for the advanced student. From the beginning of his work he has urged that the University should have an observatory, and so make some fresh contribution to our knowledge of astronomy. In this programme for astronomical advance in Canada, great interest was taken by Mr. David A. Dunlap, an enthusiastic amateur in the field of astronomy and a member of the Royal Astronomical Society of Canada, one who believed that (in Kepler's phrase) astronomers were thinking God's thoughts after Him. Some time after his death in 1924, when Professor Chant suggested to Mrs. Dunlap that she provide an observatory as a fitting memorial of her husband, she gave a sympathetic and munificent response—and the issue has been the great observatory and telescope dedicated to day to the cause of astronomical research.

There are other phases of Professor Chant's work that deserve to be mentioned. (a) He was one of the earliest experimenters in wireless. While he was an undergraduate, Hertz in Canada discovered the method of producing electric waves and thus opened the way for the future development of wireless telegraphy and radio. In 1895 when Mr. Chant was a lecturer in physics he constructed the necessary apparatus and first demonstrated to a Toronto audience the properties of these waves. In 1899 before the Canadian Institute he showed, perhaps for the first time in Canada, how to send and record a wireless message.

(b) He has played an important part in the development of the Royal Astronomical Society of Canada. He joined the original society in 1892, and from 1904-7 was its president. In 1905 he edited its transactions. In 1907 he began the publication of the *Journal of the Society*, which is now in its 29th volume and is sent to all parts of the world. At his suggestion local centres of the Society have been organized in various parts of Canada.

(c) He has written admirable and widely-used text books. In collaboration with Dr. F. W. Merchant he prepared the *High School Physics*. The first edition appeared in 1911 and a revised edition in 1923. All the Canadian provinces have used it as a text book and it has had a wide circulation in the United States. More than 275,000 copies have been issued. The same authors have published three other school text books which have had a combined circulation of 140,000.

In 1928 he wrote a small book on astronomy called "Our Wonderful Universe". This has been published in England, Canada and the United States, and has been translated into German, Czech and Polish.

In 1933, in collaboration with Professor E. F. Burton, head of our department of physics, he published a text book of college physics, and this is being used in many colleges in Canada and in the United States.

(d) He has made five expeditions to observe total eclipses of the sun. In the 1922 expedition to Western Australia evidence was secured in verification of Einstein's theory.

Surely at the close of his long academic career in this University and on the day of the opening of the Dunlap Observatory—his realized dream—we do well to honour our own Professor Chant, a scientist "indeed in whom is no guile."

I present to you, Sir, Clarence Augustus Chant, M.A., Ph.D., Fellow of the Royal Society of Canada, Professor of Astrophysics in this University, that he may receive at your hands the degree of Doctor of Laws, *honoris causa*.

Professor Harlow Shapley was presented for the degree of D.Sc. by Professor J. L. Synge in the following words:

Equipment for astronomical research has nowhere been brought to a higher degree of efficiency than in the neighbouring Republic. Some of the finest observatories of the world are there situated. Mt. Wilson Observatory has at the moment the largest telescope in existence, and the glass is still cooling for the marvellous 200-inch mirror which the Mt. Wilson organization will some day add to its instruments for observing the heavens.

One of the most famous astronomers in the United States I now present to you in the person of Dr. Harlow Shapley. From the University of Missouri from which he graduated, he went to Princeton as a Fellow. From 1914 to 1921 he was a member of the staff of the Mt. Wilson Observatory. From this he came to Harvard in 1921 to be Director of the Observatory and Paule Professor of Astronomy, and has proved himself a worthy successor to the late Edward C. Pickering. His researches have been in the fields of photometry, spectroscopy, and cosmogony.

While he was at Princeton he made notable contributions to the theory of variable stars, especially those known as eclipsing variables. This work he greatly extended at Mt. Wilson, and was able to utilize it in determining hitherto unmeasured distant portions of the universe. This led on to investigations into the nature and distance of star clusters, and their relation to our system of stars. He gained new views of the structure and dimensions of our Milky Way or galactic system, and revealed that our universe was infinitely greater than we had previously believed it to be. What a being is man, able to understand and formulate theories about this vast yet intelligible universe!

Professor Shapley is a member of many scientific societies, including the National Academy of the United States, the American Astronomical Society, the Royal Astronomical Society of Canada, and the Royal Astronomical Society of London.

He was Lowell Lecturer at Boston in 1922, Exchange Lecturer in the Belgian Universities in 1926, Halley Lecturer at Oxford in 1928, and George Darwin Lecturer in 1934. He has been awarded the Draper Medal of the National Academy in 1926; the University Medal, Brussels, 1926; the medal of the Society of Arts and Sciences 1931; and the gold medal of the Royal Astronomical Society of England in 1934.

Among his early publications were papers on astronomy in Horace and in Lucretius. From Lucretius may be quoted (in translation) lines which fairly describe our distinguished visitor: "Therefore the living force of his soul glided the day; on he passed far beyond the flowing walls of the world, and traversed throughout in mind and spirit, the immeasurable universe; whence he returns a conqueror, to tell us what can, what cannot, come into being."

We are grateful for his presence with us to-day.

I present to you, Mr. Chancellor, Harlow Shapley, one of the world's great astronomers, to receive from you the degree of Doctor of Science, *honoris causa*.

Professor Reynold K. Young then presented Director V. M. Slipper, of the Lowell Observatory, Flagstaff, Arizona, for the degree of D.Sc.:

For centuries the fiery planet Mars has been an object of interest to astronomers and laymen alike, and as telescopes became larger in size and better in quality strange details were detected on the surface of the planet. In 1877, the Italian astronomer, Schiaparelli, reported that he had seen faint straight markings on the planet, and after some years of observation he drew a map of the surface of Mars on which was a network of lines. These were so straight and regular that he called them *canali* (channels), but the world soon straight and regular that he called them *canali* (channels), but the world soon called them canals. Percival Lowell, member of a well-known Boston family and a distinguished graduate of Harvard, was greatly attracted by these observations and determined to devote himself and his means to investigations on Mars and the other planets. After a search for the best site he established the Lowell Observatory at Flagstaff, Arizona, at an altitude of 7250 feet above the sea, and gathered about him a number of able and enthusiastic young men.

One of these was Vesto Melvin Slipper, a native of the Hoosier State, Indiana, and a graduate of its State University. He joined the staff of the Lowell Observatory in the year of his graduation, 1901; he became assistant director in 1915, and director in 1926.

Mr. Slipper's special field of research is spectroscopy, and his investigations on the planets, not merely on Mars, have been of outstanding value. With skill, ingenuity and abounding patience, he has secured much information regarding the atmosphere, the surface conditions and the rotation of the planets. One of the latest results deduced from his spectrum photographs is that there is much of the poisonous gases, ammonia and methane (or marsh gas) in the atmosphere of the outer planets of our solar system.

But he has not confined his attention to our family of planets. Indeed he has attacked some of the most difficult problems of the great universe of stars beyond. From his early spectra of the stars he found evidence of the presence, throughout space, of clouds of calcium vapour, a discovery now recognized as of great importance.

His spectrum photographs were among the first to lead to the belief that the diffuse nebulae owe their illumination to the bright stars in their neighbourhood. By the inclination of the lines in the spectra of some spiral nebulae he was able to demonstrate that such nebulae are moving outward on the arms of the spiral.

In recent years we have been told that our universe is expanding, that every part is moving away from every other part. This result is deduced from Einstein's theory of relativity. One reason for the general acceptance of this view is, that, before the general theory of relativity was announced, Mr. Slipper had actually obtained photographs which showed that some of the spiral nebulae were moving in space with speeds of several hundred miles a second.

In most of his researches his work has been that of pioneer, and has proven to be of fundamental importance in formulating our theories of cosmogony.

He has been honoured by election to many scientific societies, by receiving honorary degrees, and by the award of the Lalande Prize and Gold Medal from the Paris Academy in 1919, and the Gold Medal of the Royal Astronomical Society in the year before Dr. Shapley received it, 1933. He has also been the George Darwin Lecturer in 1933.

I, therefore, present to you, Mr. Chancellor, Vesto Melvin Slipper, Ph.D., LL.D., Sc.D., director of the Lowell Observatory and a great interpreter of the starry universe, that he may receive at your hands the degree of Doctor of Science, *honoris causa*.

Professor Samuel Beatty, head of the Department of Mathematics in the University, presented Mr. W. E. Harper, of Victoria, B.C., for the degree of D.Sc., in the following words:

Canadians have taken their part in recent years in the work of astronomical research. In a new country like ours, it has not been unnatural that emphasis in the field of science should be placed on subjects that seemed to be of immediate utility. Yet pure science has made great strides on its own account, and among pure sciences astronomy has gained increased prestige and aroused fresh interest by reason of recent spectacular discoveries which have vastly enriched our knowledge of the universe. The life work of Professor Chant has not been in vain, and astronomy has come into its own. Two of our graduates, Dr. J. S. Plaskett, formerly director of the Dominion Astrophysical Observatory at Victoria, B.C., and his son, Professor Harry Plaskett, now Savilian Professor of Astronomy in the University of Oxford, have made great contributions to this fascinating and overawing science.

To-night I present to you another of our graduates whose astronomical work has been of signal value. He is in a sense Dr. Chant's "son in the astrophysical faith." Professor and pupil stand side by side in this Hall to-night to receive the highest honour their Alma Mater can bestow upon them.

William Edmund Harper came from that Ontario County of Bruce. From the Owen Sound Collegiate Institute he came to this University in 1902 and took the course in Mathematics and Physics. At Professor Chant's request, the Senate of the University in 1905 instituted a new division in this department, known as Astronomy and Physics. It was established in time for the graduating class of 1906, and Mr. Harper was the first graduate in this section. In order to encourage the study of astronomy, the Royal Astronomical Society of Canada had offered a gold medal to the candidate obtaining highest standing in the new course. The first award was made to Mr. Harper. The day of his graduation he received notice that he had been appointed to the staff of the recently founded Dominion Observatory at Ottawa. There he remained until 1919 when he was transferred to the new observatory at Victoria, B.C.

For the last twenty-nine years Mr. Harper has laboured effectively in the field of astronomical research, especially in astrophysics. With great skill and patience he has taken and measured thousands of photographs of the spectra of the stars, and from them the velocities of the stars and other wonderful results have been determined. He has discovered a large number of binary stars, which are pairs of suns far out in the depths of space revolving about each other. He has, it is said, computed more orbits of such systems than has any other man.

In co-operation with Dr. R. K. Young, of our own department of astrophysics, Mr. Harper determined, from their spectra, the absolute brightness, and hence the distance, of over 1100 stars, a great and important achievement. Only on such observations can an intelligent view of the universe be based.

Mr. Harper is a Fellow of the Royal Society of Canada and a past president of the Royal Astronomical Society of Canada. He has done much to popularize astronomy by his articles in newspapers and magazines, and by lectures and radio addresses. He is now the Acting-Director of the Astrophysical Observatory at Victoria—over which Dr. J. S. Plaskett presided for many years.

The work of this alumnus of our University has brought distinction to himself, to his native land and to his Alma Mater, who now greets and honours her academic son.

I present to you, William Edmund Harper, that he may receive from you the degree of Doctor of Science, *honoris causa*.

ADDRESSES BY THE RECIPIENTS OF DEGREES

The chairman then invited the recipients of degrees to address Convocation, and the first to respond was Sir Frank Dyson, who spoke as follows:

I am commissioned by Mrs. Dunlap to say how highly she appreciates the great honour you have conferred upon her by giving her the degree of Doctor of Laws. She wishes me to express her thanks for the many kind things which have been said of her and to her. For myself, I am very proud to be a Doctor of Laws of your famous University. I know that two of your graduates have been Prime Ministers of the Dominion of Canada. But I have a peculiar gratitude to the University of Toronto. The Medical Faculty, Banting, McLeod, Collop and Best, by the discovery of insulin, prolonged my sister's life for many years. As regards astronomy, nearly all the Canadian

astronomers, Chant, Meldrum Stewart, Young, Hogg, Millman and the Plasketts—father and son—were graduates of the University of Toronto, and many of them Prof. Chant's pupils. Now you have this magnificent as the Greenwich Observatory, which has been of service to astronomy for 260 years. As one who has been many years at Greenwich I am glad to come as a representative of the English astronomers but am very sorry to miss distinguished colleagues, Eddington and Jeans, found it impossible to leave England. The honour you have conferred on me is all the greater because it is associated with Mrs. Dunlap and the distinguished representatives of astronomy from Canada and the United States.

The chairman then called on Dr. C. A. Chant, who said:

My first wish is to thank the Senate for conferring on me the high degree of Doctor of Laws, here, in the presence of so many of my friends and in such distinguished company.

As has been intimated by the President, I have come to the close of a long period as an instructor in the University, and perhaps I may be allowed to offer a few reminiscences.

This thirty-first of May has been a wonderful day in my life. I suppose it might well be marked by a red letter or a golden number, as of old, and yet nothing of that sort is needed to remind me of it or for my complete satisfaction.

In 1891, a year after graduation, I returned to the University as Fellow in the Department of Physics, at \$500 for the session. One year later I received a permanent appointment as lecturer at \$800. During the operation of moving to our new and lovely home up at the Observatory I came across the letter from the Assistant Provincial Secretary informing me of that appointment; also a similar letter one year later stating that the salary had been increased to \$900. At the present time that salary looks small, and I suppose it was, but I had no complaints whatever to make. I was glad of the opportunity to get back to my Alma Mater.

Thus began my association with the young men and women from the high schools of Ontario who year after year came up to the University. Surely they are the finest in the world! To assist in moulding their minds and characters has been my great privilege. A teacher can have no greater joy than to see his students win high place in the world; and there have been so many classes which have gone forth in the last forty-four years that everywhere I am greeted by my former students.

Then also I entered on a life-time of the study of science and, as Charles Kingsley has remarked, "Science, like virtue, is its own reward." While that is true of all sciences, I think it is especially so of astronomy.

I think my mind was first turned in that direction when I was at the high school, by reading a school book entitled "Geography Generalized," by Robert Sullivan, of Dublin, Ireland. I have the book still. It has my eldest brother's name in it, with the date 1870, although this is a copy of the 27th edition which was issued in 1861. It contains a well-arranged and very readable introduction to astronomy. I note in the book some statements indicative of the epoch. The author continues to give the name Herschel to the planet next beyond Saturn which even then was known on the continent as Uranus. He also gives the distance of the earth from the sun as 95 million miles, instead of 93, and the velocity of light at 192,000 miles per second instead of 186,000.

Soon after being appointed to the University staff I joined the Astronomical and Physical Society of Toronto, which in 1903 became the Royal Astronomical Society of Canada. There I met a number of practical observing astronomers, who further stimulated my interest in astronomy. I was president of the Society for four years and formed some valued friendships with the members. I venture to hope that I have been of service to the Society.

In this way I was led to realize the slight attention paid to astronomy in the University and the lack of means for practical instruction, and I looked about for some way to improve matters. I suggested to my colleagues in Mathematics and Physics that I be given the instruction in Astronomy and they readily agreed. This was nearly thirty years ago, and it is interesting to look back on the growth of the department of astronomy. As the years went by my hopes for equipment for the University broadened, but I had not dared to expect anything so fine as that which, thanks to the great generosity of Mrs. Dunlap, we now possess. The completion of the David Dunlap Observatory has exceeded my rosiest dream. In the coming years, yes, in not many years, I predict with the greatest confidence, its staff will place it in a foremost position in the world of astronomy.

I conclude with one more word. Nothing has pleased me quite so much as the very generous words of appreciation of my colleagues in the University, those among whom I have lived ten, twenty, in some cases, over forty years.

The next speaker to address the gathering was Dr. Harlow Shapley.

Dr. Shapley, in thanking the University for this honour and paying further tribute to Mrs. Dunlap, stressed that the completion of the observatory was only the starting of its life and work. If it were to fulfil the purpose of its foundation, it must with care and diligence select and pursue active and fruitful research programmes. He spoke briefly of some of the difficulties and limitations encountered in attempts to solve the various problems of the universe, and jokingly condemned the universe for being so uncooperative with research workers. He referred particularly to such difficulties as the Mocking research workers. He referred particularly to such difficulties as the Mocking research workers. He referred particularly to such difficulties as the Mocking research workers.

up of regions of the Milky Way with absorbing matter and suggested that

Omnipotence should clean up the Milky Way of cosmic debris, it was all the

Yet, given the universe in its present uncooperative form, it was all the more urgent that we devise means of learning its structure in spite of difficulties. He suggested four chief fields in which he thought the equipment of the observatory and experience of its staff might produce profitable results. These comprised the determination of spectroscopic binary orbits, both for statistical analysis of characteristics of the stars; the quantitative analysis, some of the more accessible members of this group; the study of star from the astrophysical point of view, of stellar spectra; the study of star clusters and the study of the cosmic debris, as found in meteors, asteroids, and general absorbing matter in space.

The chairman then asked Dr. V. M. Slipper to speak, which he did in the following words:

I wish to express my sincere thanks to you and to the University of Toronto for honouring me with this degree. This is indeed the greater honour because it is given by an internationally great University.

I deeply appreciate such approval of my scientific work. If I have in some measure deserved this recognition it is largely because of the favourable conditions under which I have been able to study the heavens at Mt. Wilson. The Lowell Observatory where it has been my good fortune to do my astronomical work is well equipped instrumentally and is favourably situated astrophysically. It is located in the dry climate of the southwestern United States, is 7250 feet above sea level and has on the nearby mountains an auxiliary station at 11,500 feet altitude.

Thus surrounded, and possessed of some preconditions to adventure, it has not been difficult for me to undertake new and apparently very unpromising problems, even if they required extreme effort in observing. Hence, in consequence of these opportunities, of good health, and good friends, I am here this evening and I appreciate and thank you for this honour.

May I take advantage of this opportunity to speak a few words concerning another matter which is of very great moment. Reference of course is had to the formal opening this afternoon of the excellent new David Dunlap Observatory. This magnificent observatory, given to the University by Mrs. Dunlap, is sure to be of important service to astronomy and to add still further to the high repute of your University.

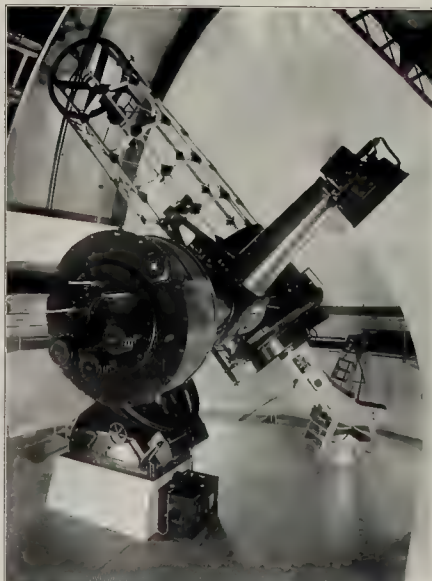
On such an occasion some persons not familiar with the developments of astronomy might ask why is another great telescope needed—is not the field covered already? But with astronomy, as with medicine and many other sciences, the more we learn the more questions arise to be answered, and often times these new ones are at least as important as were the original ones. But perhaps we ought to be more specific in answering this question. And in

PLATE III.



THE GREAT DOME
Journal of the Royal Astronomical Society of Canada, 1904

PLATE X



THE 24-INCH REFRACTING TELESCOPE
Journal of the Royal Astronomical Society of Canada, 1904

PLATE IV



THE ADMINISTRATION BUILDING, DAVIS DUNLAP OBSERVATORY

PLATE XI



THE CASE FOR ASTRONOMICAL OBSERVATION

Journal of the Royal Astronomical Society of Canada, 1904

Journal of the Royal Astronomical Society of Canada, 1904



Journal of the Royal Astronomical Society of Canada, 1904

THE LIBRARY OF THE DAVIS DUNLAP OBSERVATORY
Journal of the Royal Astronomical Society of Canada, 1904

doing so I cannot do better than to draw examples from my experience at the Lowell Observatory. Founded in 1894 by Dr. Percival Lowell for the study of Mars and the other planets, such study has been carried on there continuously since, employing every new means that could be applied to its researches. And whereas much advance has been made in our knowledge of Mars regarding his atmosphere, temperature, seasons, etc., there is still much to learn about this the best known of the planets.

The same story is to be told of the other planets. Thirty odd years ago when taking up the study of the spectra of the planets it was possible with commercial plates to photograph only from the violet to the orange. But now, thanks to repeated improvements in sensitizing dyes, it is possible to include three times the old range of wave-lengths. By thus being able to reach far out through the red into the heat spectrum, we have added enormously to our knowledge of planetary atmospheric absorptions. And recently we have been able to determine that the remarkable absorptive action of the atmospheres of the giant planets (which has long been a most puzzling question) is due to the containing surprisingly large amounts of methane gas. This substance, commonly called marsh-gas, is produced in our air by decaying matter but but that on the giant planets must have arisen not from organic matter but from inorganic processes of these planets. Here again important advance in knowledge raises many questions and it encourages us to push forward with the aid of this knowledge to further understanding of these neighbour worlds.

Then, breaking new ground, in 1906, prompted by the publication of Chamberlain and Moulton's planetesimal theory of the origin of our solar system, which assumed a primordial body of the nature of a spiral nebula, we set about to study spectrographically these remarkable nebulae. While this work did not tell us what was the primary state of the solar system we did find it could not be such a nebula but that these nebulae were very distant, extremely large bodies and were flying through space with most enormous speeds. And thus was supplied the first observational basis of the present day theory of the expansion of the universe.

And once again, the search for Dr. Percival Lowell's predicted Planet X which led to its finding at Flagstaff in 1930, within a few degrees of his predicted place, encouraged further searching of a very large belt of the sky for still more distant members of our planetary system.

So we see how one undertaking has led us to a greatly expanded universe and the other to a much enlarged solar system, both inviting much subsequent investigation. Yet these are only a few outstanding examples of progress in astronomy. Still others might be given, but these may suffice to make it evident that there is real need for new and well equipped observatories, that there is ever so much more to learn, and that there are many more fields to explore.

In conclusion I bring to the David Dunlap Observatory, to Mrs. Dunlap, its donor, to its director, Dr. Chant and his colleagues, the cordial greetings and good wishes of the Lowell Observatory. May your labours meet with results more fruitful of knowledge than your fondest hopes.

The last of the recipients of degrees to address the audience was Dr. W. E. Harper, who said:

I am deeply conscious of and grateful for the high honour just conferred upon me. I feel that any little thing I may have accomplished in the past does not merit such recognition; nevertheless, the honour will stimulate me to greater endeavours in the future.

I have often been asked, "How on earth did you come to take up the study of astronomy?" The answer can be given very definitely. I entered the University intending to qualify as a teacher of mathematics in High Schools and Collegiate Institutes. Early in my second year, though, I attended lectures on astronomy given on certain nights each month in the lecture room of the Chemical Building. I was attracted to the subject by reason of these interesting addresses given by Professors Chant, DeLury and other members of the local Astronomical Society. On entering my third year to continue my Mathematics and Physics course it seemed the logical thing to select the option which led to specialization in astronomy.

At that time I was not aware of a new observatory being under construction at Ottawa, so my choice was not in any way influenced by considerations of a "bread and butter" job. However, about the time I was ready to graduate they needed some one at Ottawa and, being recommended from here, I had the somewhat unique experience of being informed a few minutes after the graduation ceremony that I had been appointed to the Ottawa Observatory. That was how I entered astronomy.

The every day uses of astronomy do not obtrude themselves upon our notice, yet they are none the less important. The determination of accurate time, so necessary in our modern age, the application of our science to precise surveying and the delimitation of national boundaries, the calculation of the tides, the determination of the position of a ship at sea, or the aeroplane on a long flight, are all illustrations of where a knowledge of astronomy is essential.

But there are other uses, not necessarily of practical value, but worthy of consideration, nevertheless. In our quest for the beautiful in life we miss one of the richest sources of all if we have not an acquaintance with the stars. Emerson expresses the idea in this way: "If the stars were all to be blotted out and were to make their appearance only at long intervals, say once in a hundred years, how, as the critical time approached, the whole world would await with expectant awe, the oncoming of the gorgeous spectacle and how they would treasure for generations to come the glimpse they had been given of the City of God." Yes, the stars are a source of wonder and beauty.

Astronomy is a subject also that we should have some little acquaintance with if only to appreciate the many allusions to it in literature. Poetry of past ages has been rich in references to what was known as "the music of the spheres." The planets and other celestial bodies were each supposed to be affixed in a sort of crystalline shell or sphere and as these spheres, rotated about their common centre, the earth, the solid gliding of one over the other produced a harmonious strain always referred to as the music of the spheres. Numerous such allusions abound in literature and can be appreciated only by those having some knowledge of the subject.

A knowledge of astronomy, moreover, prevents one being a prey to fear when catastrophes are predicted for the earth. In earlier days, before comets were known to be regular members of our system, one can understand the alarm the unexpected appearance of a comet would produce. They were supposed to be harbingers of some dire catastrophe to come upon the earth. Pestilence and death followed in their train, so it was thought. Such concern and anxiety could not exist in our day, and yet it does.

Again, if we knew even the rudiments of astronomy we would realize the hoax that is perpetuated in the casting of horoscopes which even in our enlightened 20th century still clutter up the pages of the daily newspapers. Many people are still living in the superstition of the middle ages when it was believed that the stars and planetary configurations affected one's destiny. How worthy is a knowledge of the subject if only to rid us of this all-too-prevalent gullibility.

We need not enlarge further upon these more or less practical claims for consideration, for they are unimportant in comparison with the contributions which the pure knowledge side of astronomy has made. Professor Shapley has just referred to the difficulties the astronomer finds himself in by reason of the thin atmosphere covering that surrounds us. That is bad enough but can we picture our present state of civilization if the earth were externally shrouded in clouds, thick enough to obscure the heavenly bodies from view, though not sufficient to cut off the sun's energy. We would, of course, know something of the earth's surface and a mile or so of the atmosphere but we would be creatures of restricted outlook, nevertheless. We would go on with our work of raising potatoes and wheat for food, always looking down, never up, knowing nothing of the great universe, save this thin stratum of earth and air—such would be our condition to-day. Indeed, with the heavens in open view such might even be given as a picture of life 400 years ago before the days of the telescope, the spectroscope and the photographic plate.

The main purpose of astronomy, then, is to understand the universe. It is not that such knowledge will yield any financial returns, although there is rarely ever a discovery in pure science but contributes directly or indirectly to others which have practical applications. However the astronomer works with no such aim in view, but simply for the sake of finding things out.

I have thought that to illustrate something of the advances made in solving

the riddle of the universe during the 29 year interval since I last stood upon this platform on a similar occasion, I might refer to the item of radial velocities mentioned in the all-too-generous citation just read. In 1906 the number of stars whose radial velocities were known was about 400, chiefly determined at the Lick and Yerkes Observatories. The work taken up at Ottawa and for the past seventeen years at Victoria has resulted in Canada having a very creditable share in this phase of astronomical work. Of the 7000 stellar velocities now known, approximately 2000 have been determined at Victoria, although not all have as yet been published. In this work I have had some small share.

Much as has been learned from such researches we can nevertheless say with a great scientist of a former age that what we have explored is but the smallest portion of the great ocean of truth.

President Cody then spoke of the various distinguished guests on the platform and welcomed them individually. He asked Mr. Mackenzie King to address convocation.

Mr. King stated that it was his purpose to attend the opening ceremonies at the observatory without being seen in Toronto or missed in Ottawa. He had been seen in Toronto; how well he had avoided being missed in Ottawa he could learn when he returned. He wished to pay tribute to all those who had contributed to the realization of the construction and completion of the observatory. It was a great day for the University and a great day for the country, but also a day of remembrance for the whole scientific world.

He most warmly thanked Mrs. Dunlap and praised her selection of such a memorial to her husband. The David Dunlap Observatory was a combined monument, resembling a double star, of which the components are so close as to seem to be one, to the generosity of the donor, and to the memory of her husband. He spoke kindly of the long-enduring and unflagging enthusiasm and determination of Dr. Chant, and rejoiced in this happy culmination of his plans.

From his own personal point of view, he spoke of his happiness at being able frequently to return to his Alma Mater. He did not come for relaxation, nor recreation, but because, in the University, he found real inspiration for his own life. And in the realization of such a magnificent, unselfish, and noble project he found an especial inspiration.

Convocation was then dismissed.

SOME ADDITIONAL GREETINGS

From the Dominion Astrophysical Observatory, Victoria, B.C.

(Telegram to C. A. Chant)

Birthday greetings and hearty congratulations to all your staff.

(Signed) STAFF, Dom. Astro. Obsy.

From the Victoria Centre, Royal Astronomical Society of Canada—to Professor Chant

The Victoria Centre of the Royal Astronomical Society of Canada has learned with much pleasure of the completion of the David Dunlap Observatory with its reflecting mirror of 74 inches diameter.

The Executive desires to extend, on behalf of the Centre, its heartiest congratulations to you upon the fulfilment of your cherished ambition to secure such an observatory for Canada and in particular for the University of Toronto. Through your association with the *Journal* as Editor the members have come to feel that they know you personally and wish to send you this message of congratulation and best wishes for the success of the observatory.

We should like to place on record also our appreciation of the great contribution Mrs. Dunlap has made to the higher life of the country in the gift of such a splendid observatory. The aid she has thus rendered the cause of pure science will, we trust, be an example to other Canadians to do likewise.

To the University which receives this splendid gift to administer, we would offer our congratulations also, realizing that through the use of this new equipment researches will be possible which will add additional lustre to the University.

Signed on behalf of the Victoria Centre, R.A.S.C.

H. BOYD BRYDON, President.

GORDON SHAW, Secretary.

From the Vancouver Centre, R.A.S.C.—To Mr. Harper

I wish to thank you for your kind letter, and request that while in the east you convey from the Royal Astronomical Society, Vancouver Centre, hearty greetings and congratulations to Prof. C. A. Chant, to whom we have for many years been deeply indebted for his valuable services to Astronomy and particularly for his being instrumental in securing the new David Dunlap Observatory for our country.

(Signed) C. A. McDONALD, Secretary.

From the Edmonton Centre, R.A.S.C.—To Mr. Harper

I have been instructed by Dr. Gowan, president, to say that we would be glad to have you convey to Dr. Chant greetings and congratulations from the Edmonton Centre of the R.A.S.C.

(Signed) G. A. CLARKE, (Secretary)

From Prof. J. W. Campbell, University of Alberta, formerly President of the Edmonton Centre.—To C. A. Chant

Our Centre has been glad to join with other Centres of the Royal Astronomical Society in extending to you our best wishes and congratulations on the opening of the new observatory, but I should like to add a personal word of greeting. I want to express my appreciation of your cordial and kindly interest in all the matters that I have had to bring to your attention, and for the generous assistance you have always given when called upon;—and not only for the academic interest and assistance you have given, but also for the cordial hospitality of your own home which it has been my privilege to enjoy.

I regret that Mrs. Campbell and I are not in the east this week to participate in felicitations to you and Mrs. Chant on the occasion of this very fitting recognition and reward for your achievements in the interest of astronomy, but please accept this intimation of our very best wishes to you both.

(Signed) J. W. CAMPBELL

From the Winnipeg Centre, R.A.S.C.—To C. A. Chant

It is with great pleasure that we ask Mr. W. E. Harper, who is attending the 1935 meeting of the Royal Society of Canada, to convey to you the greetings of the Winnipeg branch of the R.A.S.C.

We have followed the planning, building, and equipping of the new Observatory at Toronto with great interest and pride, and realizing how much of this fulfilment is due to the untiring efforts of you and your associates we, on the occasion of the official opening of the David Dunlap Observatory, send you our felicitations, greetings, and hopes that you may be spared long to direct the Observatory, and that, notwithstanding the noteworthy contributions made by Victoria and Ottawa to the science of astronomy, Toronto may outshine them both.

Good luck and good seeing, always!

(Signed) WINNIPEG CENTRE,

per L. T. S. NORRIS ELYE (President)

From Prof. A. Vibert Douglas, McGill University, Secretary of the Montreal Centre, R.A.S.C.—To C. A. Chant

It is a matter of great regret to me that I cannot be present on the momentous occasion of the official opening of your observatory. I would have liked to add my humble congratulations to those of the many distinguished well-wishers who will be with you on that occasion.

May health and strength be granted you to enjoy the fruit of your labours and to see the David Dunlap Observatory achieve a reputation for valuable

research. May great happiness be yours on your birthday and through the succeeding years.

(Signed) A. VIBERT DOUGLAS.

Mr. GEORGE R. LIGHTHALL, president of the Montreal Centre, was present at the opening of the Observatory.

Dr. RALPH E. DELURY and Mr. CLARENCE B. HUTCHINGS of the Ottawa Centre were also present.

Professor H. R. KINGSTON, former president of the London Centre, with Mrs. Kingston, attended the opening.

Of the Toronto Centre a large number of members were present, including Professor L. Gilchrist, president of the R.A.S.C. and Mr. A. R. Hassard, K.C., chairman of the Centre.

THE DAVID DUNLAP OBSERVATORY*

By R. K. YOUNG

With Plates X-XVII

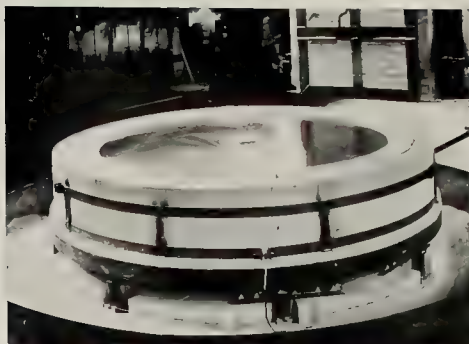
THE David Dunlap Observatory, the gift of Mrs. Jessie Donaldson Dunlap to the University of Toronto as a memorial to her husband, will be formally opened on May 31, 1935. It is very fitting that the Observatory should be opened on this date, which is the seventieth birthday of the Director, for Dr. Chant has spent the greater part of his life in fostering astronomy within the University and in encouraging its study throughout Canada. The purpose of this article is to tell briefly the story of this magnificent gift and to say something of the research which will be carried on by the new observatory.

ASTRONOMY IN THE UNIVERSITY

The progress of astronomy as a department of the University is due to the continued efforts of Dr. C. A. Chant to emphasize its importance as a cultural subject in education and as a training for the advanced student. It was a part of his plan, even from a very early date, that the University should have an observatory and contribute to the knowledge of the subject, but it was hardly expected that the money for its erection would be obtained from the provincial grant to the University. In an institution striving to meet the needs of the province and expanding rapidly, chief emphasis in the field of science was likely to be placed on subjects more immediately utilitarian. Not until these had been taken care of would the claims of a pure science like astronomy be considered. The interest in the subject in recent years has been much increased by the spectacular discoveries which have greatly extended our knowledge of the universe and which have appealed to the imagination.

*Reprinted with slight revision from the *University of Toronto Quarterly*, Vol. IV, No. 3, April, 1935.

PLATE XIII

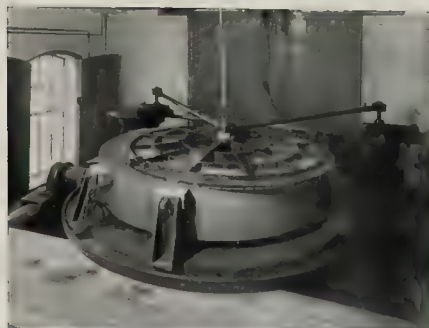


THE DISK OF GLASS FOR THE 74-INCH MIRROR IN ITS MOLD
AT CORNING, N.Y. JUNE 21, 1933



THE DISC IN THE SHOPS AT NEWCASTLE-ON-TYNE, MAY, 1934
The disc has been ground approximately to shape; the hole at the centre is being bored out.

PLATE XIV



THE DISC FOR THE 74-INCH MIRROR
After the grinding of the surface is complete. The disc is now ready for the rough grinding of the back. When cut the disc was 74 inches in diameter and 1 1/2 inches thick. (Photograph at Newcastle, December, 1934.)
Courtesy of the Corning Glass Works, Corning, N.Y.

PLATE XVI



THE GREAT MIRROR
The mirror has just been lifted from its case and is about to be placed in the mold.
May 2, 1934
In front: T. J. Healy, left; J. S. Hagg.
Behind: A. Young, left; E. H. Shaw, right.
Courtesy of the Corning Glass Works, Corning, N.Y.

PLATE XVII



THE GREAT MIRROR
The mirror has just been lifted from its case and is about to be placed in the mold.
May 2, 1934
In front: T. J. Healy, left; J. S. Hagg.
Behind: A. Young, left; E. H. Shaw, right.
Courtesy of the Corning Glass Works, Corning, N.Y.

Astronomy owes much also to the great body of amateurs whose interest in the subject has strengthened the desire that an efficient telescope might be situated within the province. The late David Alexander Dunlap was one of these. He was a member of the Royal Astronomical Society of Canada and attended the meetings of the Toronto centre. Dr. Chant in all his lectures before the society and throughout the country has emphasized the observational side of astronomy and the need of an observatory. If those who give instruction in astronomy are to possess that inspirational touch which comes only from an intimate knowledge of the subject, they must either be actively engaged in astronomical research, and thus be extending the frontiers of the science, or must be thoroughly informed regarding the work carried on in the observatories of the world. Lacking the equipment for the former, Dr. Chant followed the latter course; but he constantly emphasized the need of a first class observatory. It was his hope that aid in this project would be received from Mr. Dunlap, but the latter's death in 1924 prevented this. When, some time later, Professor Chant suggested to M. Dunlap that she should provide the observatory as a memorial to her husband, the suggestion met with a sympathetic response. Indeed, Mrs. Dunlap shared her husband's interest in astronomy.

GENERAL PLANS AND LOCATION

In 1927 Mrs. Dunlap expressed her willingness to provide the observatory, but it was not until June of 1928 that we were in a position to call for the construction of the various instruments which would be installed. The original plan, which has been closely adhered to, contemplated two buildings: one, a steel structure to house a large telescope; the other, an administration building for office work and the reduction of the observations. There was no haste about the construction of the latter building since it offered no particular difficulties, but the main telescope was ordered as soon as possible because the time required for its construction was somewhat uncertain, this being especially true of the large mirror which forms the essential optical part of the telescope.

The location of the observatory was an important point to decide. Dr. Chant and the writer spent many afternoons inspecting the maps of the neighbourhood of Toronto and visiting the possible sites. It was not thought advisable to go more than twenty or twenty-five miles away from the city, and locations north or north-west were much preferable to those east of the city. Most of our clear weather comes with west or north-west winds, and at these times the smoke from the city is blown south-east or east. A considerable amount of experimenting was carried on to determine the transparency of the air thirty, fifteen, and four miles from the city, as well as the amount of sky-illumination. In this regard the stations thirty and fifteen miles away proved far superior to that near the city, especially in the matter of the sky-illumination. The gain between thirty and fifteen miles did not seem to warrant placing the observatory at the more distant station. The site finally chosen is about twelve miles north of the city limits and about five-eighths of a mile to the east of Yonge Street on the highest land in that immediate locality, the summit being eight hundred feet above sea level. The land falls away in all directions, giving a remarkable view of the surrounding country. At present the land around the observatory is quite open, with a few trees and shrubs scattered here and there. From an astronomical point of view it would be better if the land were more heavily wooded. On an open plain the ground becomes hot during the day; and when, after sunset, the heat is given off again, the warm air flowing upwards creates an unsteady atmosphere which interferes with the use of the telescope. Trees and vegetation absorb a great deal of the sun's heat so that a steady state of the air is reached much more quickly. Accordingly, it is intended to make the land into an arboretum to be known as the David Dunlap Park.

AWARDING THE CONTRACT

Comparatively few firms possess machinery large enough to handle the massive castings of a great telescope, and there are still fewer with experience in telescope-building. The tentative specifications were sent to four firms: Carl Zeiss in Germany; Sir Howard Grubb, Parsons and Company in England; Warner and Swasey Company of Cleveland; and J. W. Fecker of Pittsburgh. The Warner and Swasey Company did not submit a tender, and the design of the Carl Zeiss firm was considered less satisfactory than the one selected. There was not much difference either in the designs or prices of the other two firms, but after due consideration it was decided to accept the tender of the English company. This was a very fortunate choice because the decrease of the pound sterling made the cost very much less than it would have been if the contract had been let in the United States. It was very satisfactory, also, that the complete contract for the mounting, for the steel building to house the telescope, and also for the optical parts, could be let to one firm. This made it much easier to ensure that the finished equipment would assemble without difficulty when it arrived. Each of these three items merits some description.

THE CIRCULAR STEEL BUILDING

The building to house the large telescope was ordered in November, 1931, and it was received in Toronto on July 31, 1933, (in pieces.) As has been already stated, the entire building is of steel construction. The circular drum and the hemispherical dome resting on it have double walls, and an open space at the base of the building admits air, which circulates to the top and goes out through openings with baffle plates at the upper part of the dome. The inside and outside of the dome are covered with "agasote," a hard paper product, and the outside is further protected by a sheeting of copper. By this means the interior of the building is kept cool during the day, and in the evening when thrown open for observations the whole building soon assumes the temperature of the outside air. Thus the "definition" of the telescope is not interfered with by heated air-currents, as would undoubtedly be the case if any quantity of heat were stored in the walls. The dome, which weighs about eighty tons, rests on twenty-four rollers, twenty-seven inches in diameter, running on a circular track. The entire dome may be rotated so that the opening, which is fifteen feet wide and extends from the bottom to seven feet beyond the zenith, may be made to face any part of the sky. Two parallel shutters run on rails at the top and bottom of the dome and cover or uncover this opening. The building cannot be heated because of the air

currents which would be set up. In the cold weather the observer must dress for the occasion.

THE MOUNTING OF THE TELESCOPE

The telescope which the building houses is comprised essentially of the big mirror and a suitable means to hold it and direct it toward any part of the sky. The mirror, which consists of a block of glass seventy-six inches in diameter and about twelve inches thick, weighs about five thousand pounds, and consequently the mounting must be correspondingly heavy. This mounting is so constructed that the telescope can be pointed to any desired part of the sky and be continuously moved by clockwork in order to follow any object from the east to the west as it passes across the sky owing to the rotation of the earth. Although the moving parts of the mounting weigh upwards of thirty tons, this motion must be perfectly steady. It demands great perfection in the construction. The completed mounting, with the exception of the mirror, was received in Toronto on October 15, 1933.

THE GREAT MIRROR

When the telescope was ordered, in 1930, we knew that the portion which would probably take the longest to complete was the big mirror. At that time the Grubb-Parsons Company controlled the Parsons Optical Glass Works at Derby; and Sir Charles Parsons, the head of the C. A. Parsons and Company, of which these other companies were subsidiaries, was confident that they could manufacture a suitable disk of glass for the telescope mirror. (Incidentally, Sir Charles Parsons was the youngest son of the Earl of Rosse, who completed a six-foot reflector in 1845.) But Sir Charles was in his seventy-sixth year when the order for our telescope was placed, and unfortunately he did not live to see the disk made. Had he lived I have no doubt that his active interest and ingenuity would have solved the difficulties and pushed the task to completion. But in 1932, after his death, the disk had not yet been cast, and it

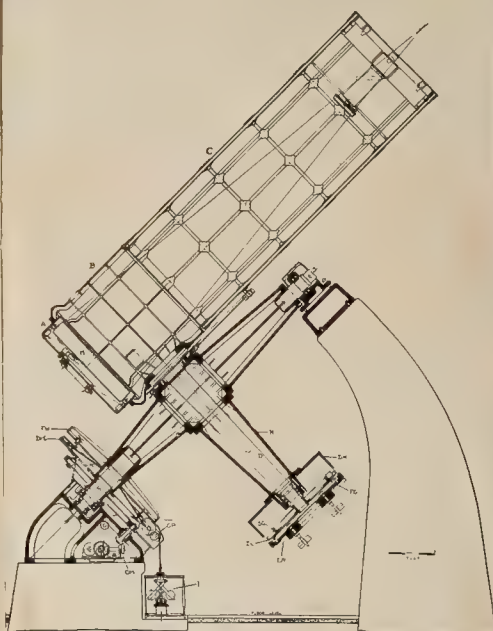


FIG. 1—Vertical north-south section of the 74-inch telescope

seemed that the project for our observatory might be unduly delayed.

However, in 1932 unexpected help arrived in connection with the manufacture of telescope mirrors, which was not available in 1930. In the latter year the only firms which would undertake the manufacture of large disks were Carl Zeiss in Germany and the Glass Works at Derby in England. On the American continent the Corning Glass Works, of Corning, N.Y., had made some small "pyrex" disks of glass which were superior in every way to any that had been previously made, but this firm was extensively engaged in the commercial manufacture of pyrex articles and was not prepared to undertake the expensive experimenting necessary to manufacture so large a disk as we required. But between 1930 and 1932 conditions changed. Plans had been put forward for the manufacture of a disk for a two-hundred-inch telescope and time and money spent in finding out the most suitable material. In the end it was decided that "pyrex" glass offered the best hope of success for this disk. The Corning Glass Works was prevailed upon to install the necessary furnaces and annealing ovens for the task. We were informed late in 1932 that they were prepared within six months to cast our disk. From the first we should have chosen this material for the large mirror had it been available at the time. The Grubb-Parsons Company gave the contract for the manufacture of the raw disk of glass, which was to be shipped to England to be ground and polished into the final mirror. It was with great hope and satisfaction that our party from Toronto, including Mrs. Dunlap, Dr. Chant, and others, was present in Corning on June 21, 1933, and saw the disk poured from a special type of "pyrex" developed for telescope mirrors. Possibly with still greater satisfaction Dr. Chant and the writer inspected it on September 29 of the same year after it had come from the annealing oven, and saw a disk whose appearance met all our hopes.

No time was lost in shipping it to England. It was still to be ground and polished, and Mr. C. Young, the manager of the Grubb-Parsons Company, estimated that this would take eighteen months. Only those who have attempted to make a telescope mirror can really appreciate the difficulties that are encountered in such a task. The material is fragile and the disk heavy, and the surface has to be



FIG. 2—Unloading the Great Mirror from truck into Dome, May 2, 1935

The David Dunlap Observatory

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shaped true to an extraordinary degree of accuracy. The figure of the mirror, we are informed, now warrants the makers in ceasing work on it. Measures and photographs which have been received lead us to believe that the surface is in no place at fault more than one millionth of an inch.

THE ADMINISTRATION BUILDING

The large telescope will be an effective instrument of research. In addition to the building containing it, it is necessary to have a suitable place where the plates taken with the telescope may be studied, where the staff can have office space, and the activities of the observatory be directed. The construction of the Administration Building was conducted by the University's Superintendent of Buildings and Grounds, who also had charge of the erection of the steel building, of securing the water supply, of the electric installations, and of other essential details. Great credit is due to his management and to the architects, Mathers and Haldenby, for the fine appearance of the building, both within and without. It is a stone structure, built on classical lines, and is situated about two hundred and fifty feet from the great dome.

On the roof of the building three domes were erected for the accommodation of smaller instruments. The central and north domes are at present vacant. In the south dome is a nineteen-inch telescope which was built at the University. This telescope, we hope, will prove a valuable accessory to the larger instrument and may be used on many programmes of research when the large telescope is not available.

In the Administration Building is also the library of the Observatory. Original plans for a special endowment for the library could not be carried out and this essential adjunct to research is not yet satisfactorily provided for. The Observatory and the University of Toronto are indebted to the Royal Astronomical Society of Canada for the loan of the more technical part of its library to the Observatory for a term of three years. The library is not nearly as complete as it should be and will have to be supplemented either by generous grants from the University or from funds obtained from other sources.

THE WORK OF THE OBSERVATORY

Almost the entire research programme of the Observatory will be carried on photographically. It is estimated that during a year we shall have about one hundred and twenty-five clear nights with twelve hundred observing hours. The large telescope will be in use from sunset to sunrise, either in directly photographing regions of the sky or in the employment of auxiliary instruments such as the spectroscopic with which the spectra of the stars are secured. Considerable time will have to be utilized at first in testing the accuracy of the instrument and in determining methods of work in order to make the labour of the limited staff as efficient as possible. Only after operating for a number of years can an observatory have accumulated that wealth of astronomical photographs which forms the basis of research on many astronomical problems. After this is secured there is forthcoming a steady stream of valuable work. It will probably take a number of years for us to reach this goal, and the first duty of the staff will be to accumulate the observational material. While a new observatory is thus handicapped for a time, it has compensating advantages: it can utilize all previous experience in order to obtain the material in a superior form, so that it can be used for more lines of investigation.

On account of the large size of the new telescope, the second largest in the world at the present time, we shall be able to secure much material which has never before been obtained. Astronomers throughout the world, through their organizations and publications, are well informed regarding the work which is being carried on, and the various observatories co-operate in the prosecution of researches for which their equipment is best suited. The David Dunlap Observatory will participate in the field of stellar velocities, spectral photometry, and in other co-operative labours. The major part of the observational programme will be devoted to such researches, and by contributing to the advances in astronomy it is hoped to justify the generosity of the donor in presenting the University of Toronto with this magnificent instrument.

During the last twenty years the nature of "time" has been a subject of much lively debate. The publication of Einstein's General Theory of Relativity caused great intellectual quickening, not only in academic circles but among factory workers, in mining camps and in outposts of empire. Is time absolute? Has it a separate independent existence or does it depend on something else? For two centuries Newton's definition was universally accepted—at least by students of physical science:

Absolute, true, and mathematical time, of itself, and from its own nature, flows equally without relation to anything external, and by another name is called duration.

Space also is treated by him in similar simple terms which seem clear to all! And yet Einstein and his supporters, who sprang up everywhere, boldly stated that time and space were simply relative notions. According to Minkowski,

From henceforth space in itself and time in itself sink to mere shadows, and only a kind of union of the two preserves an independent existence.

However when a person is told that he has reached the age of three-score-and-ten and that he must therefore retire, whether he likes it or not, he feels that time is pretty absolute after all.

About two decades ago I had the opportunity of examining the records made at the Meteorological Office at Toronto for many years past. For Wednesday, May 31, 1865, the day of my birth, the following entry was found:

Temperature at 2 p.m., 76°; mean for the day, 68°·4. Fair day. Hazy round the horizon. Frogs noisy at night.

The site at which these observations were taken is now on the grounds of the University of Toronto. For May 31, 1935, seventy years later, at a station half a mile farther north, the atmospheric conditions are embodied in the following summary which is ab-

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stracted from the elaborate record as now taken:

Temperature at 2 p.m., 69°·9; mean for the day, 64°·01. At 2 p.m. the barometer read 29·759 in. which, reduced to freezing point and sea level, was 30·062 in.; relative humidity, 41; visibility, 8. From 7·45 a.m. to 12 noon, partly cloudy. Light rain from 7·35 a.m. to 9·30 a.m. At 2 p.m., partly cloudy; at 4 p.m. generally clear and sunny.

It need hardly be remarked that to hear the frogs now one would have to travel several miles from this station.

There is another semi-meteorological or semi-astronomical matter which might be referred to. We speak of a blooming maiden of sixteen summers or a hoary veteran of eighty winters, meaning that these individuals are sixteen or eighty years old, respectively; but I have seen only sixty-nine summers although I can claim seventy-one winters. The anomaly is to be credited to the eclipse of 1922, of relativity fame. Leaving our country in June of that year, a summer in Canada was exchanged for a charming winter in Australia, with the numerical result indicated.

It is in no spirit of protest that I withdraw from active service on the staff of the University of Toronto and put on the quiet gown of the emeritus professor. The regulation requiring retirement at a definite age is a good one. There are younger men, keen of mind and vigorous of body, who will carry forward the work. I shall miss the peculiar pleasure experienced by the instructor in his contact with his students, especially when they are growing into manhood and womanhood. But such will not be entirely absent. Life for me began many years ago, assuredly before forty, and it will continue after seventy. In my pleasant quarters at the Observatory I shall have more leisure to pursue investigations in which I am interested.

Further, the projects in the University and the Royal Astronomical Society of Canada, to which I have given special thought and effort, are, I think, now firmly established. They will assist in the moulding of human thought in the future just as astronomical discovery and research have in the past centuries.

During the many years in which I have conducted this JOURNAL I have not printed any verses of which I could claim authorship. Indeed many moons have displayed their phases since I ventured

upon the sea of poetry. I hope I may be pardoned if I now introduce some stanzas which were written over forty years ago; and although they do not require any elucidation they may carry a little greater interest if I explain how they came to be produced.

In 1887 after completing my first year examination at the University I went with a class-mate—long since passed away—on my first camping trip. We launched our canoe in the Stony Lake region where there were then very few summer cottages. I find the following brief record:

On Wednesday, July 13, a very hot day, we went to Jack's River with the intention of climbing the (so-called) Blue Mountain. After walking several hours and losing our way we at last reached the summit. From it was a glorious view. Over a dozen small blue lakes could be seen.

In the autumn of 1891, I was asked to send a contribution to *What-Not*, a magazine published by the Literary Society of the St. Catharines Collegiate Institute. It appeared in the Christmas number. Perhaps it is more appropriate now.

"OLIM MEMINISSE JUVABIT"

I

We camped upon the lakelet's shore,
And paddled on the river's breast;
We slept upon the grassy floor
When earth was in her sable dressed.

We climbed the mountain's rocky side
Beneath a scorching August sun,
And many times our strength was tried
Before the lofty crest was won.

And though at night, so sore of foot,
With toil and heat quite overworn,
Refreshed with sleep, we urged our route
When Phœbus rose the morrow morn.

But now a kindly memory knows
No thought of pain or blistering ray;—
The thorns of pleasure's fragrant rose
All vanish in the after-day.

On Retiring

II

For years, upon the broadening stream
Of learning has our little bark
Been moving on,—too slow, we deem,
But ever on towards the mark.

We search the much-indented shore
And chart the reefs and shoals we see,
Though oft, beneath the heavy oar,
The body sinks in agony.

Upon the bank, with nightly toil,
We seek to climb the rugged steep,
But ere we trudge a single mile
We fain would lie to rest in sleep.

Yet time a rosy glamour throws
O'er all the hardships by the way;—
The thorns of learning's fragrant rose
All vanish in the after-day.

C. A. Chant

According to tentative plans the observatory will be officially dedicated May 31, Dr. Chant's seventieth birthday. Already the huge dome housing the telescope has been erected and the grey sandstone administration building completed. All that is awaited is the all-important mirror, or "eye" of the telescope, now undergoing its final grinding and polishing process at the Parsons Optical Works, Newton-on-Tyze, England. This firm is a subsidiary of the Sir Howard Grubb Company which undertook the contract to build the observatory.

Manufacture of the mirror was the greatest accomplishment of the whole undertaking, in Dr. Chant's opinion. The experiment in casting the disc has been a long and tedious process, and for a time the experts assigned to this technical work were compelled temporarily to acknowledge defeat. The commission ultimately was transferred to a firm in Corning, N.Y., where success attended the pouring of a mammoth 306-inch mirror for a California observatory.

The Dunlap mirror was poured into its mould on June 21, 1935, and on October 1, the same year, was declared a complete success. Then it was shipped to England to undergo further technical treatment. Even if the mirror is delivered in Toronto on schedule, nearly two years will have elapsed since work on the telescope "eye" was started.

"We shall certainly be happy when the mirror is installed, but we have never been impatient," Dr. Chant said. "I would suggest those afflicted with impatience might try astronomy as a cure." In that field results were far from spontaneous, he said, with a premium rather paid to concentration, determination and patience.

Journal of the Royal Astronomical Society of Canada, 1935.



VIEW OF THE DAVID DUNLAP OBSERVATORY FROM THE AIR.

PAGE 11.

Designed and produced by ROUS & MANN Limited Toronto.

